

## ***X-ray Absorption Fine Structure Spectroscopy***

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Interactive lecture (Quicktime 7) at <http://pangea.stanford.edu/~farges/SB2005>

# Plan :

## I. Introduction

- history
- general principles : absorption coefficients, electronic mechanisms
- synchrotron sources : principles, available facilities
- data collection outlook: transmission, fluorescence, dispersive...

## II. Analysis and data reduction in real

- pre-edge region : electronic transitions
- edge region : Fermi level and redox
- interference region : XANES/EXAFS
- more complicated: multiple-scattering, disorder, multi-electronic excitations
- data reduction strategies, softwares comparison

## III. Some examples : glasses, mineral surfaces

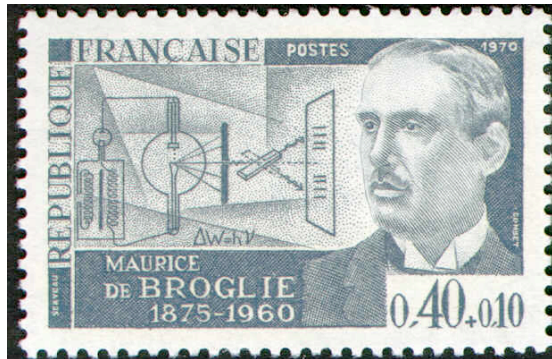
## IV. Prospectives...

# I. Introduction



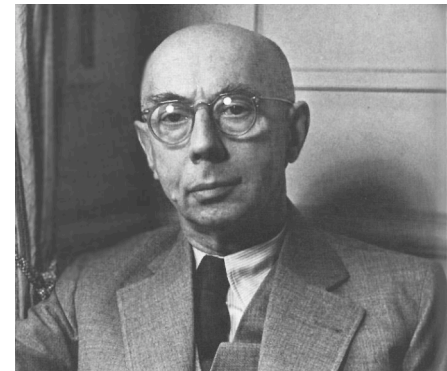
# One upon a time (*brief summary*)

1895 : Röntgen discovered  
x-rays



1913 : Maurice de Broglie measured  
the first absorption edge

1920 : Fricke and Hertz  
(*shown*) reported the first fine  
structures



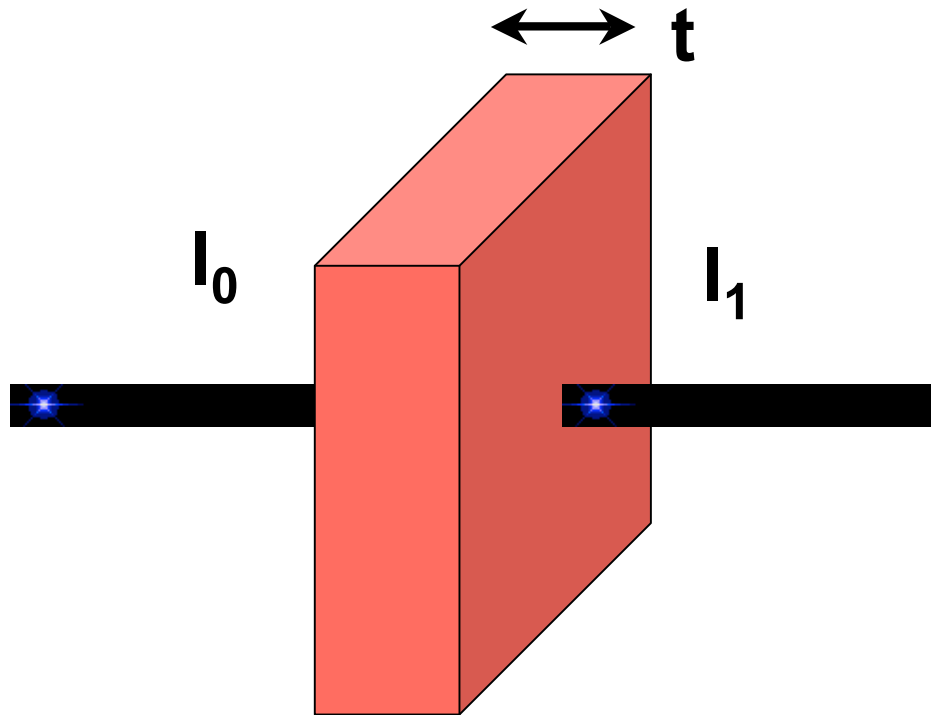
1931 : Kronig wrote the  
first theory of EXAFS

1971 : Stern, Sayers and Lytle  
introduced the most of the current theory,  
incl. Fourier transform of EXAFS

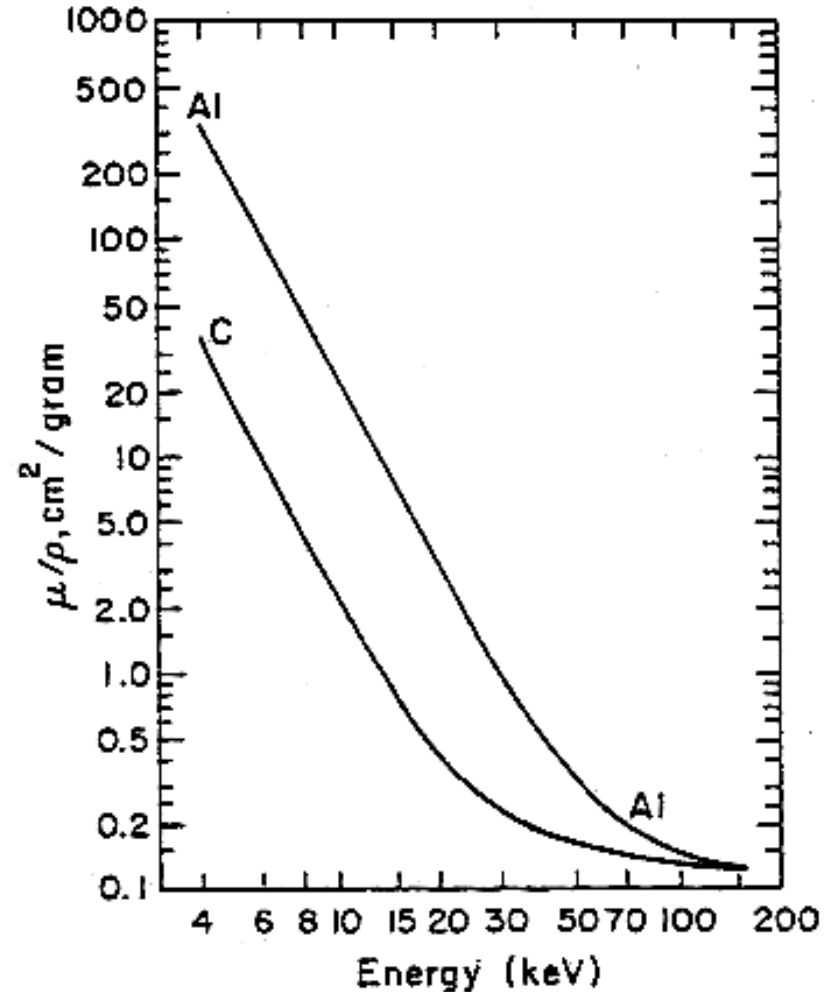


More ? See <http://www.exafsc.com/techpapers/>

The (x-ray) photon intensity ( $I_1$ ) that passes through a sample of thickness  $t$  is function of the absorption coefficient,  $\mu$  :

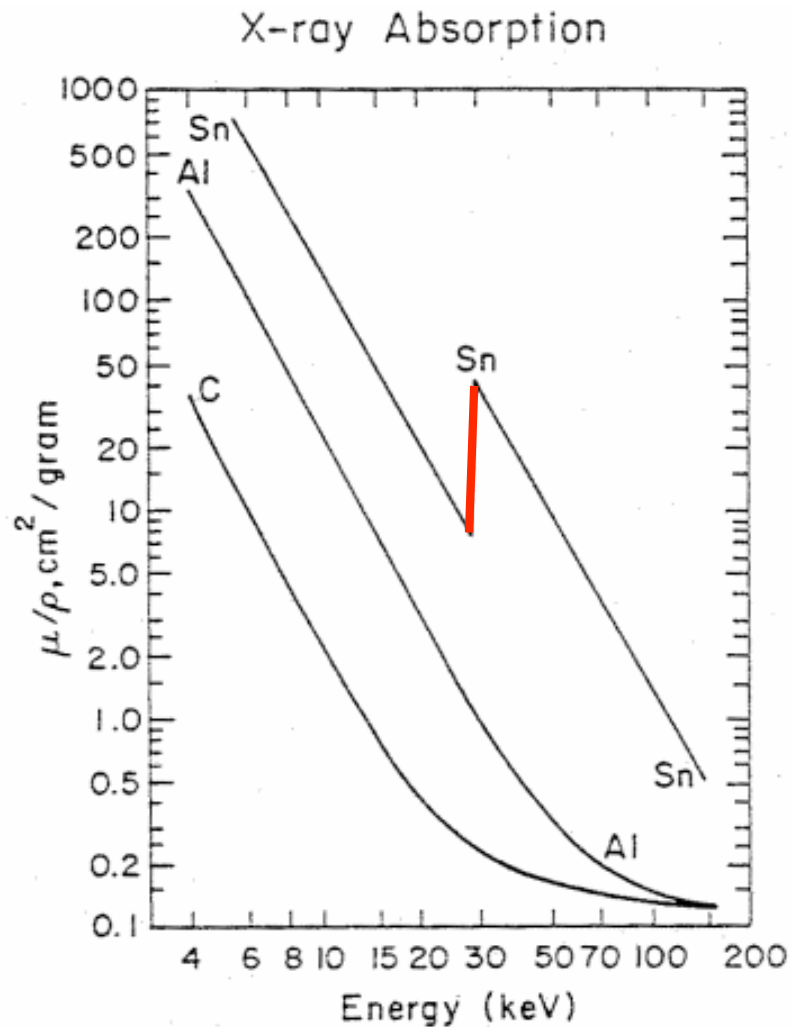


$$I_1 = I_0 e^{-\mu t}$$



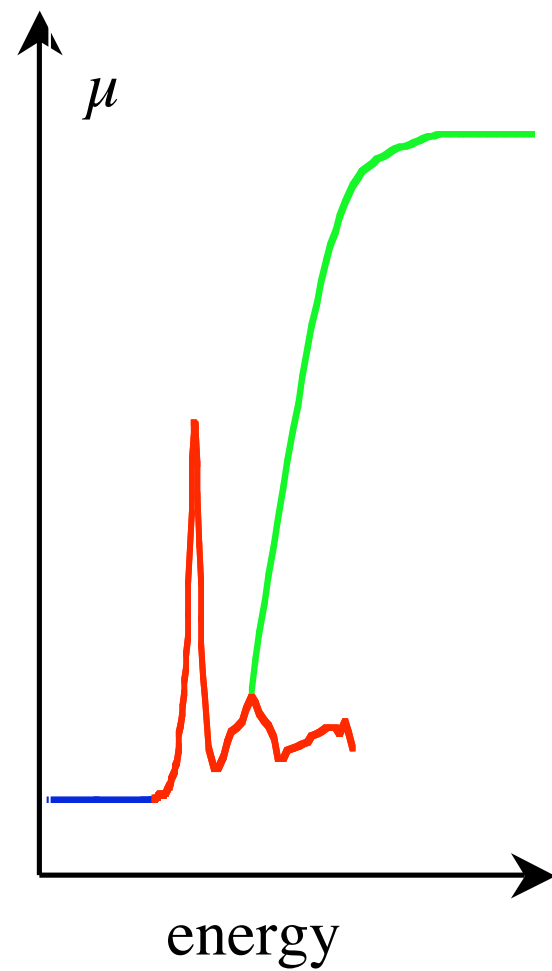
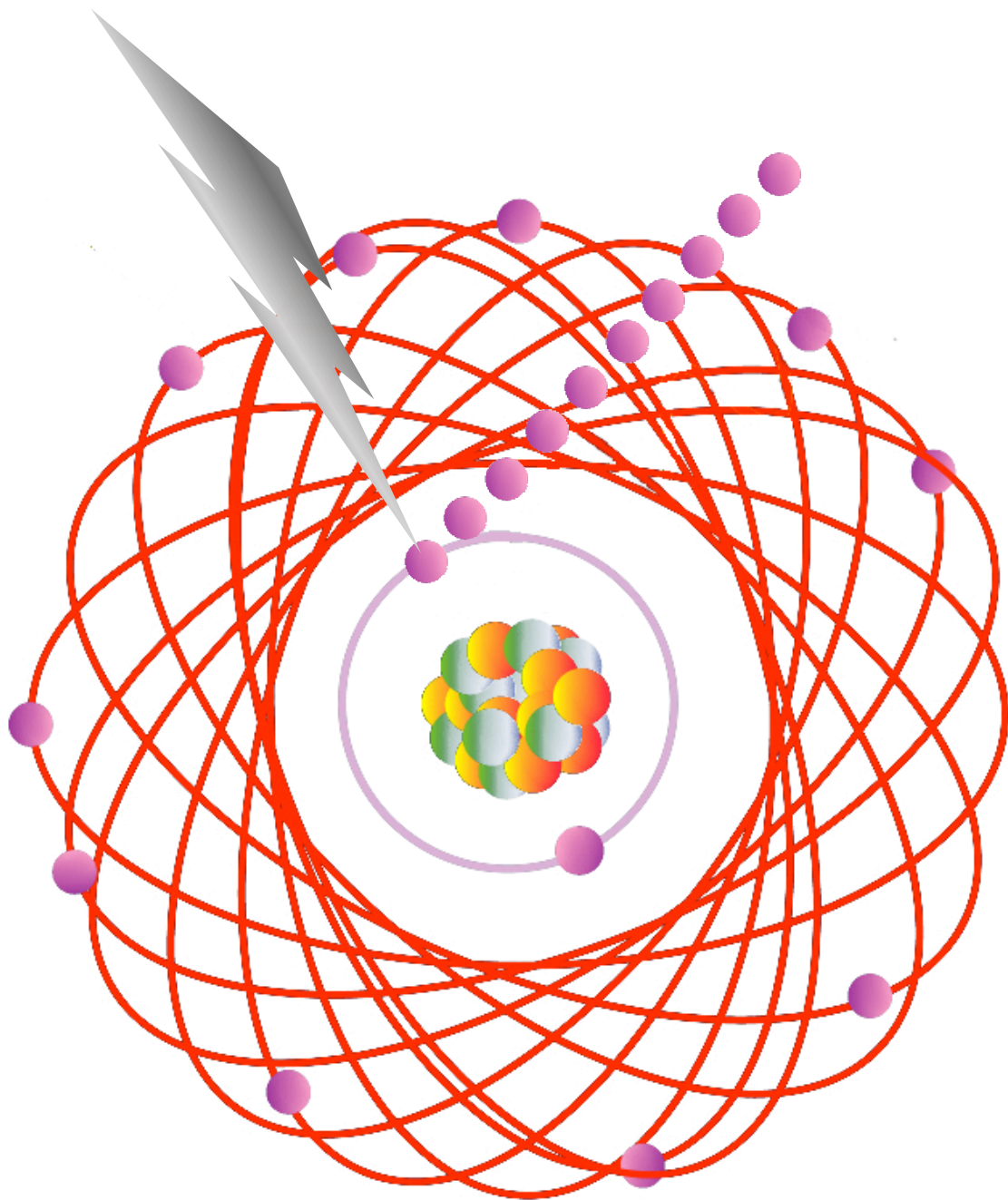
$\mu$  is function of  $Z$ , density ( $\rho$ ) and photon energy  $\sim \rho Z^4/AE^3$   
( $A$  is the atomic mass)

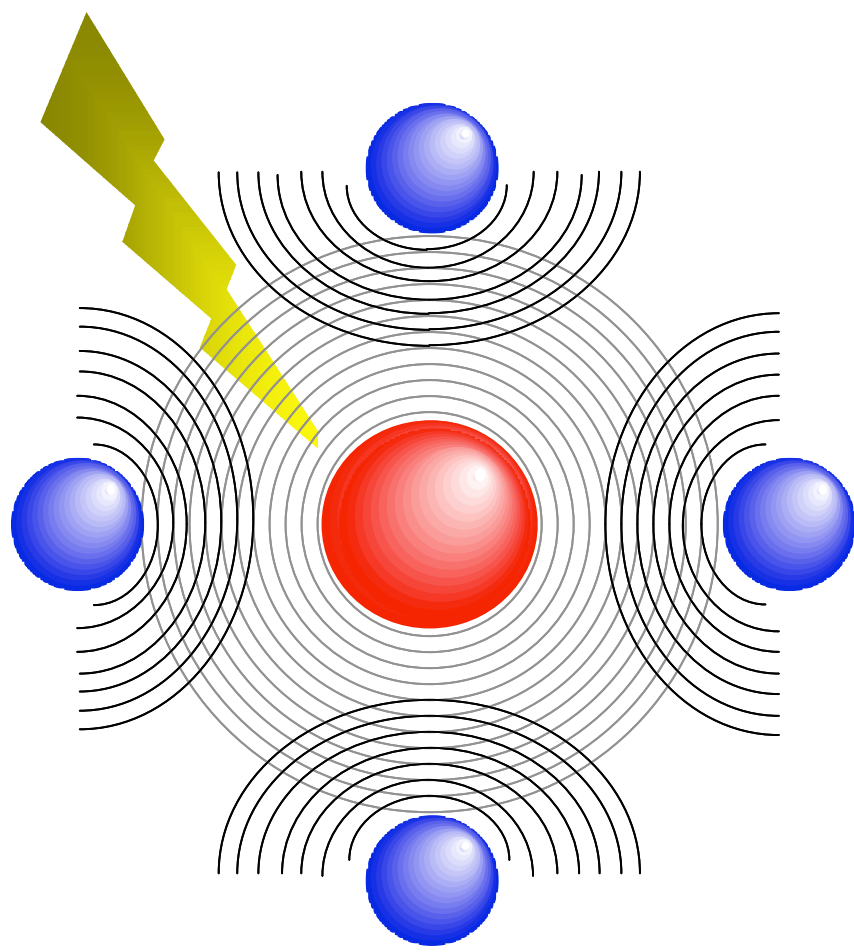
**Discontinuities in  $\mu = f(E)$  are observed,  
called absorption edges  
and related to the excitation of core electrons**



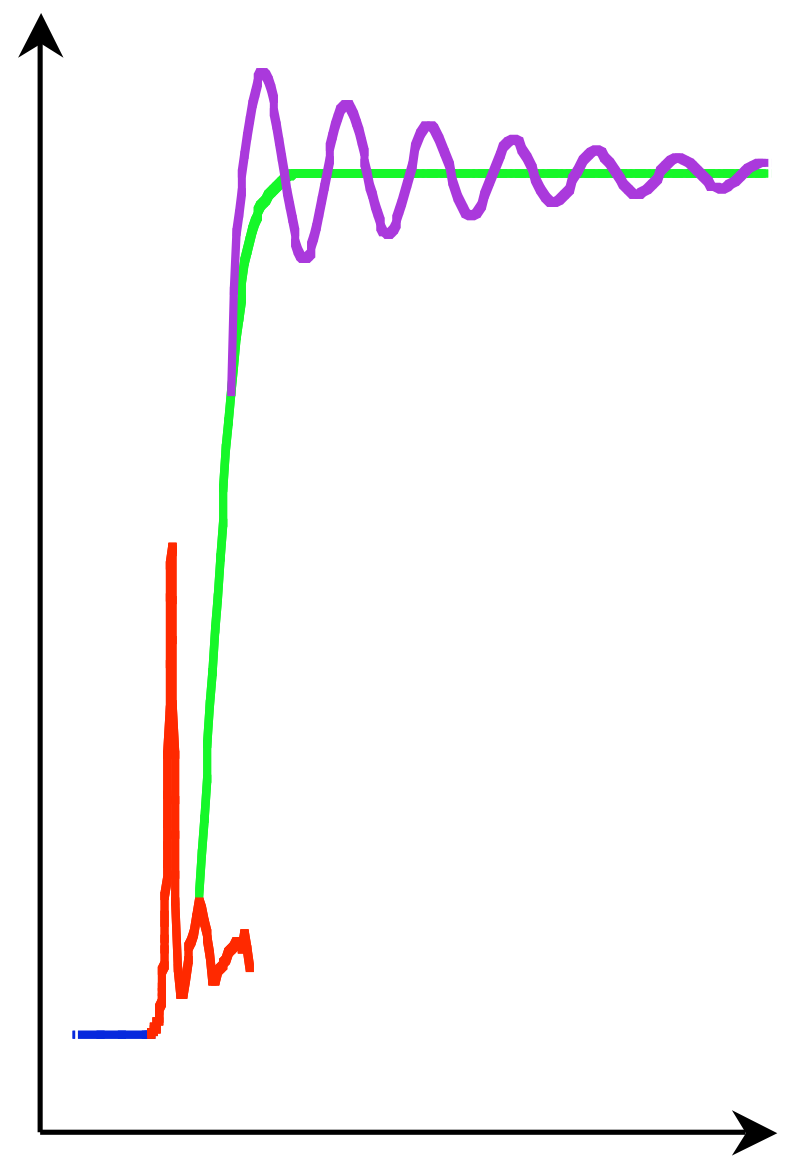
element	1s (K-edge)	2p (L edges)
H	13.6	...
Si	1840	149 - 99
Zn	9657	1194 - 1020
Sn	29200	4465 - 3929
U	115606	21755 - 17166

 available on synchrotron





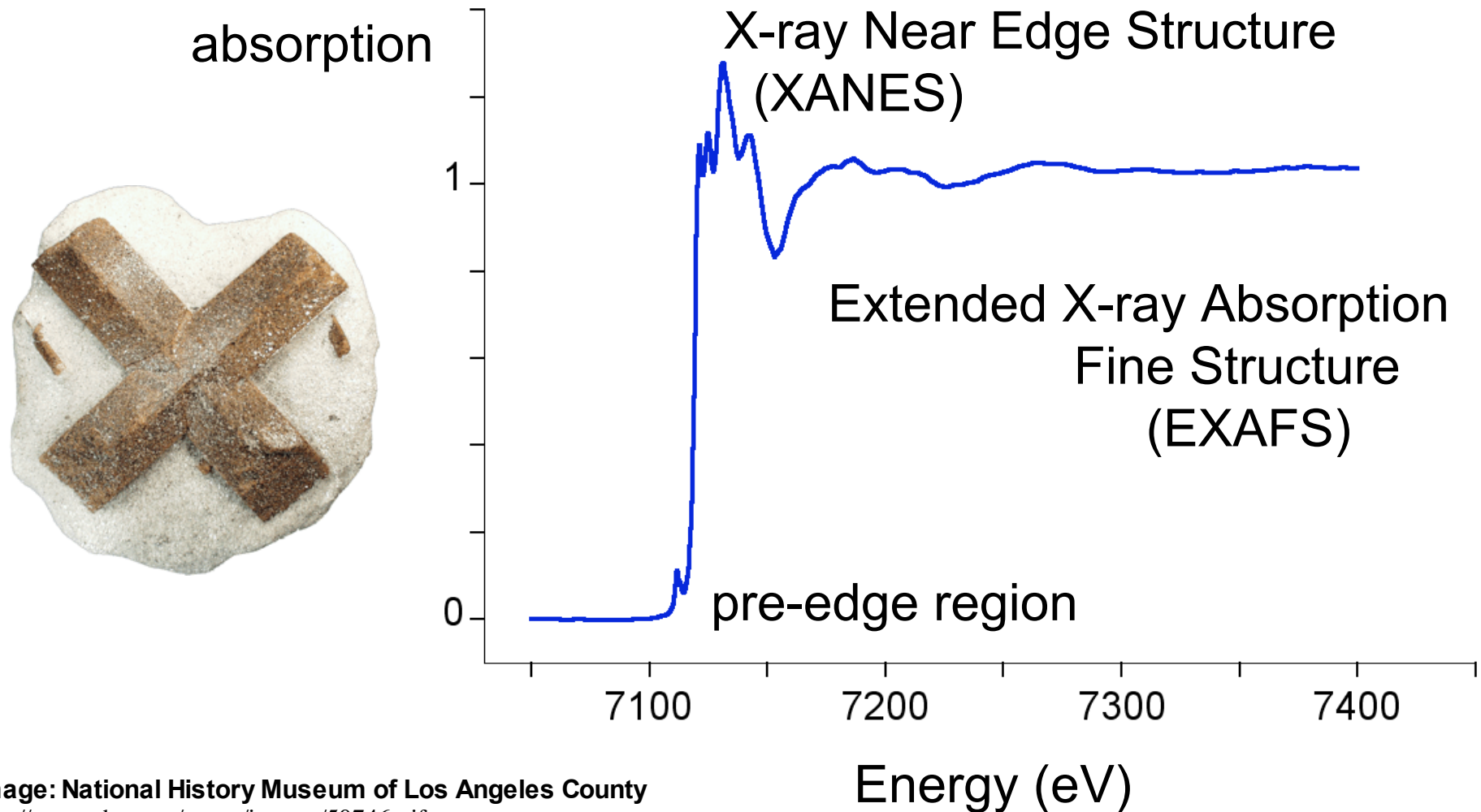
$\mu$



energy

# example

$\text{Fe}_2\text{Al}_9\text{Si}_4\text{O}_{22}(\text{OH})_2$  (staurolite) at the iron K-edge



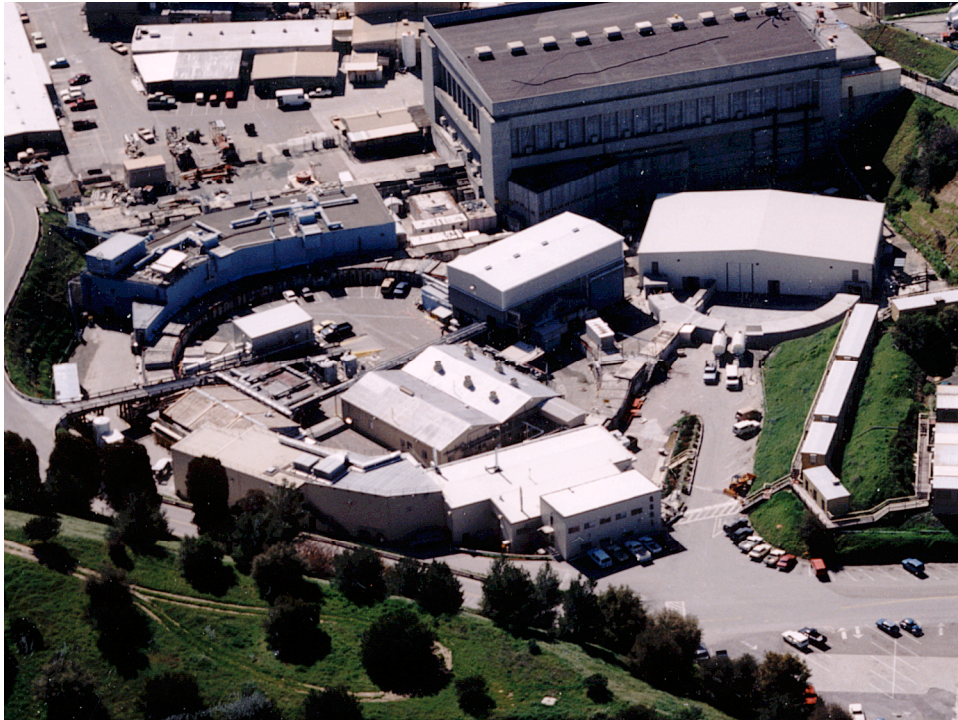


- **chemically selective**
- **available from boron ‘till actinides**
- **sensibility as down as ppm (in the best cases)**
- **sensitive to redox, speciation (incl. Its distortions)**
- **detect “medium-range environnement” (3-5 Å)**
- **works with “amorphous” matter, organic, wet, iron-rich matter**
- **OK with any natural/industriel “real” samples**
- **sample preparation usually minimal**
- **fewer beam/matter interactions (precious samples OK)**
- **very active community : theoreticians, sources, applications**

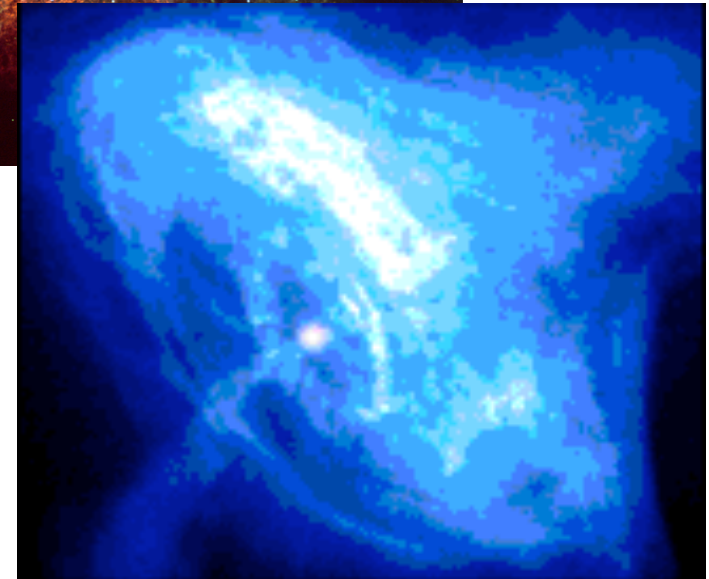
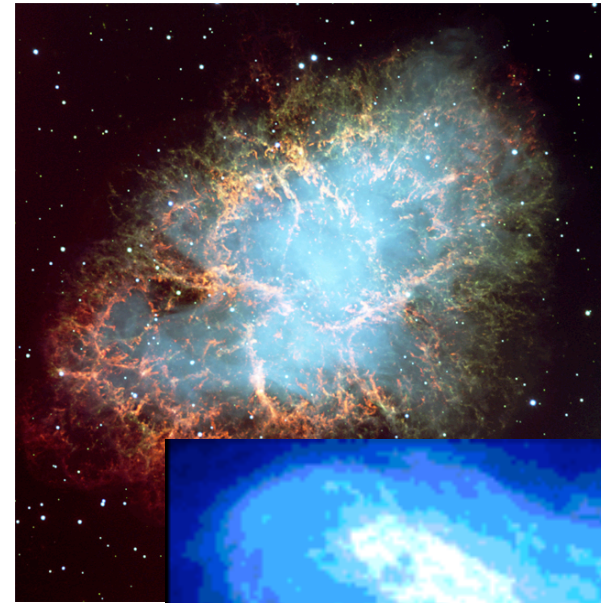


- **not sensitive anymore when disorder is too high**
- **not sensitive to neighboring ions with close Z-numbers**
- **access to large instruments is tough and highly competitive**
- **only a few beamtime days a year**
- **must be ready for experiment - no improvisation**
- **expensive, centralized - writing proposals and activity reports**
- **interface with theoricians not always easy**
- **spectrometers are often “home-made”, not standardized**
- **must read Phys Rev, Amer Min, JACS, GCA and many others**

**synchrotron = coherent source of x rays (among others),  
directional and intense**



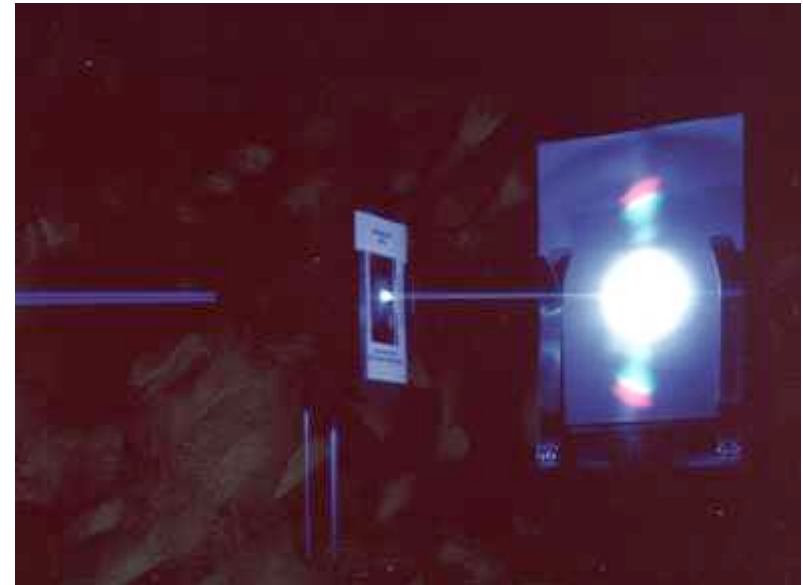
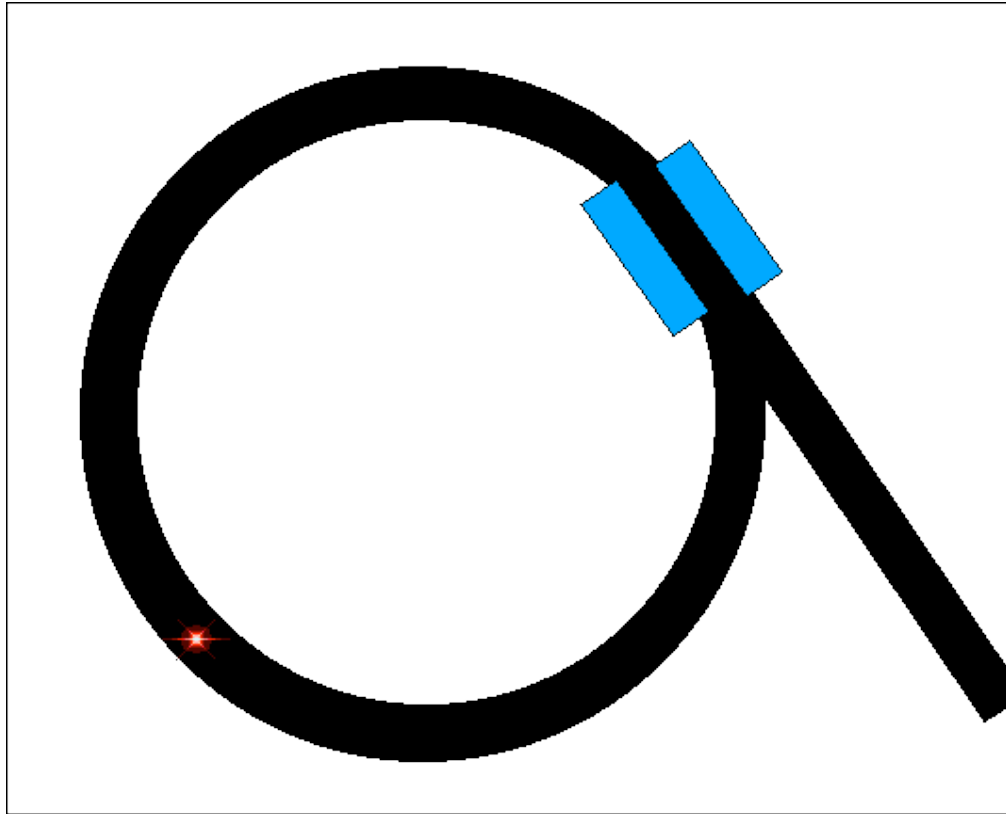
**SSRL**



**pulsar of the  
Crab nebulae  
(supernova of 1054)**

**principle :**

**Particle (positron) field = electrostatic field + radiative field**

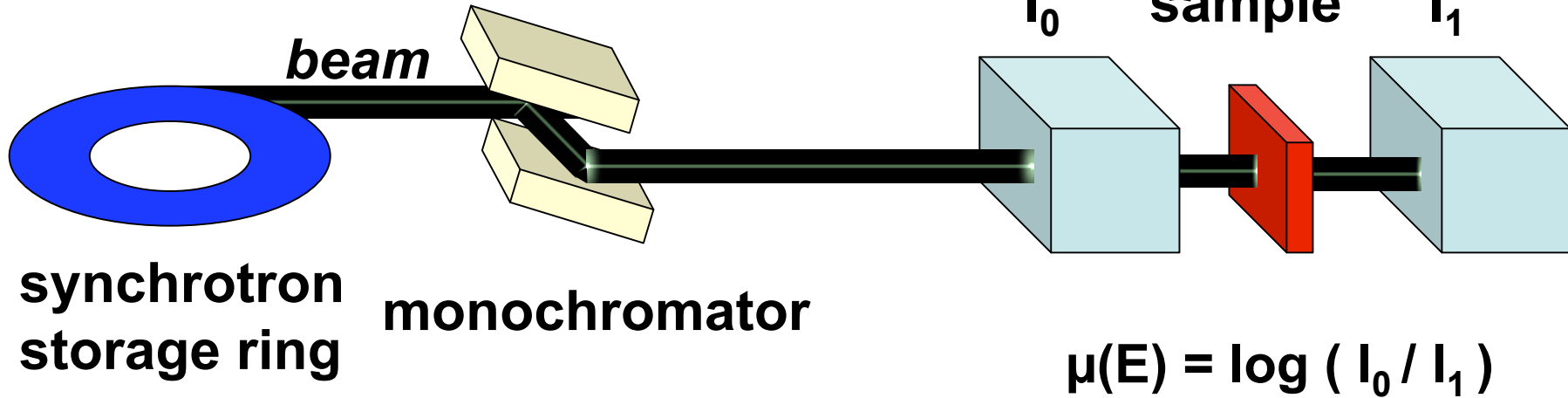


**under a magnetic field,  
the radiative field dissociate from the electrostatic field  
of the accelerated particule**

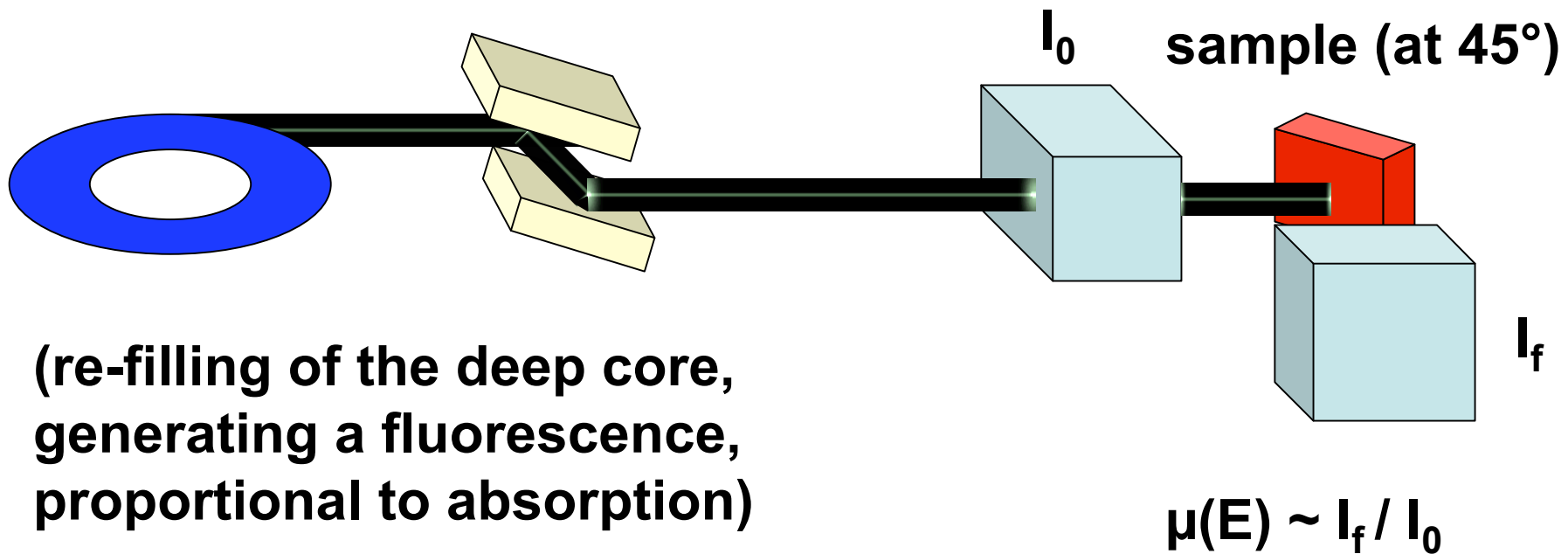


Spectra are collected in :

- the transmission mode



- the fluorescence mode



# Monochromators :



**Bragg's law**

**to select one ( $\pm\Delta E$ ) energy**

**from a continuum of radiations from the ring**

**gratings (100-2000 l/mm)**

**beryl**

**YB<sub>66</sub>**

**KTP**

**Quartz**

**Si (111)**

**Si (220)**

**Si (311)**

**Si (511)**



# Monochromators :



[www.accel.de](http://www.accel.de)

**flux**



**(111)**

**(220)**

**resolution**

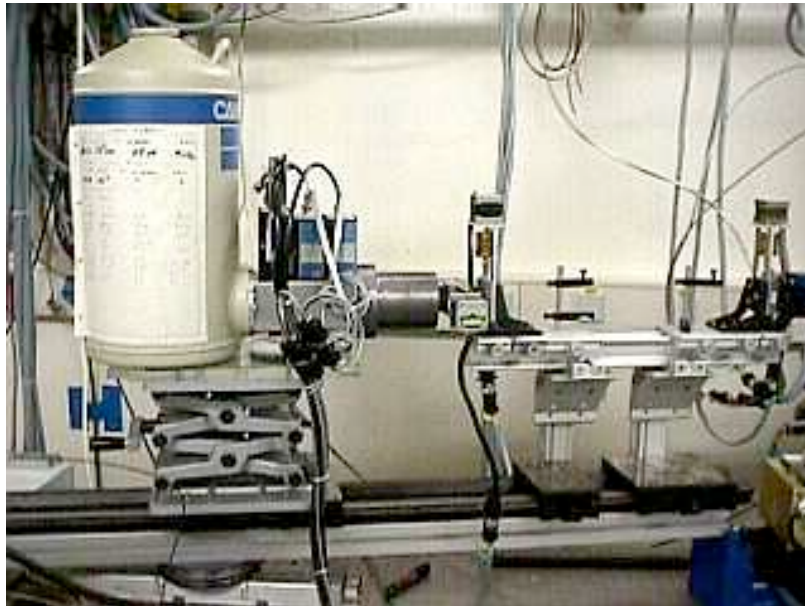


**(311)**

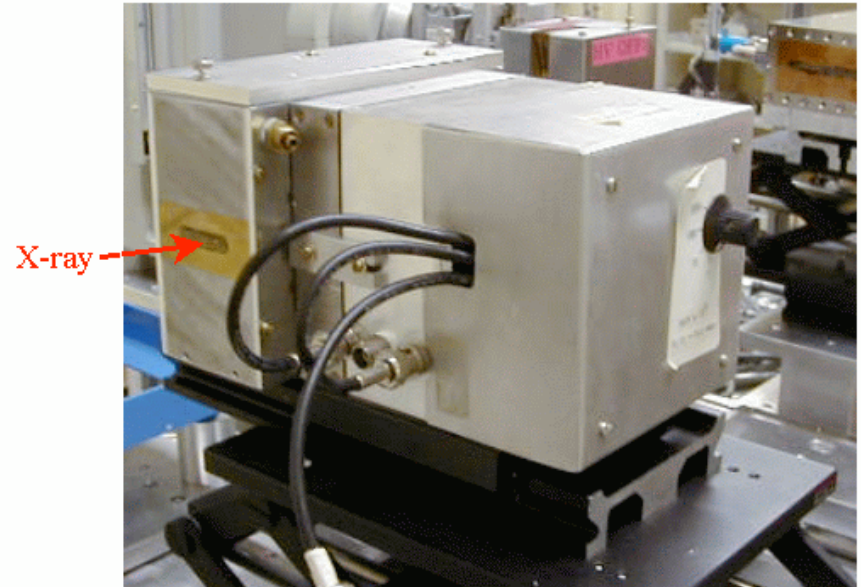
**(511)**

## Detectors :

**Gaz chambers : *flexible, easy***



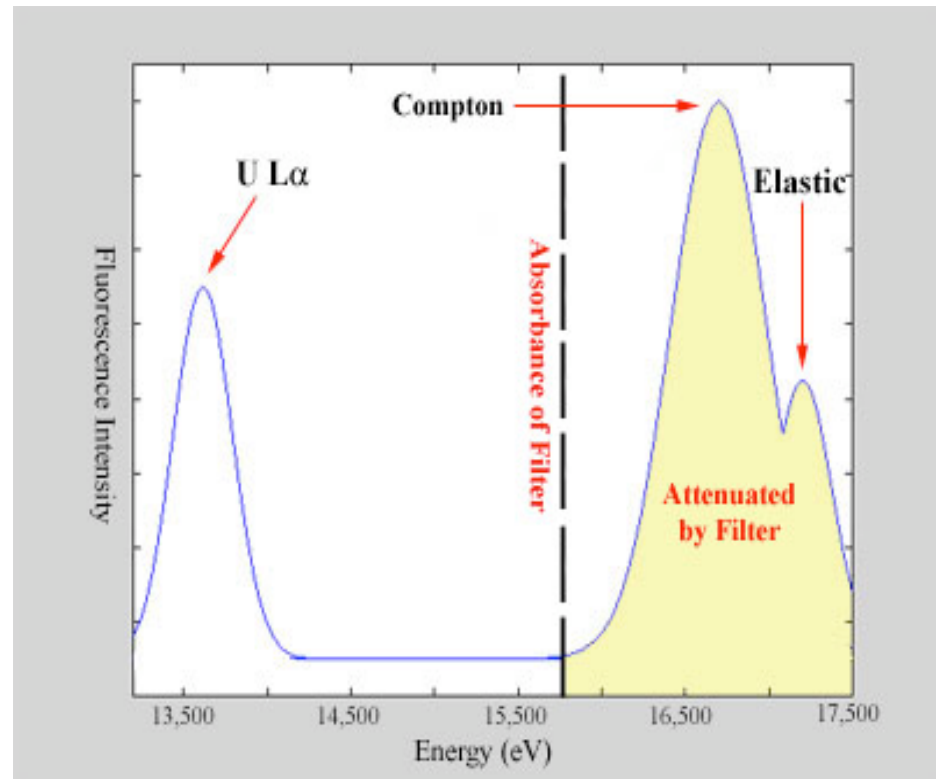
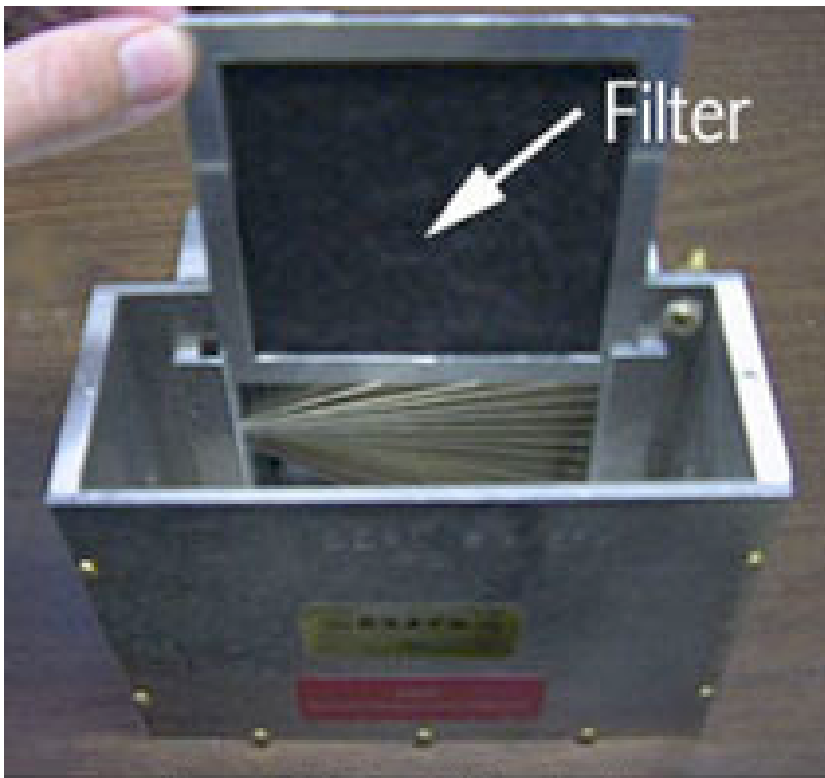
**Si-strips : *the future, all in one***  
***(no cooling)***



**Solid state detectors :**  
***more sensitive, expensive***



- data collection “in fluorescence” (“through the fluo”):

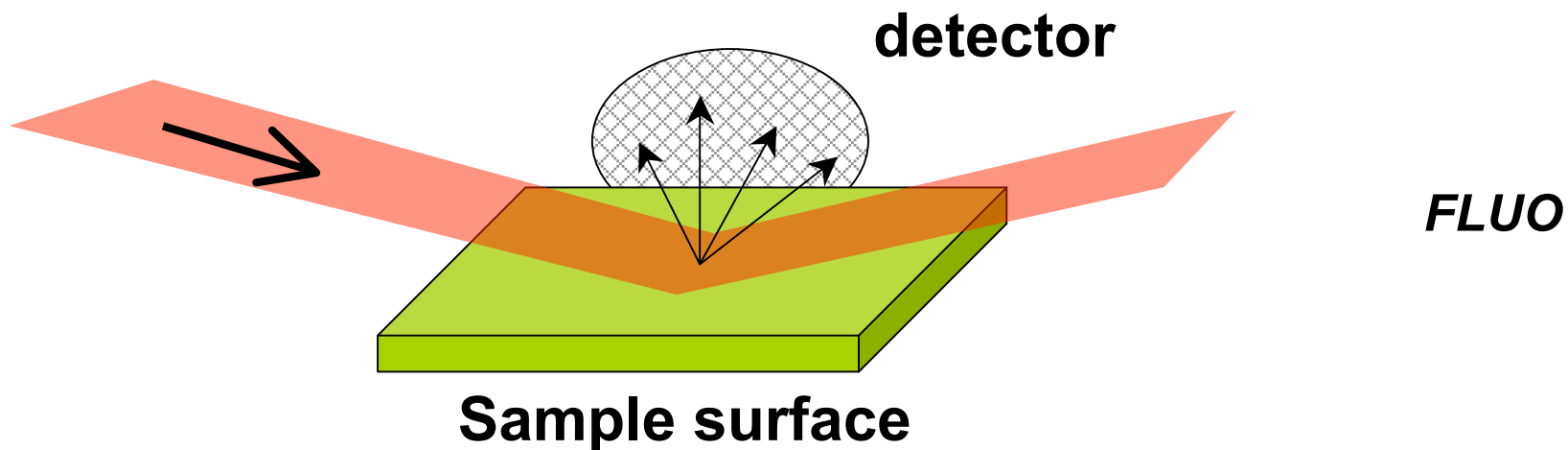


*MES 11-2 SSRL beamline documents*

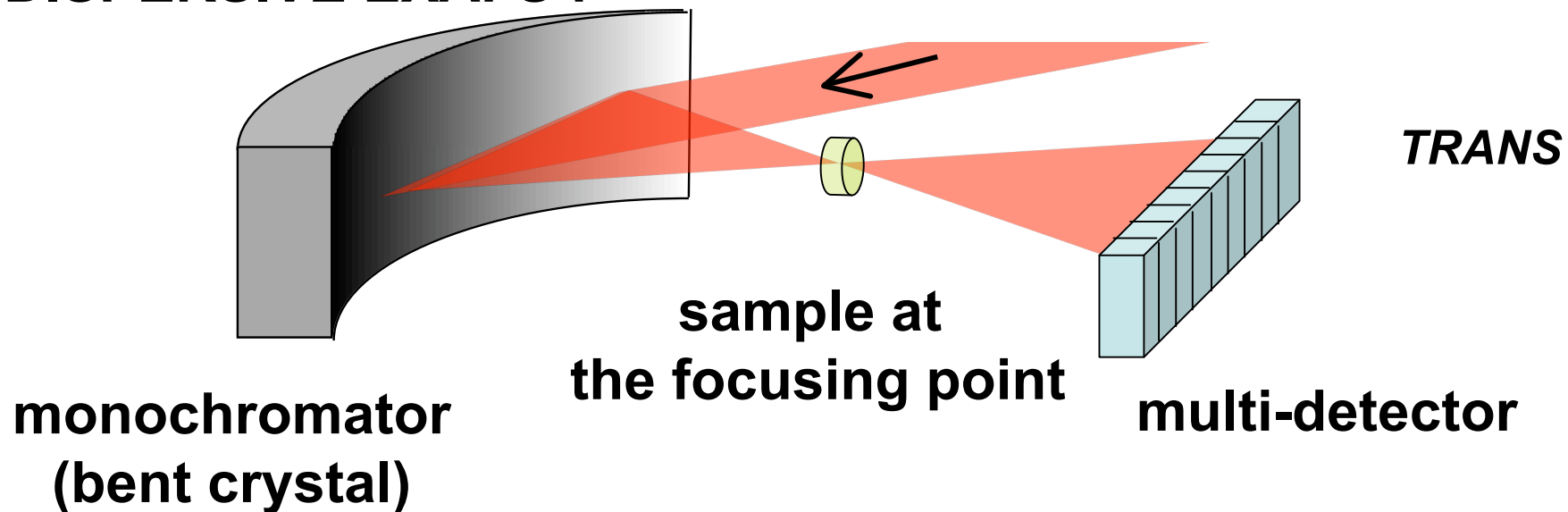
- intensity of the fluo. is proportional to the absorbed x-rays
- use a filter to cut unwanted “elastic scattering”

## OTHER MODES :

- REFLEXAFS (SEXAFS / GI-EXAFS) :



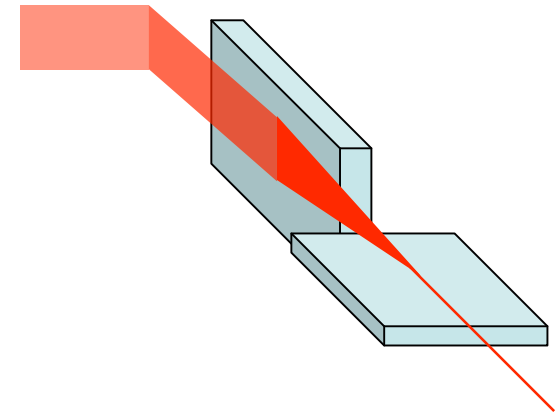
- DISPERSIVE EXAFS :



## top : microXAFS (and nanoXAFS...)

- various focusing devices :

- Kirckpatrick-Baez (“KB”) mirrors
- capillaries
- Fresnel zone plates



- advantages

- lateral resolution down to the  $\mu\text{m}$
- coupling chemical mapping / speciation
- 3D XAFS tomography

- problems

- auto-absorption at the point of impact (correctible)
- baseline distortions (correctible)
- high quality EXAFS is harder to get
- local concentration rather high ( $> 0.1 \text{ wt. } \%$ )



## **II. Analysis & data reduction in real**

**I) corrections**

**II) pre-edge**

**III) edge**

**IV) EXAFS**

# (1) Deadtime corrections

*(in fluorescence, using solid state detectors)*

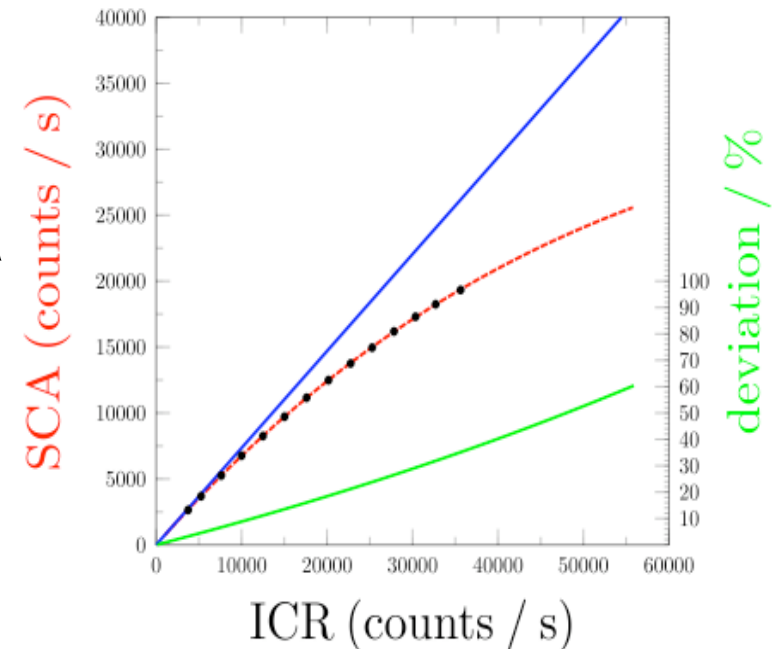
It is a fluorescence correction factor,  
measured for each energy available on the  
spectrometer used

Minor but important intensity corrections

Can significantly reduce noise in your data

Specific to each beamline, often negligible  
(negliged), sometimes really worth to  
consider !

Ask your beamline scientist about that  
(even if he/she hates...)



Konarev et al. (1999)  
Hasylab-Desy


## 2) non-linear baselines :

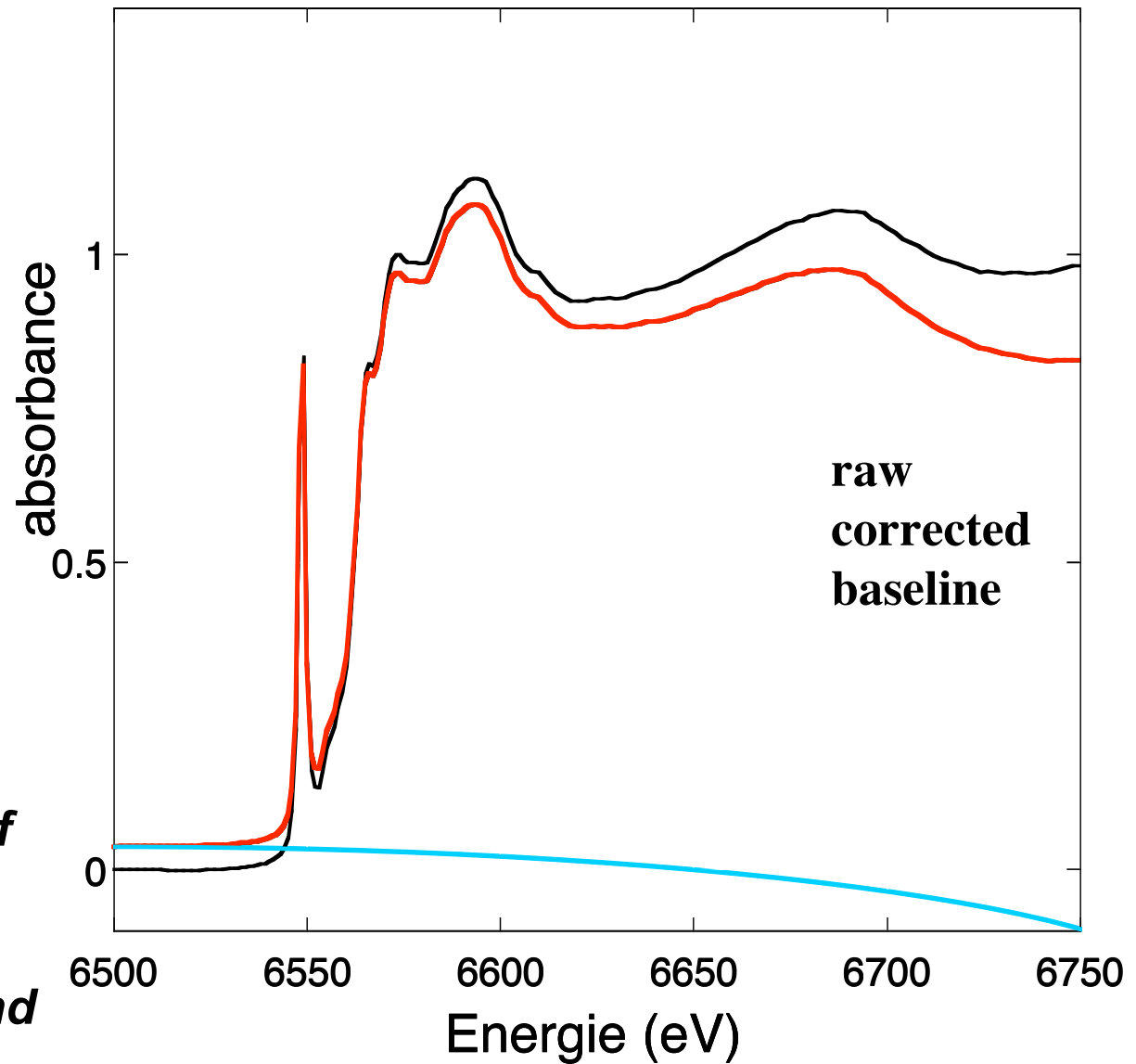
-  $\mu$ -XAFS

- use of solid state detectors in some cases

self-made code

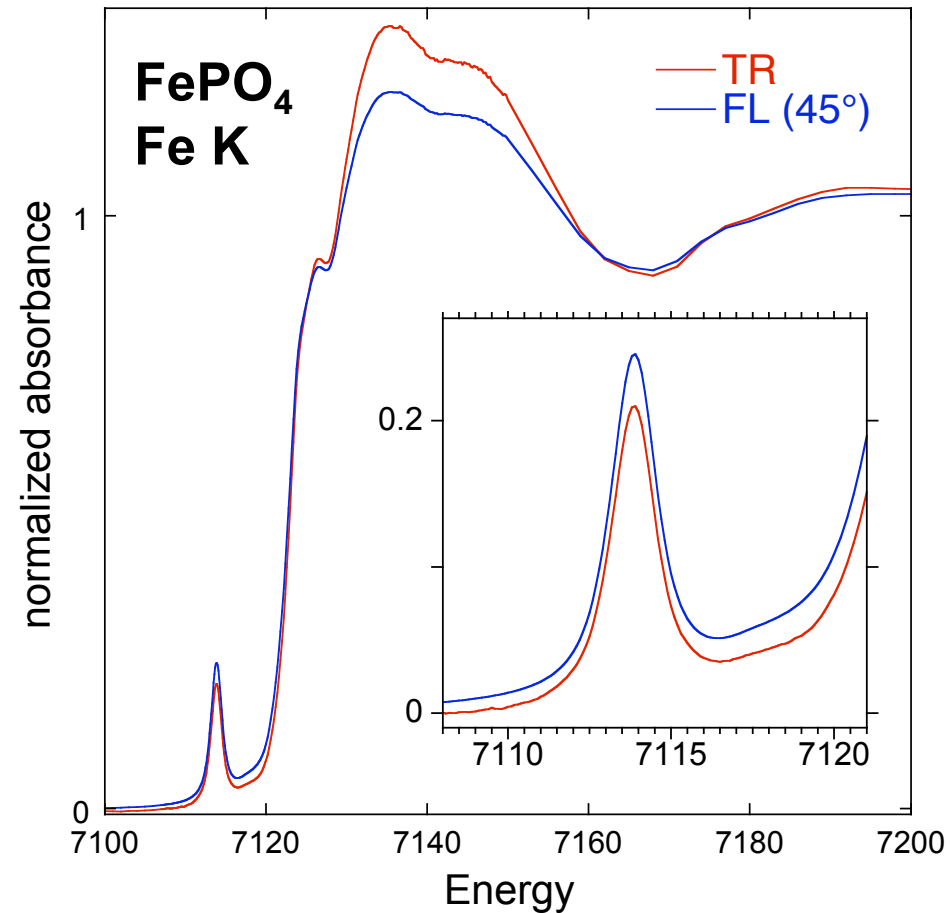
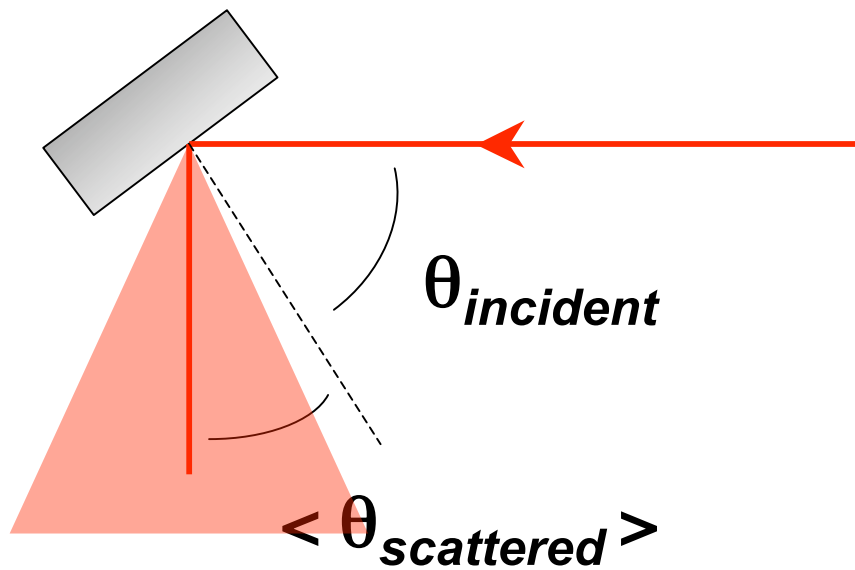
using combination of polynomials 

to "flatten" before and after the edge 



### 3) Self-absorption corrections

(in fluorescence)



Haskel et al., 1999 :

$$\frac{I_f(\omega)}{I_0(\omega)} = \mathcal{E}_f(\omega) \frac{\Omega}{4\pi} \frac{\mu_e(\omega)}{\mu_t(\omega) + \mu_t(\omega_f)} \frac{\sin \theta_{incident}}{\sin \theta_{diffusé}}$$

***S.A. occurs when :***

***45° geometry***

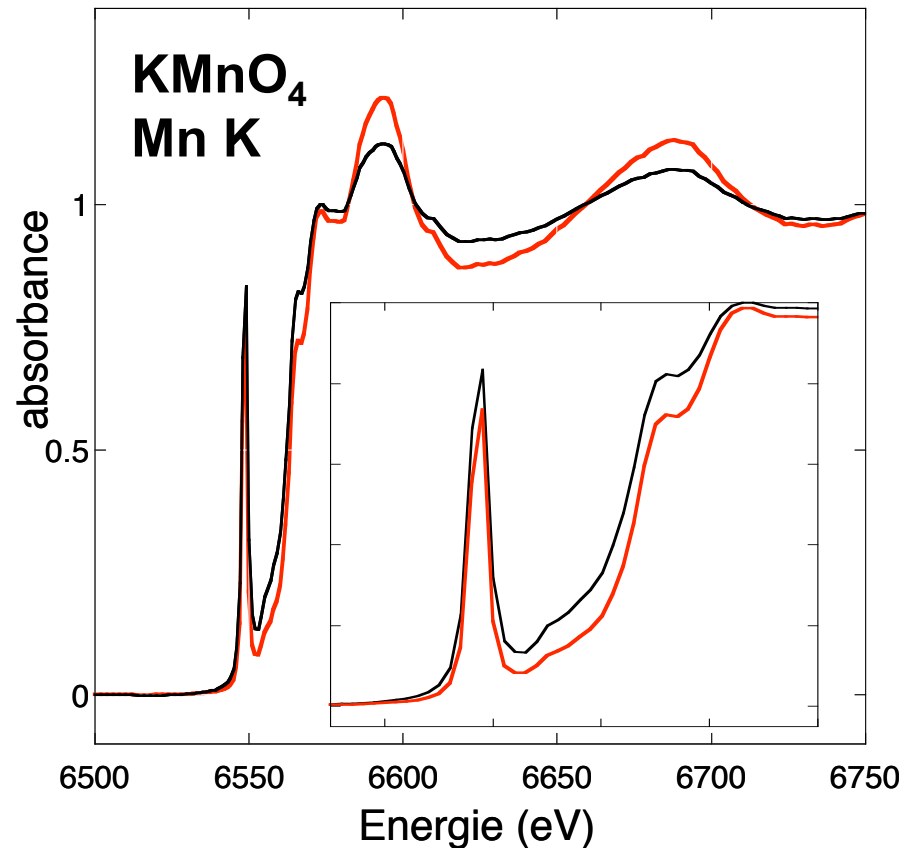
***+***

***Absorbing atom is  
more than ( $\pm$ ) 20 mol.%  
in sample on the sample  
surface illuminated***

***Tricks :***

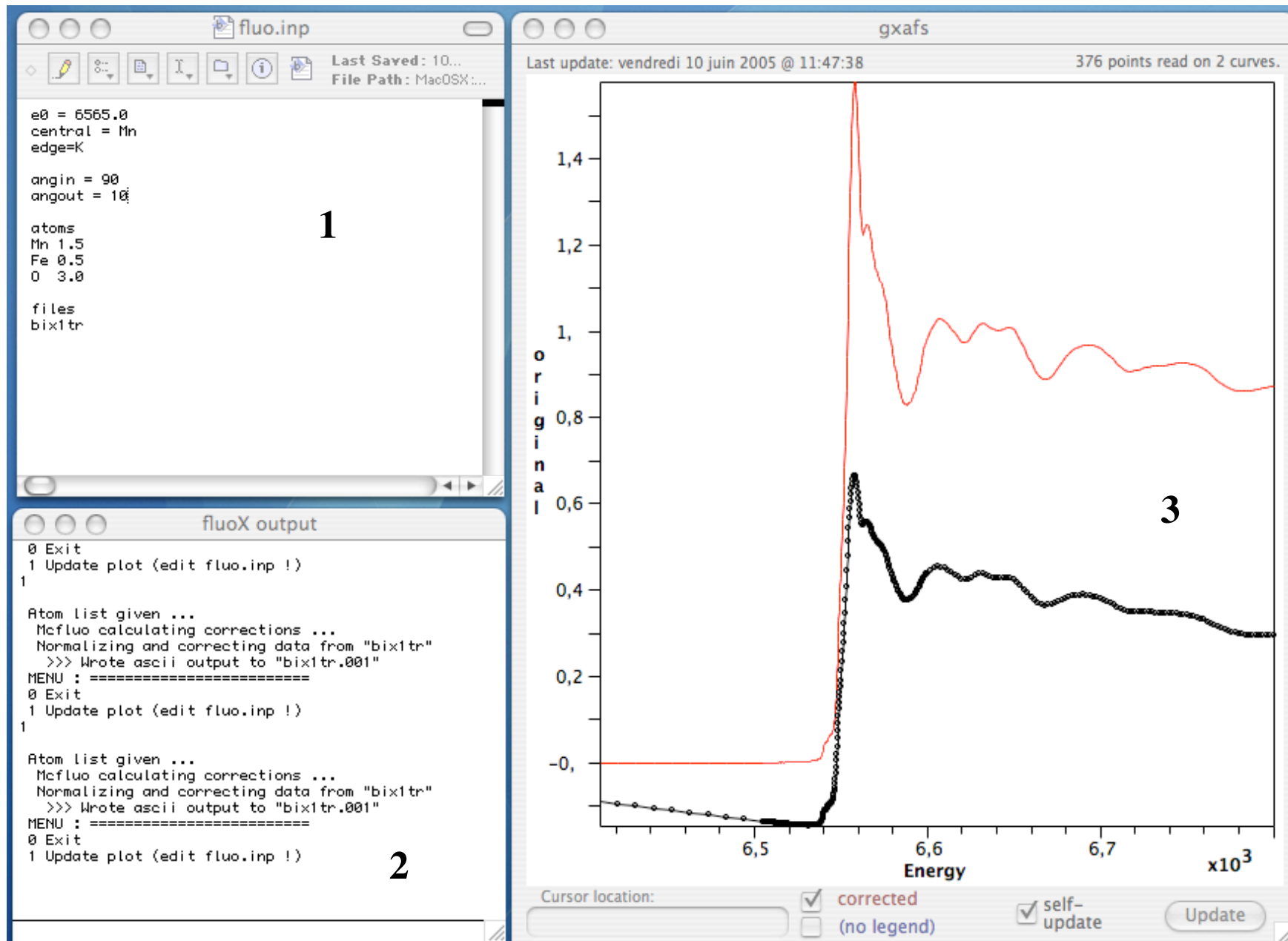
***Use either 90° OR 45° geometries***

***Diluting in a non-absorbing matrix won't help !!***



# FLUO-X (home-made, with graphical interface)

1-2-3

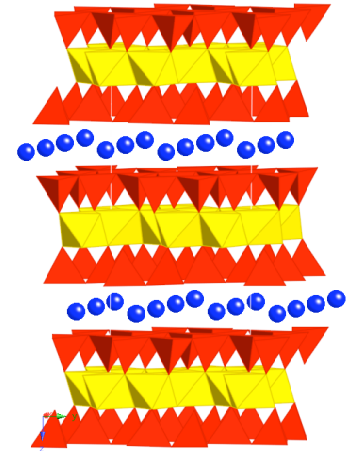


## 4) Polarization artefacts :

atoms along the polarization direction of the X-ray beam are preferentially probed (and  $\pm$  attenuated)

important for layered structures :

*micas, clays, phyllosilicates ...*



angle-dependant approach :  $\alpha$

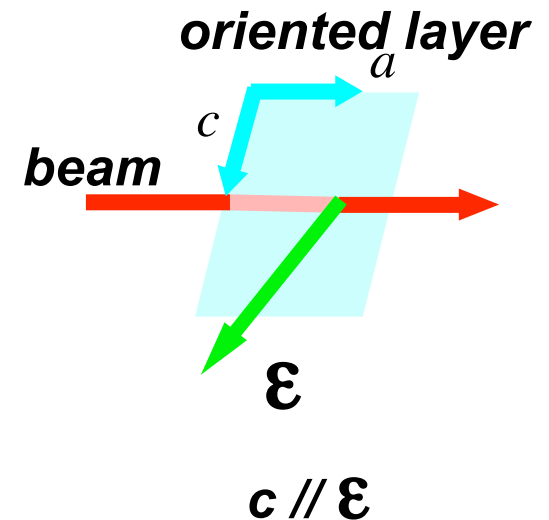
angle between the  $\mathbf{\epsilon}$  vector and the **layer plane**

yields P-corrected amplitudes

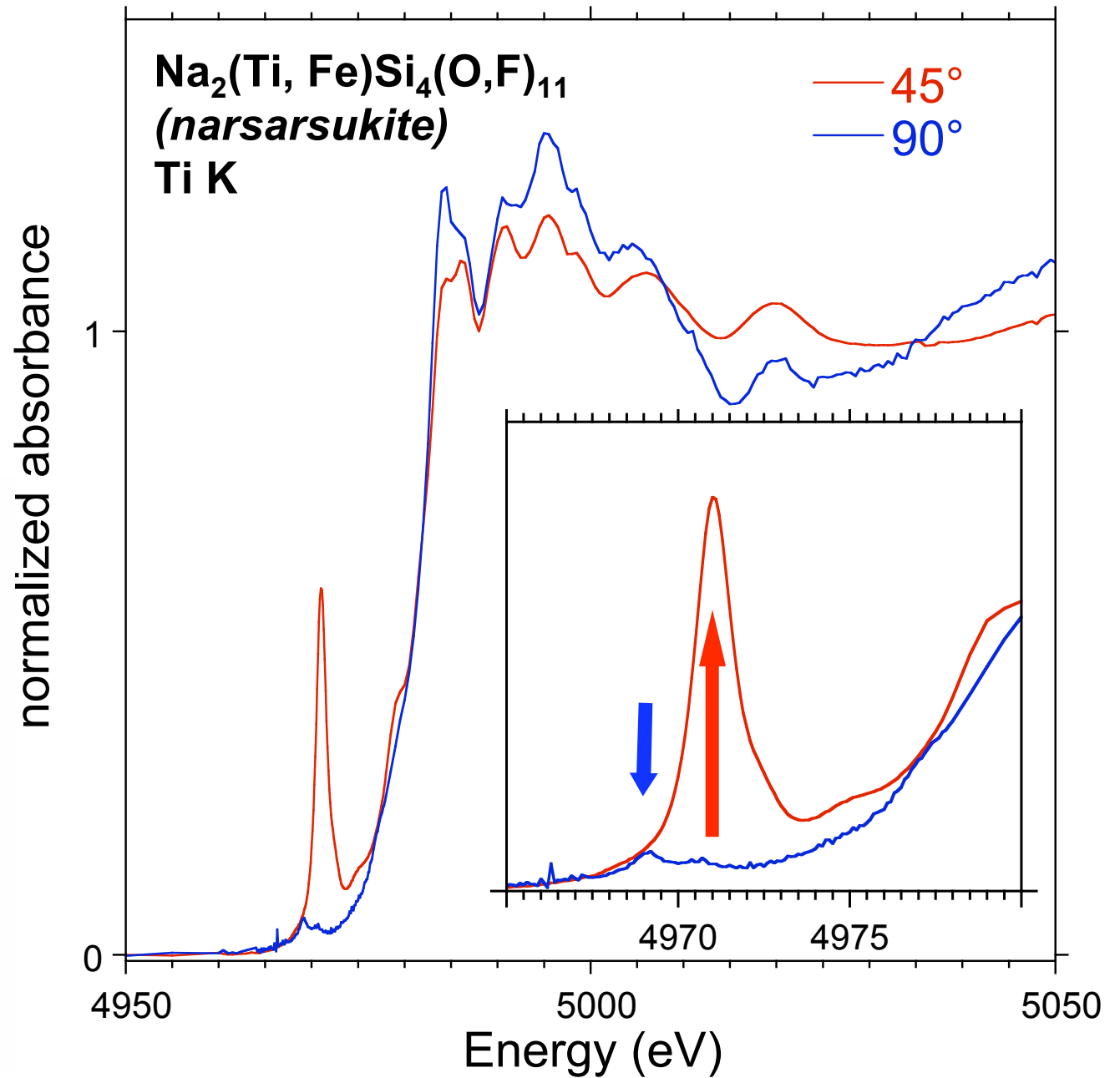
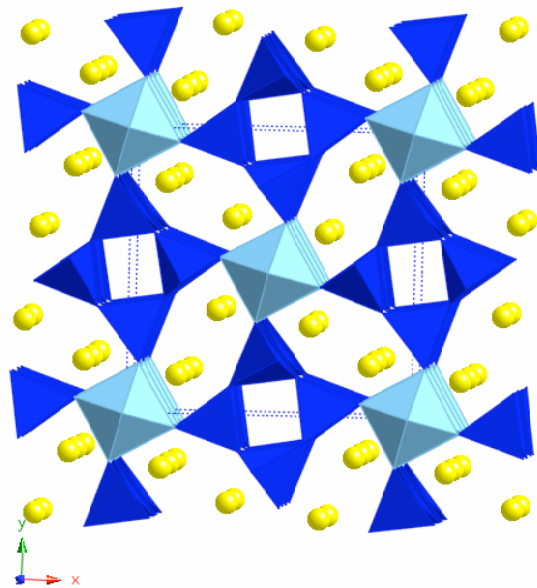
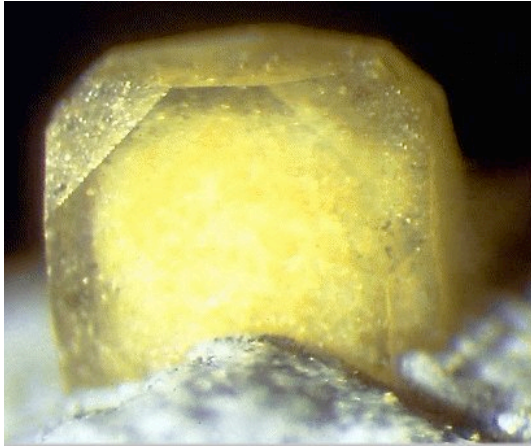
*Manceau et al. 1988, 1990; Manceau and Schlegel, 2001*

help discriminate between neighbors

can be applied to textural analyses

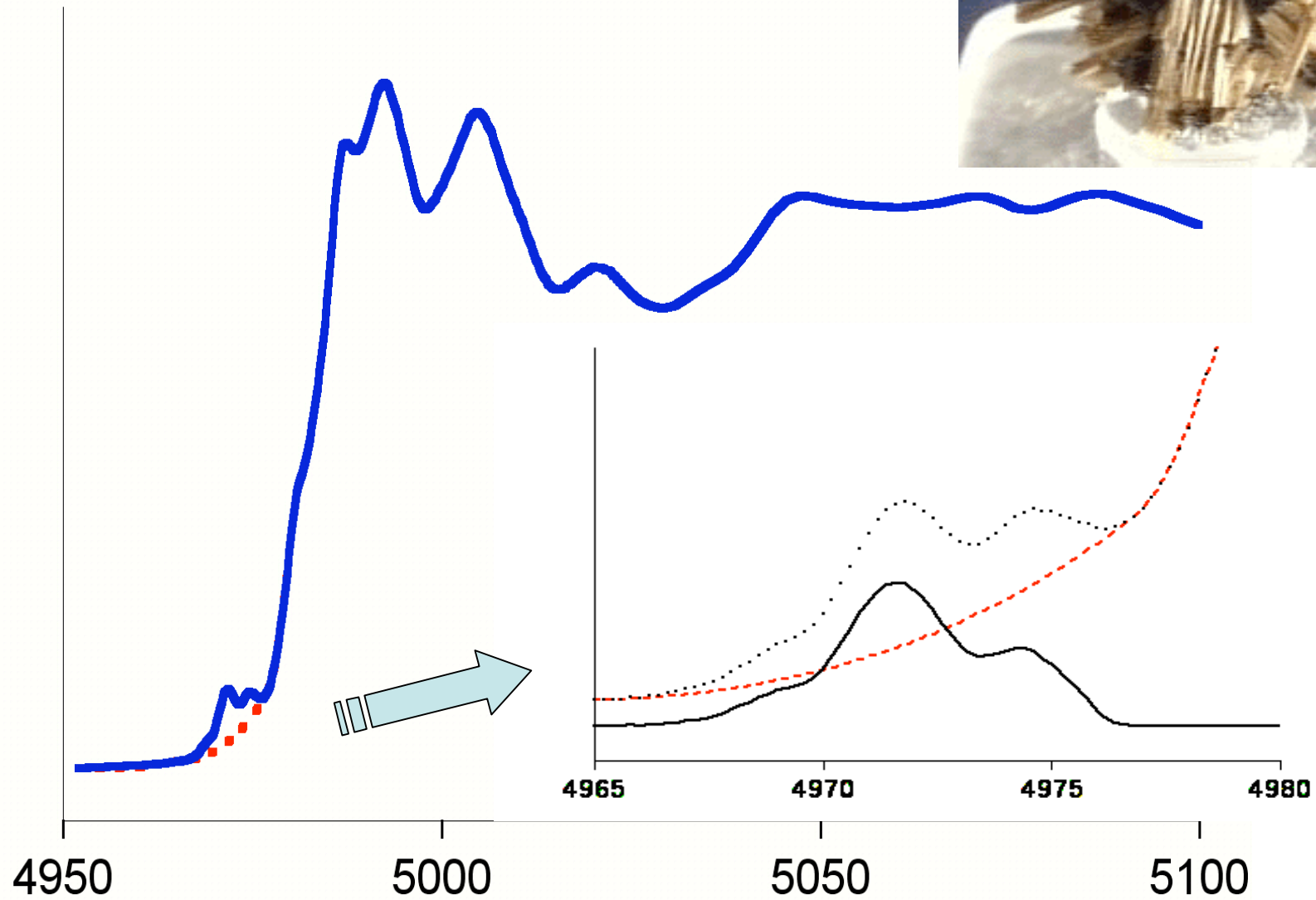
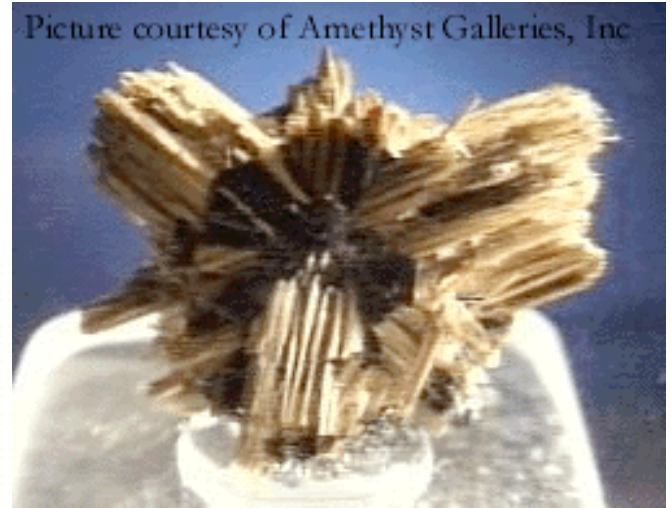


# WARNING : self-absorption vs. polarization

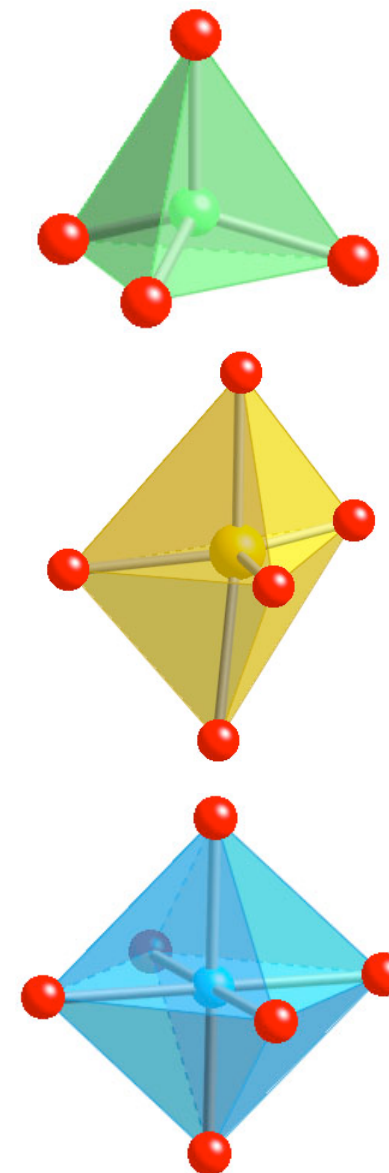
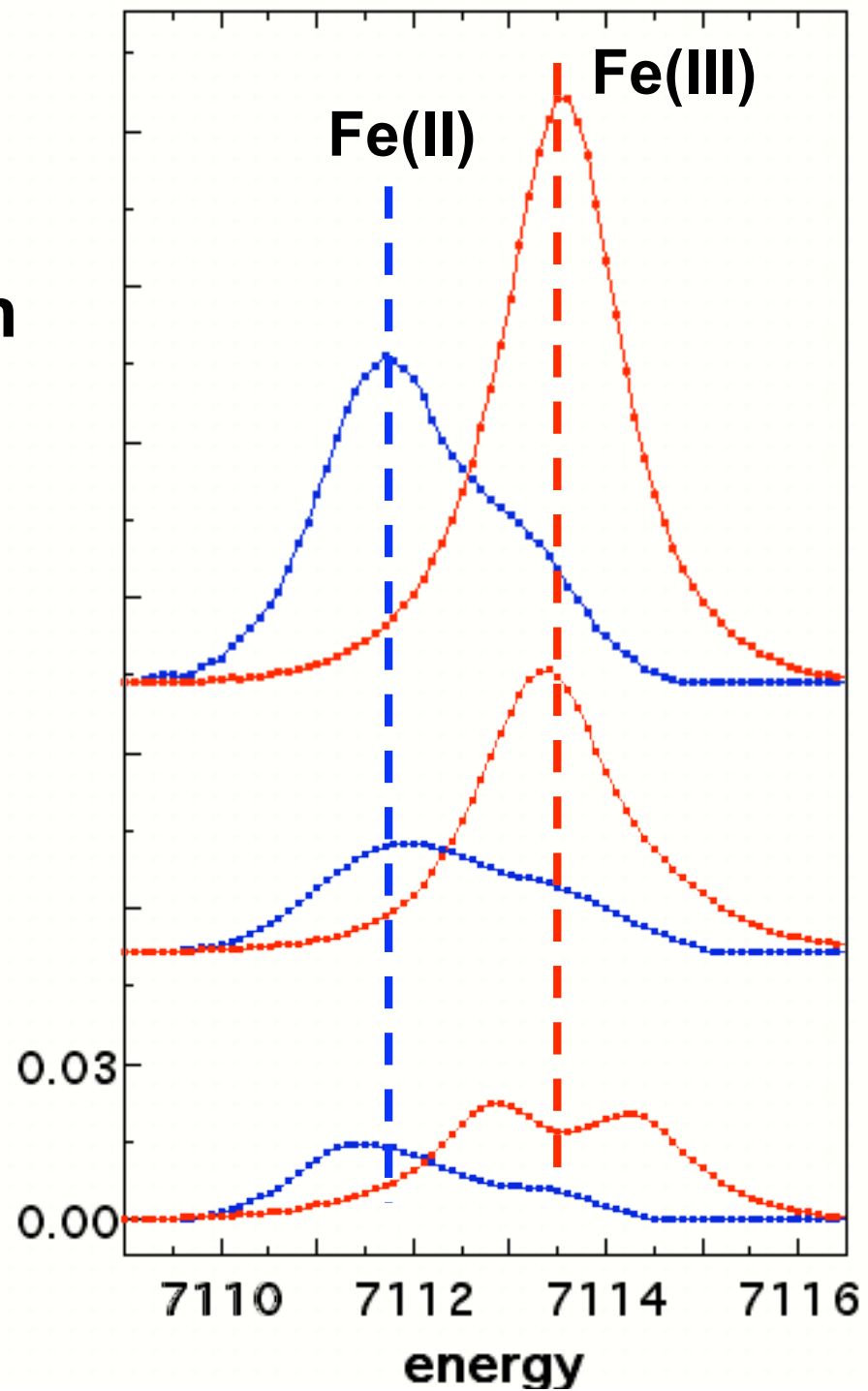


## **II) pre-edge region**

rutile ( $\text{TiO}_2$ ) at the Ti K-edge :



- variation with
  - redox
  - symmetry



## Modeling :

- « peak-fitting » or deconvolution

- number of lines to fit ?

- with which shape ?

- ◆ core hole : *lorentzian*

*tabulated by Krause & Oliver (1979)*

- ◆ experimental resolution

(mono, source size....) :

*gaussienne*

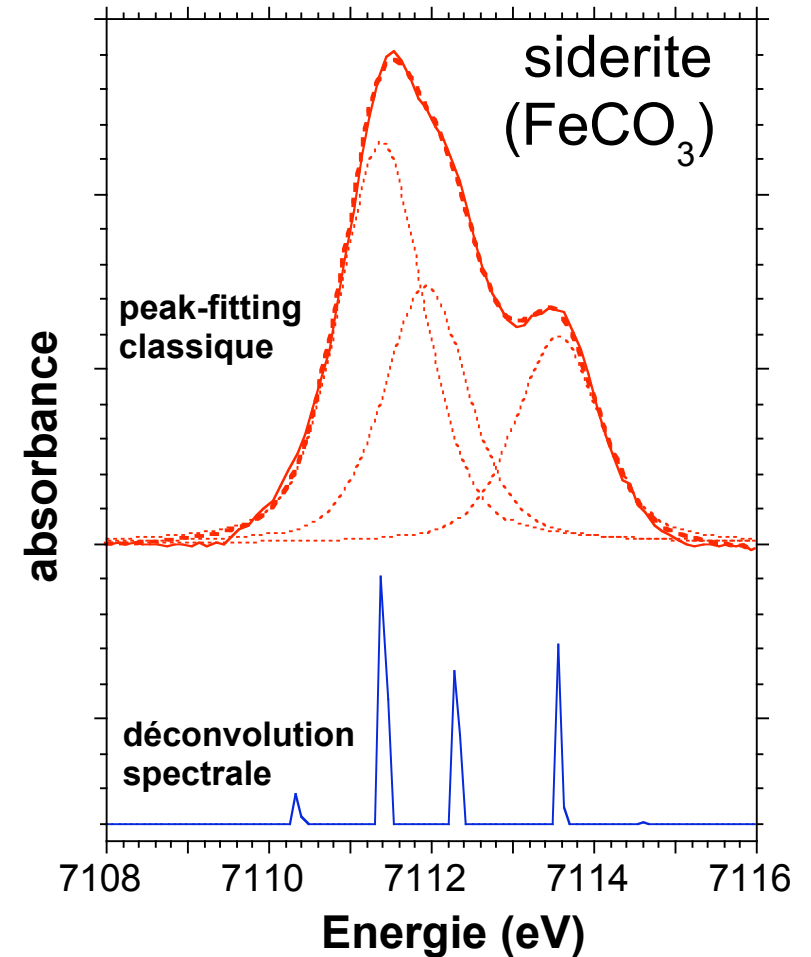
*measured by the beamline scientist*

true Voigt ~ pseudo-Voigt  
(sum or product L/G)

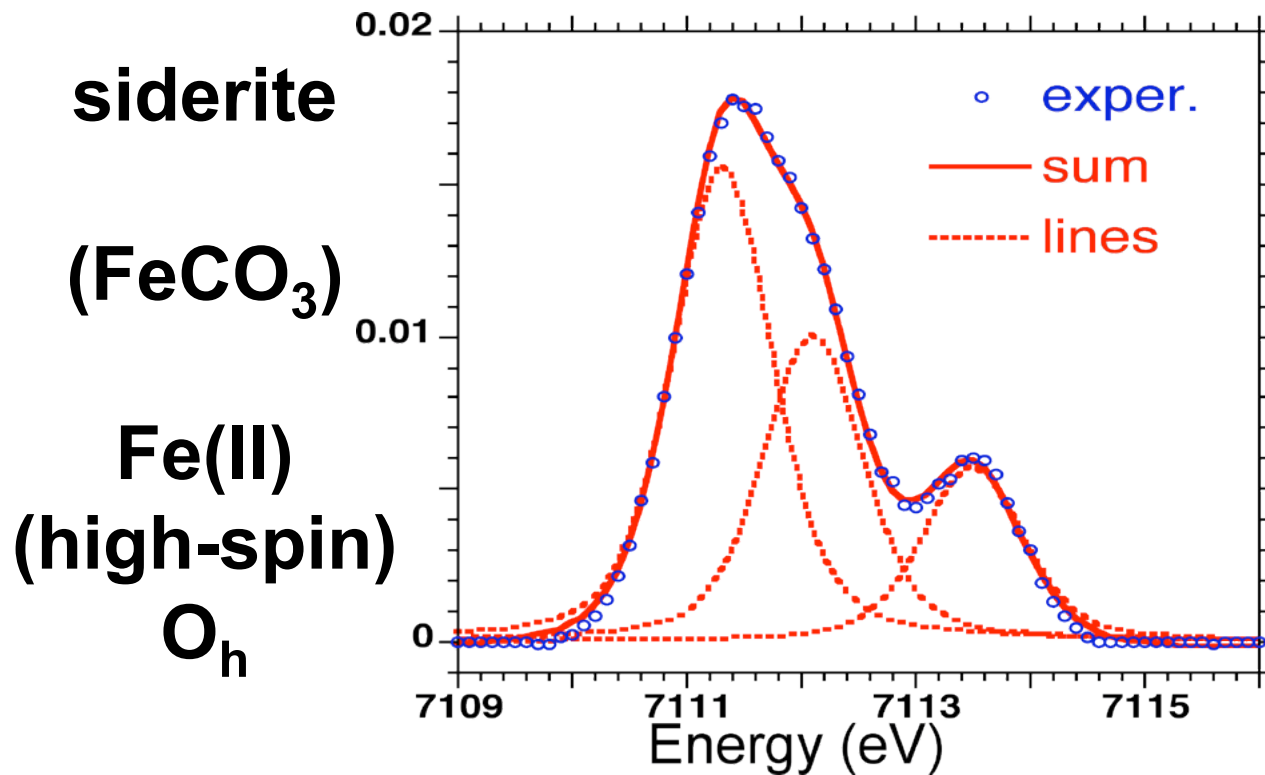
% Gaussian ?

width ??

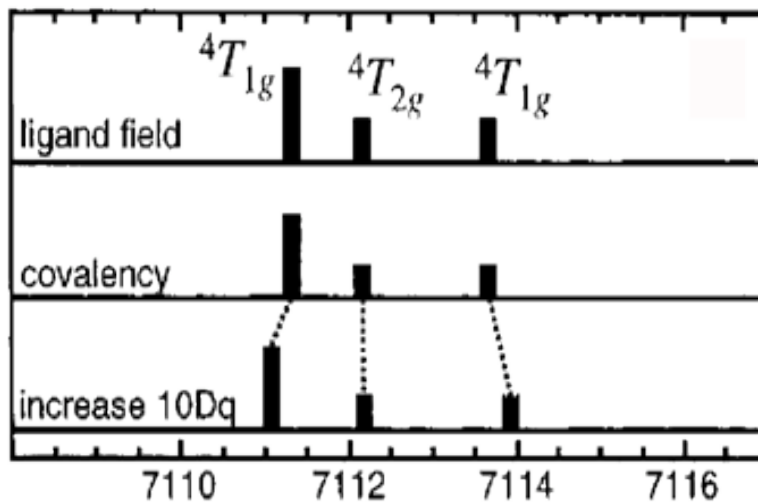
intensity ???



# correlation with theoretical models (multiplets):

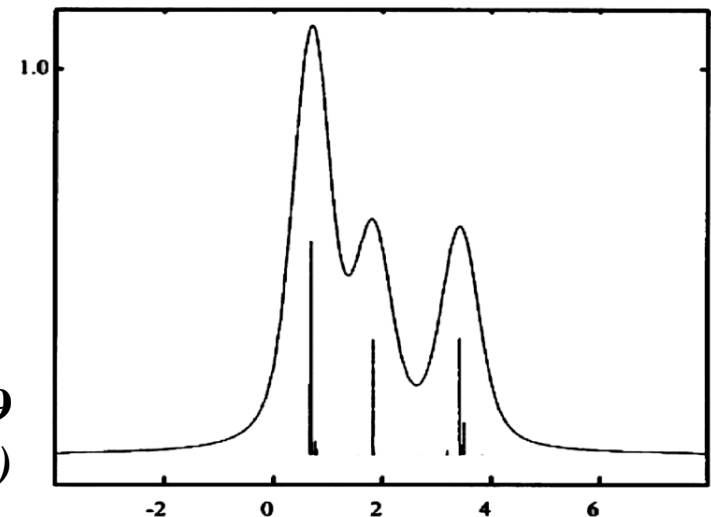


Siderite, mine "Pffannenberger Einigkeit",  
Salchendorf, Siegerland, Germany. Height 4 cm.  
Collection: Mineralogisches Institut Universität  
Mainz. Photo: © Rainer Bode, Haltern



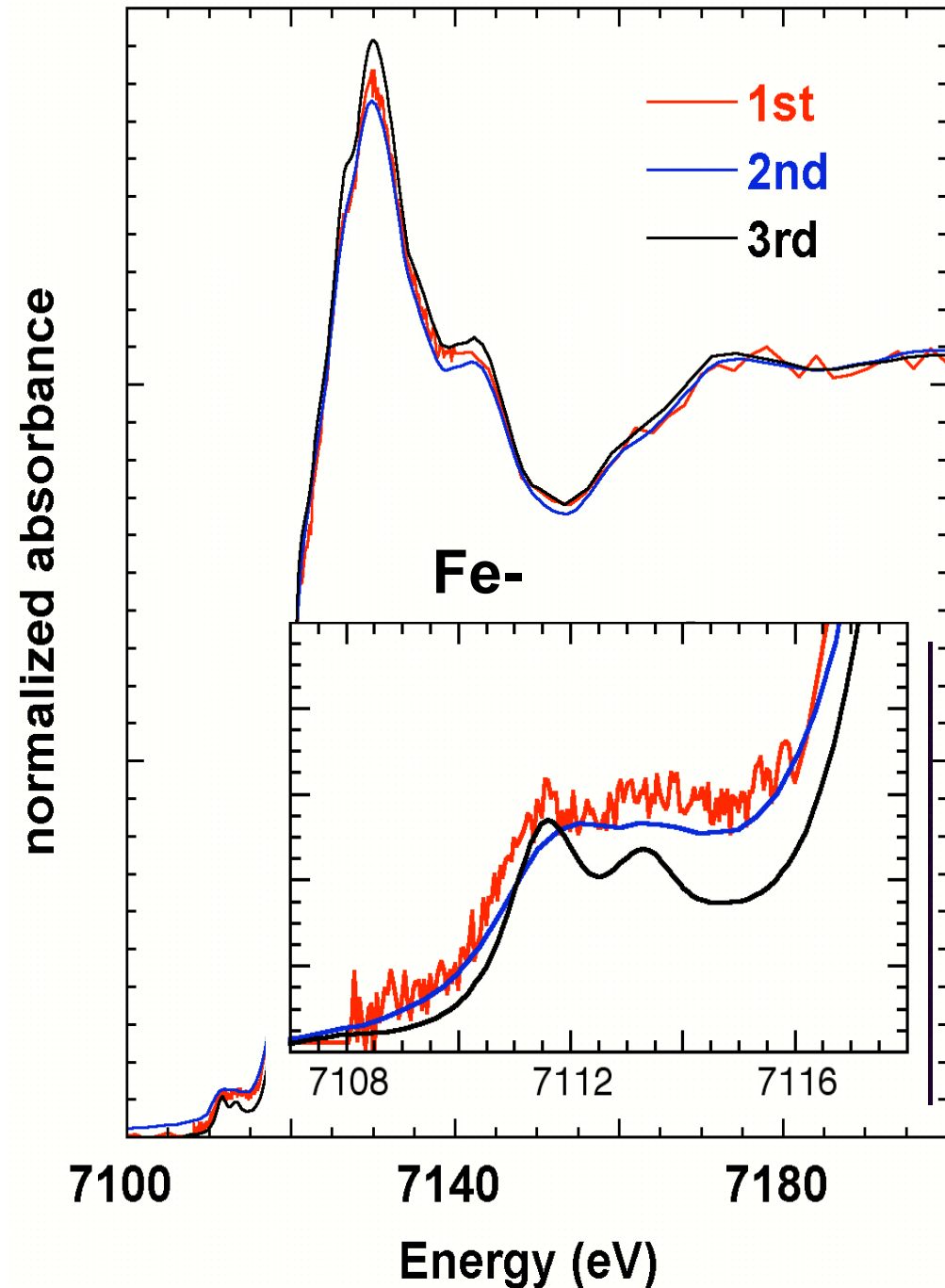
Westre et al. 1997  
JACS

de Groot, 1999  
(pers. com.)



## warning :

- need energetic resolution
- high signla/noise ratio
- baseline tough to model
- some features are still not well understood
- measure, as usual with XAFS, an average redox

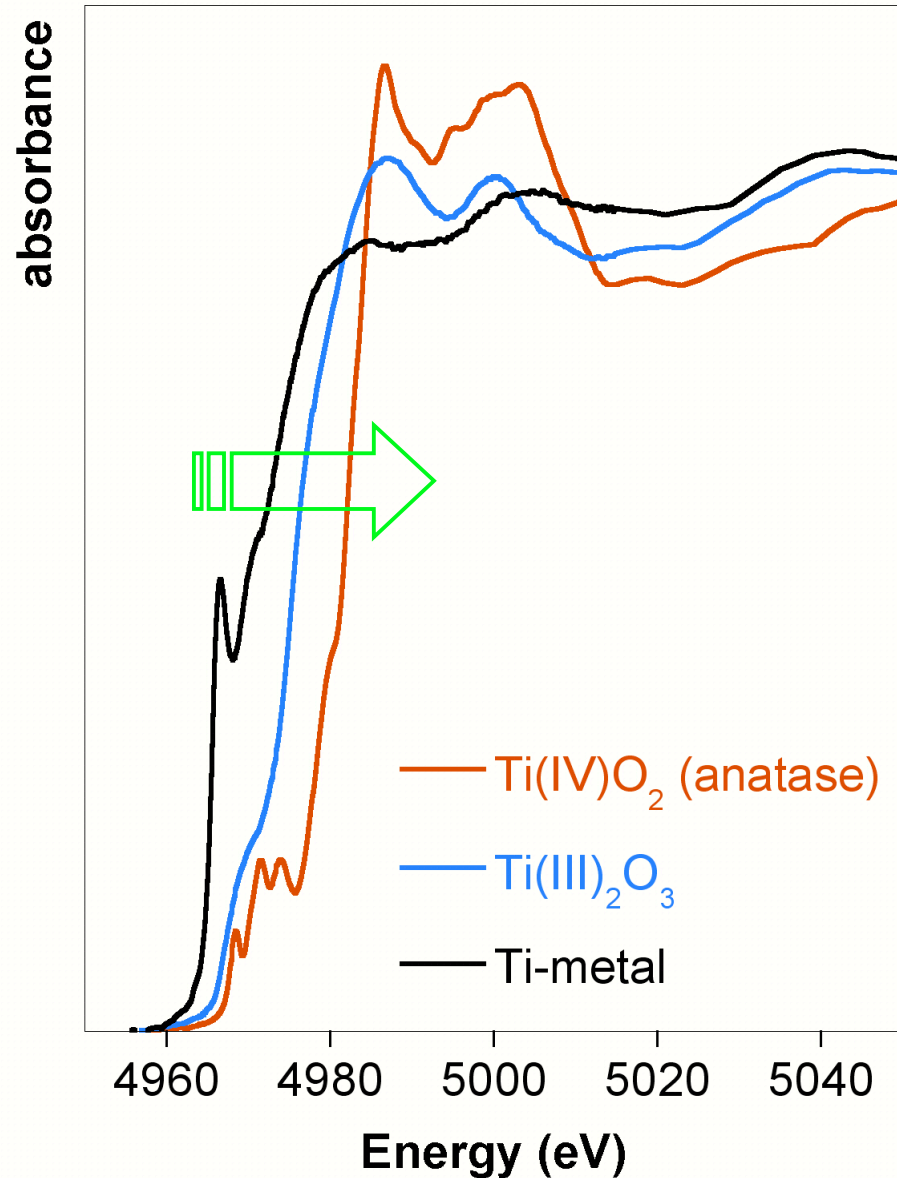


### 3) Absorption edge :

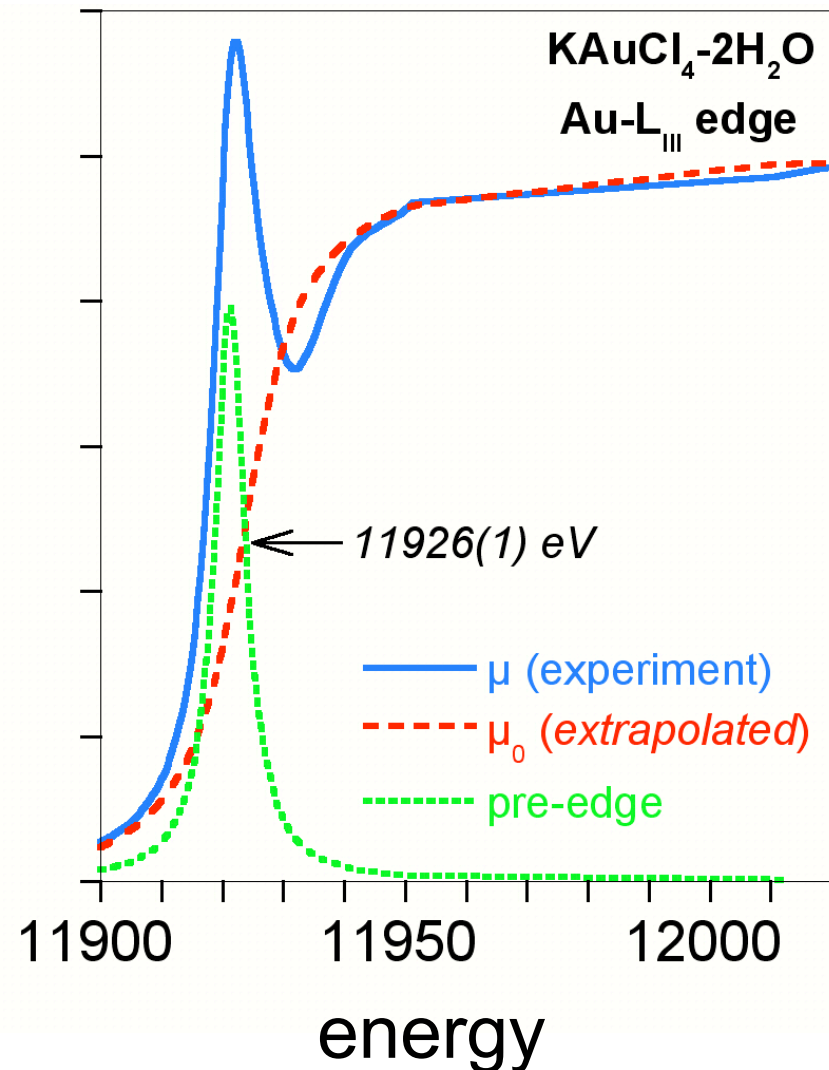
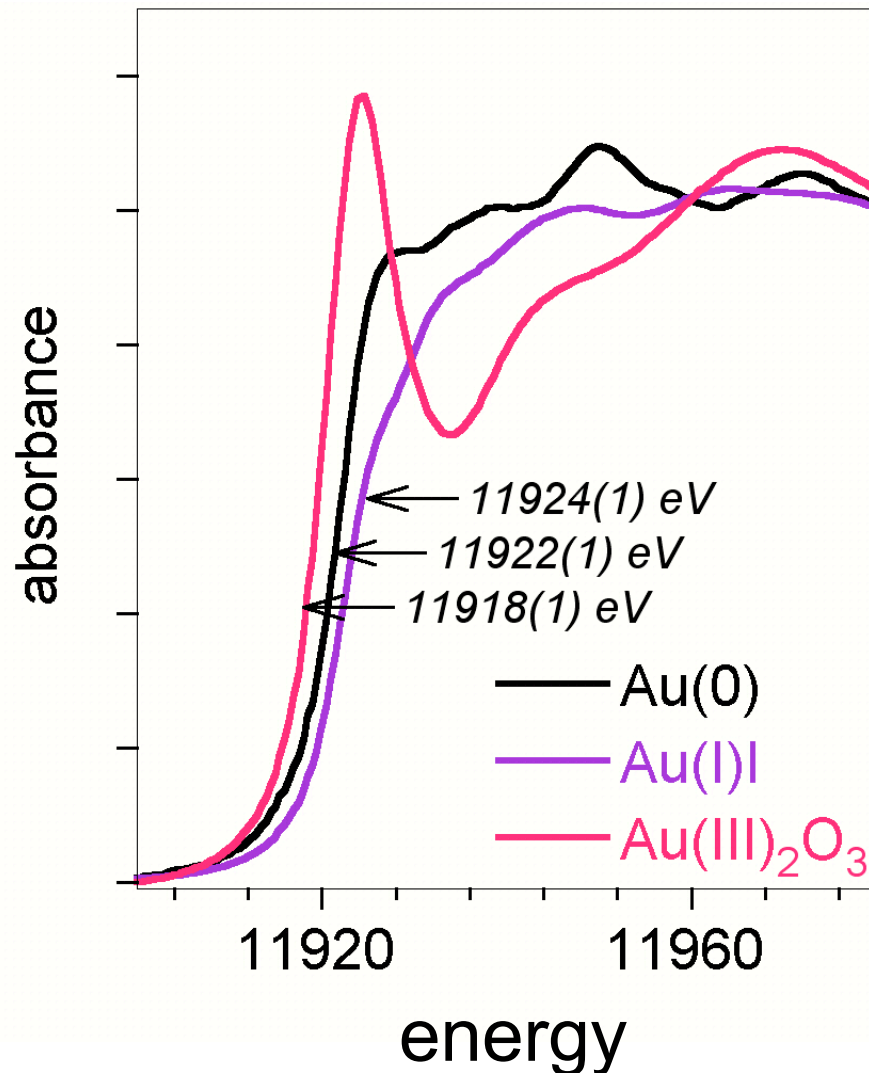
The edge position

is sensitive to redox

(roughly  $\sim 1$  eV par e-)



# Exemple:

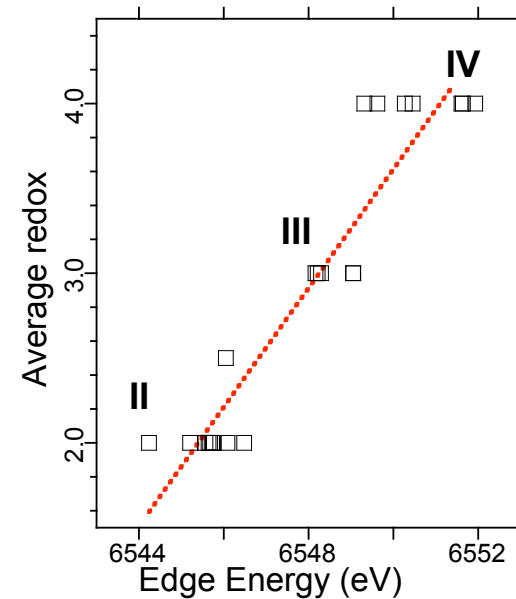
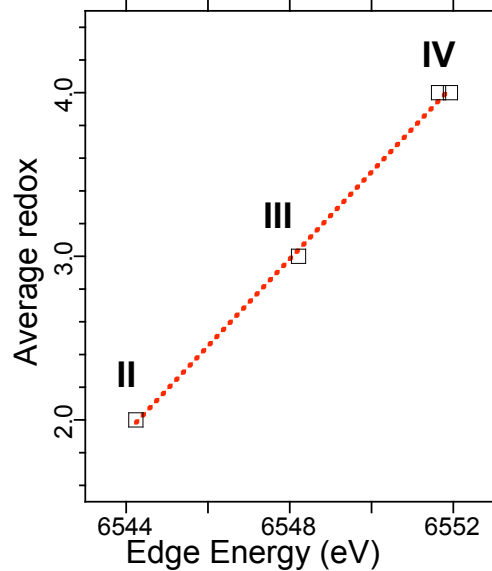
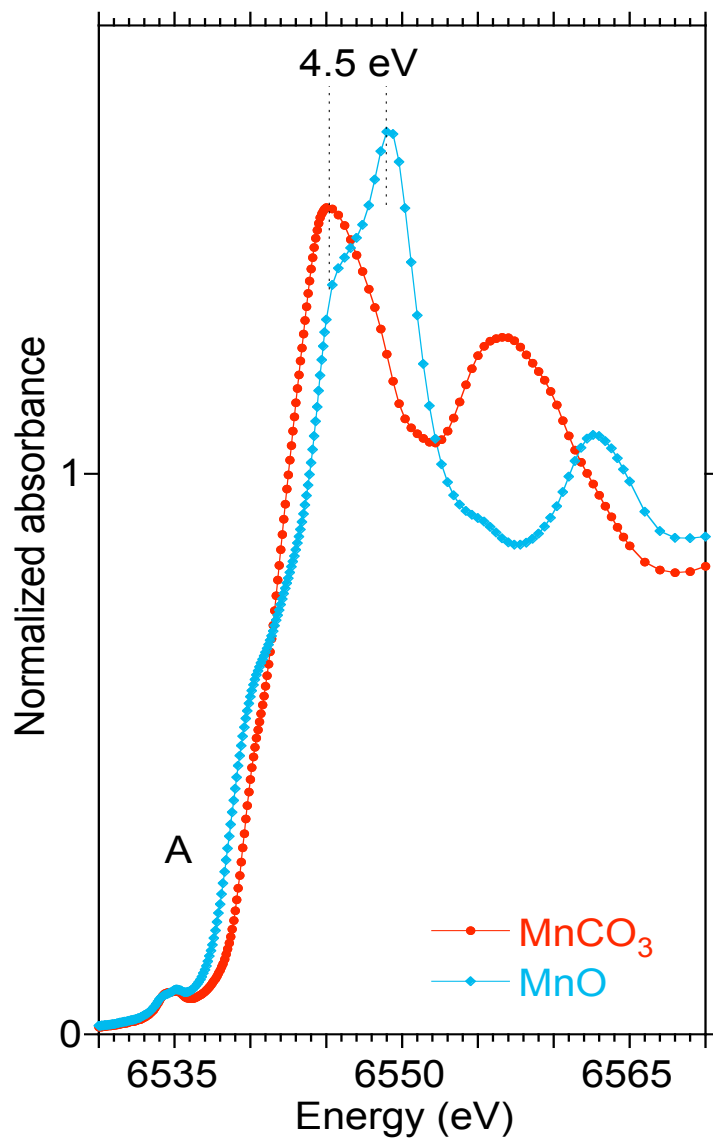


Edge of Au(III) BEFORE that of metal ...

=> The edge is, in fact, a gigantic pre-edge !

**Bad :**

# Inventing ourselves new junky edge theories or protocols

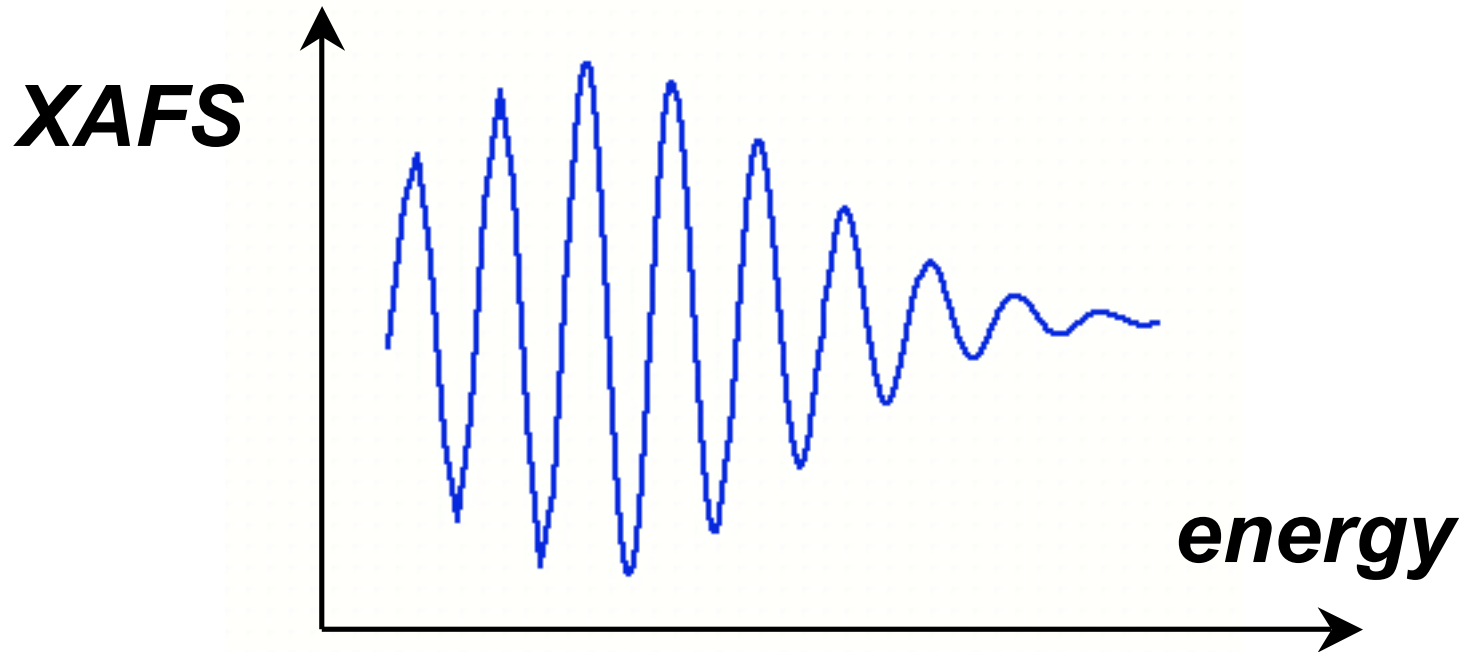


## 4) XAFS Interferences

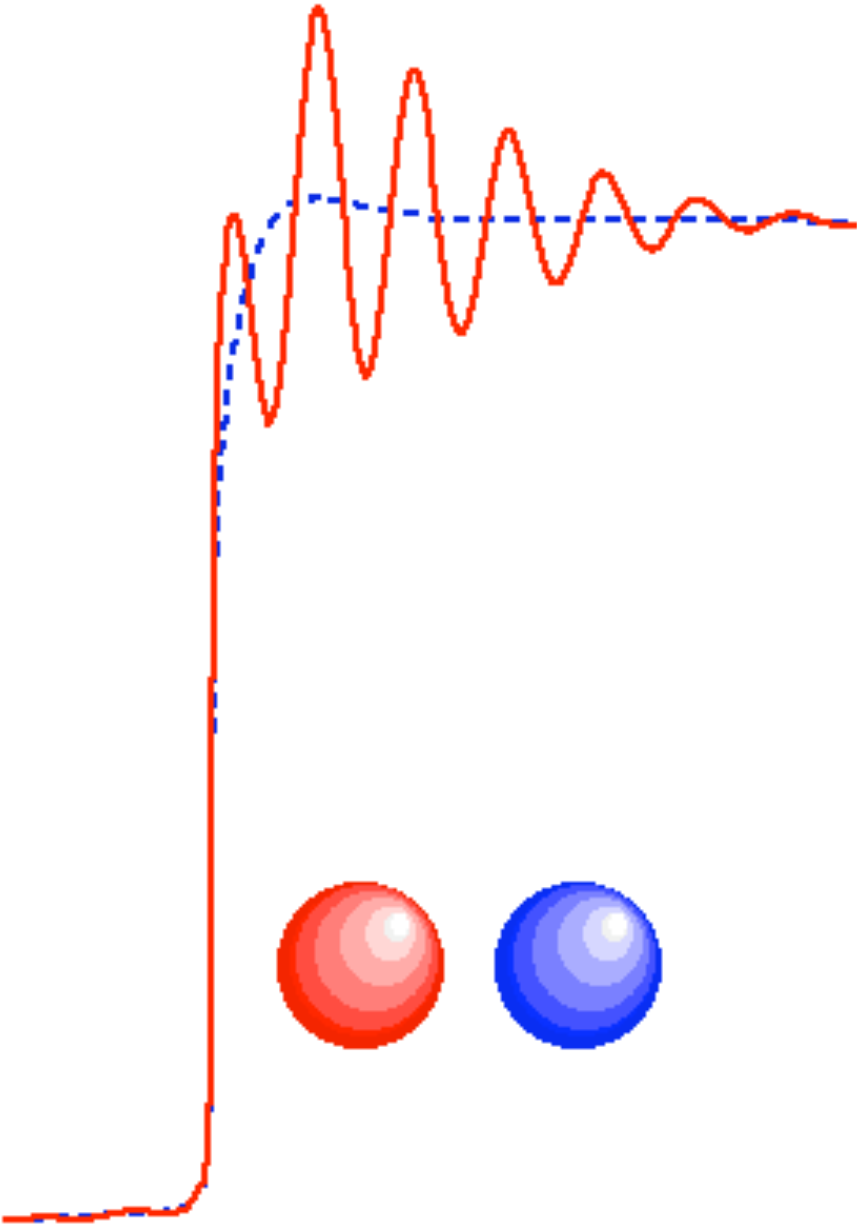
$$\chi = A \cdot \sin(\phi)$$

$A = f(\text{Type}, \text{Number}, \text{disorder} \dots)$

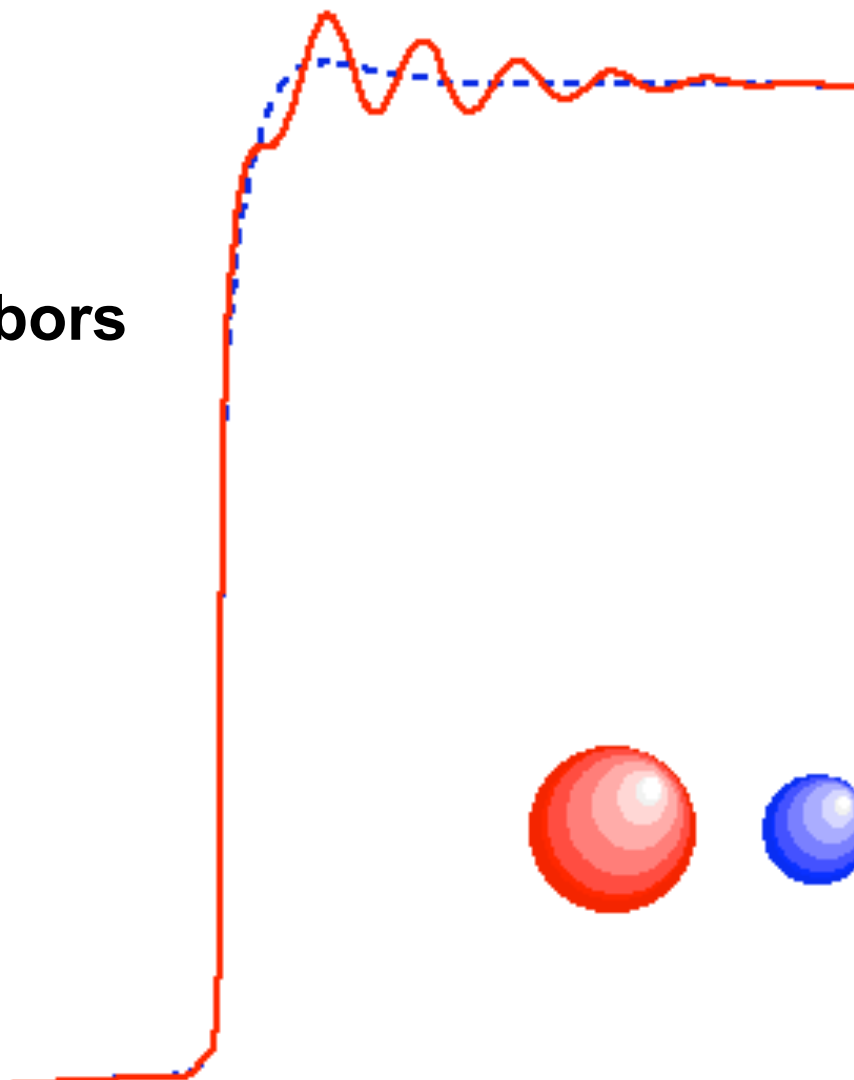
$\phi = f(\text{Type}, \text{distance}, \text{disorder} \dots)$



**effect of  
inter-atomic distance**

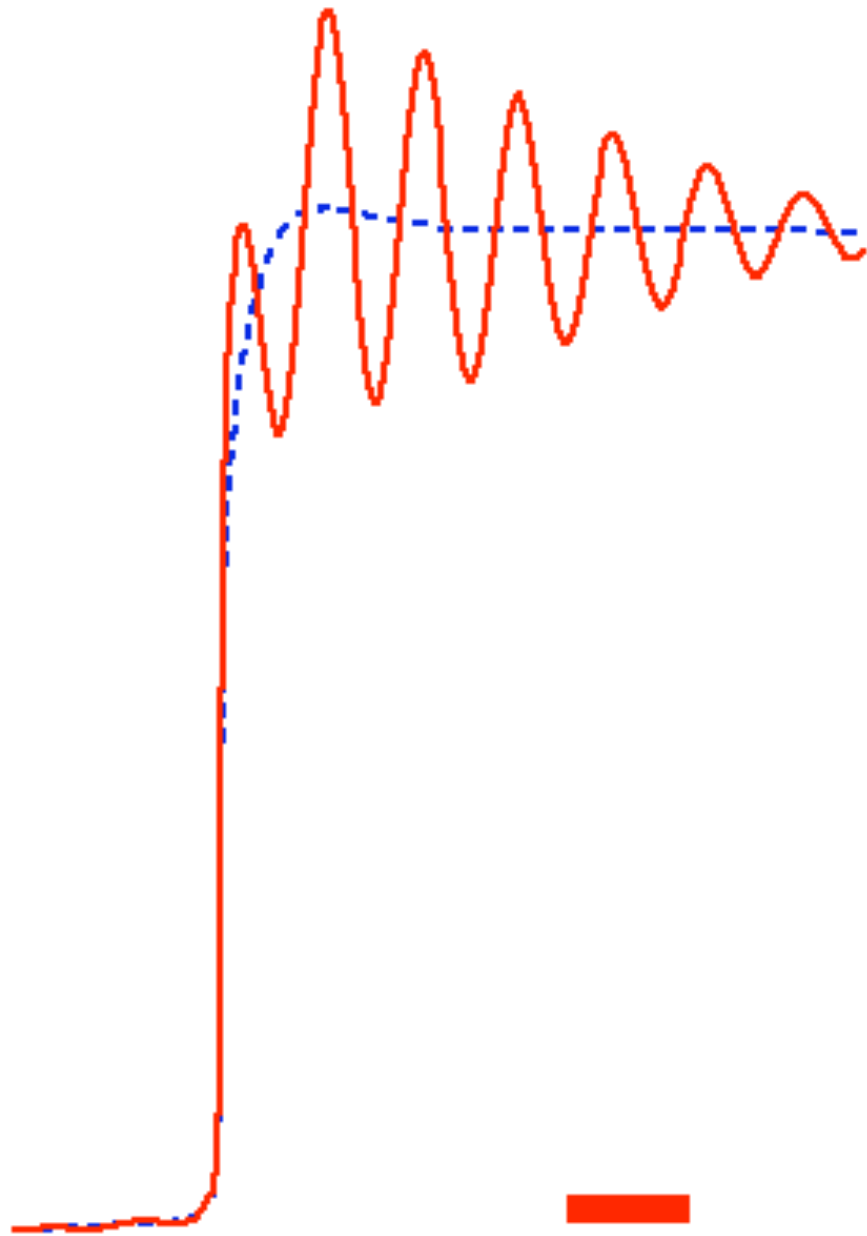


**effect of the number of neighbors  
( and NOT coordination !!)**

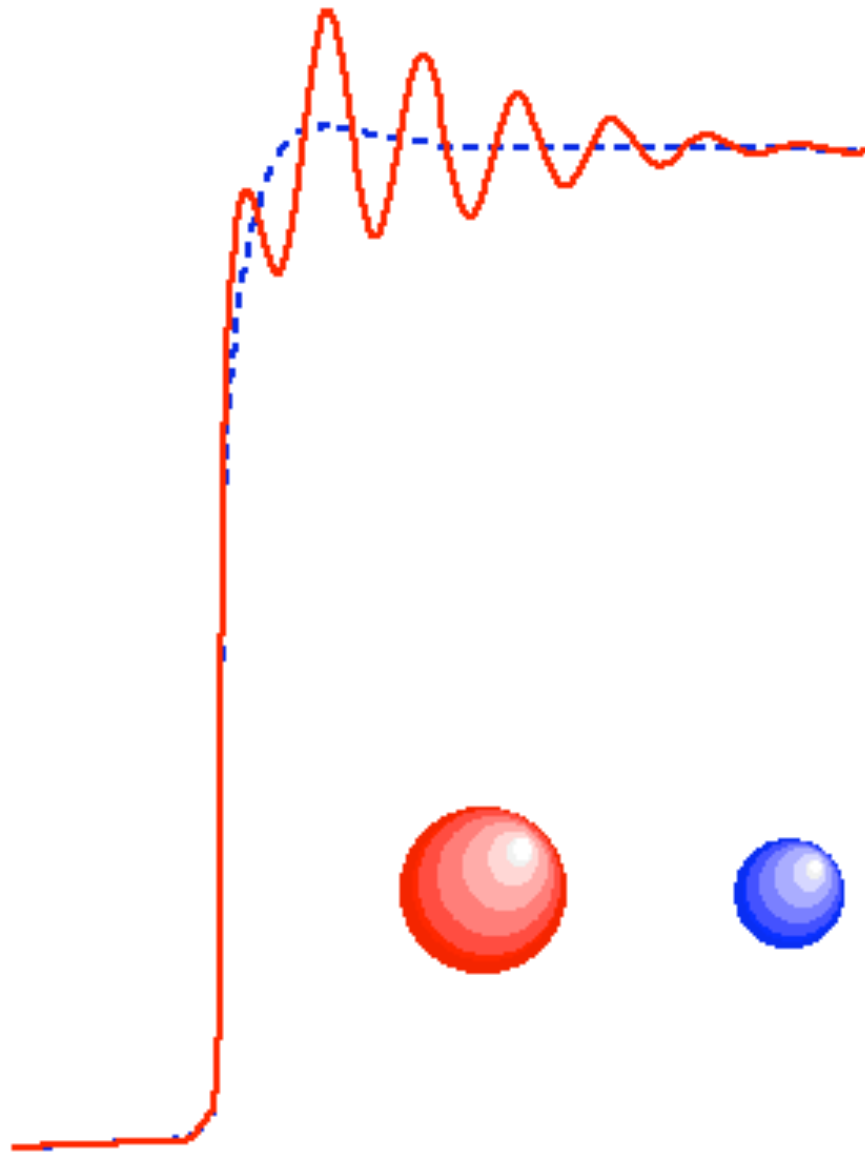


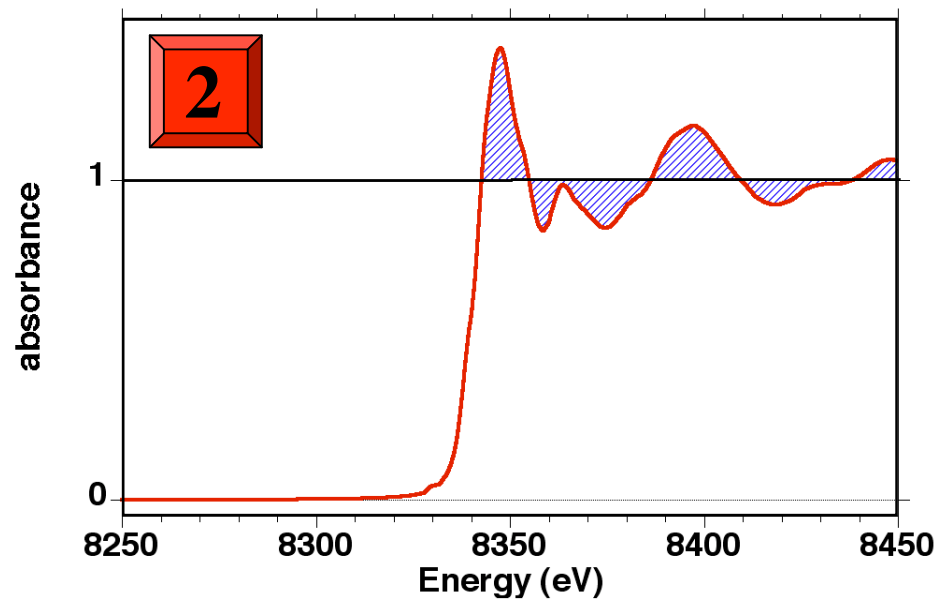
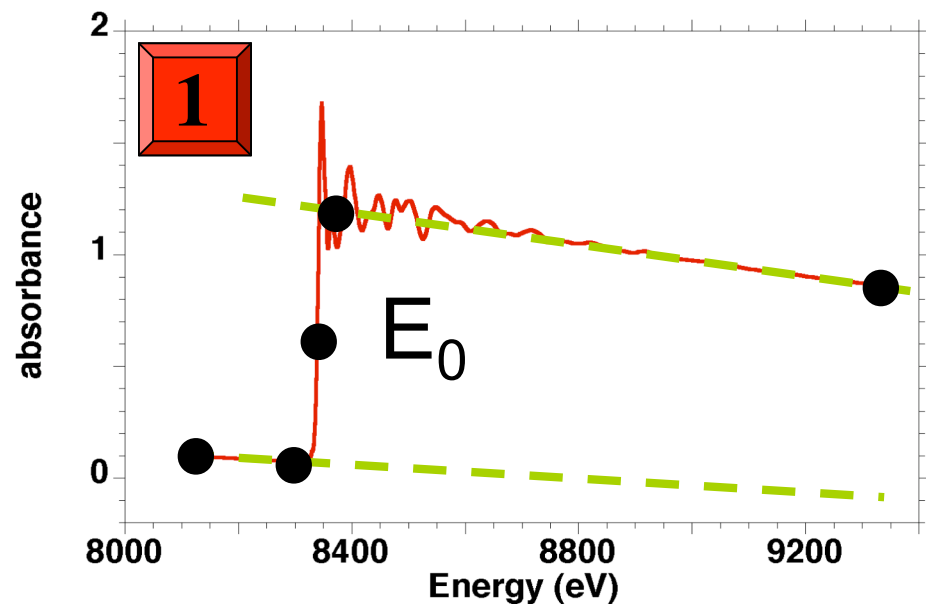
**effect of disorder**

- thermal ( $kT$ )
- static (structural)

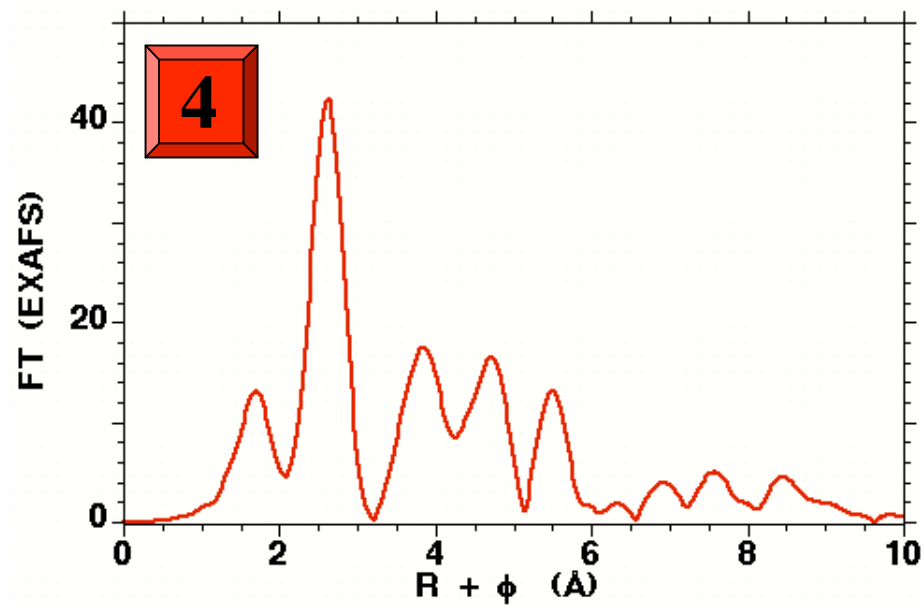
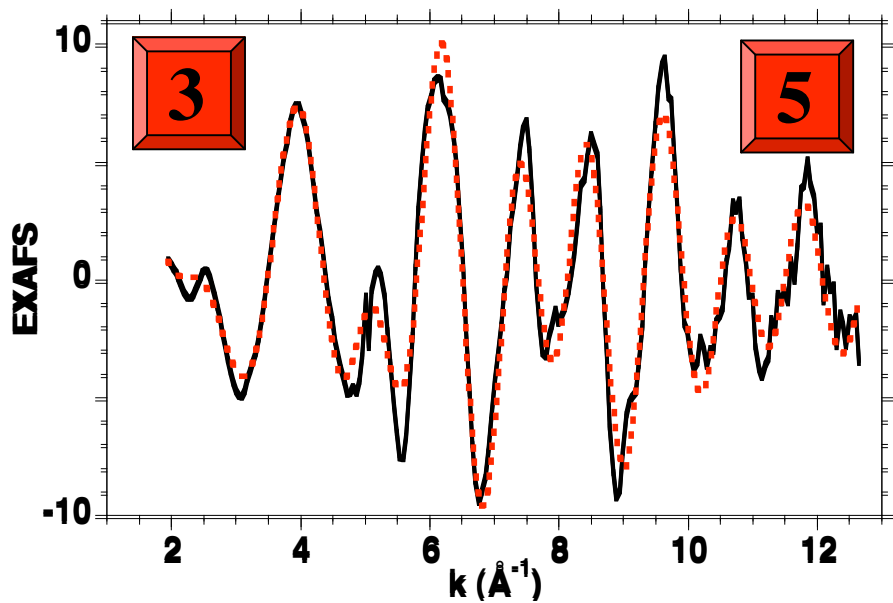


**effect of the  
next-nearest neighbors  
Z-number**



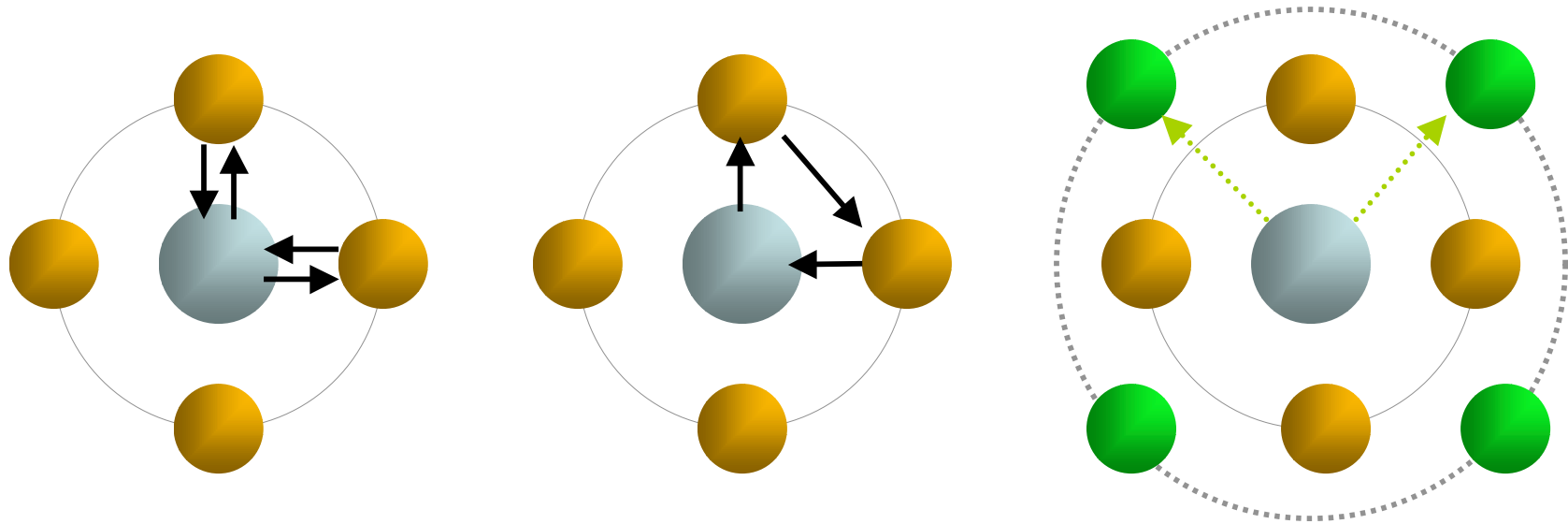


### 5 important steps in EXAFS data reduction



## complex : multiple-scattering (MS) vs. second neighbors (SN)

$$\mu(E) = \mu_{\text{atomic}}(E) \left( 1 + \chi_{\text{SS}}(E) + \chi_{\text{MS}}(E) + \chi_{\text{SN}}(E) \right)$$



- MS is important when atoms are in a special 2D geometry  
2D (colinearity, triangles, square planar...)
- symmetry should be estimated thanks to other methods  
(pre-edge, Raman, MCD, Vis-UV-IR)
- and/or *XAFS ab-initio calculations (FEFF, gnxas etc)*

# More complex : disorder effect ....



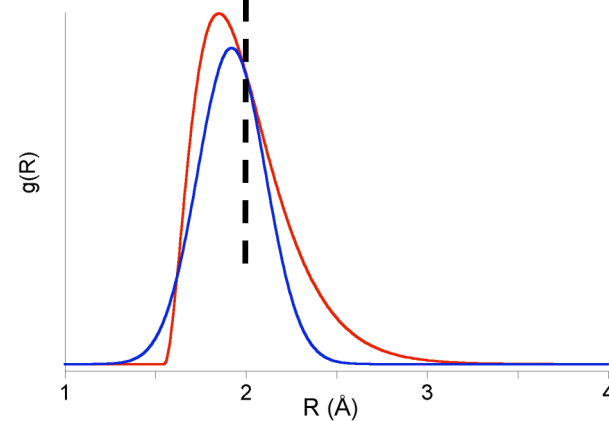
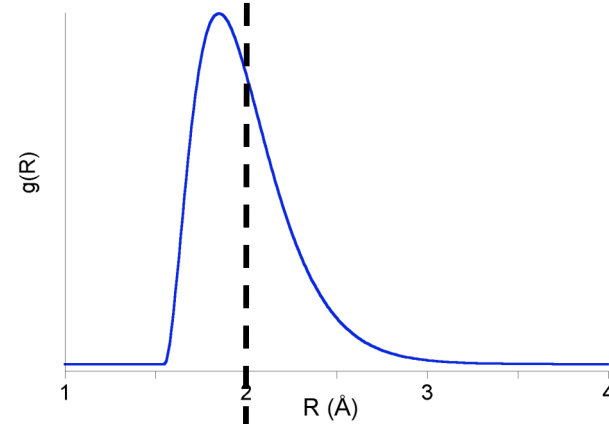
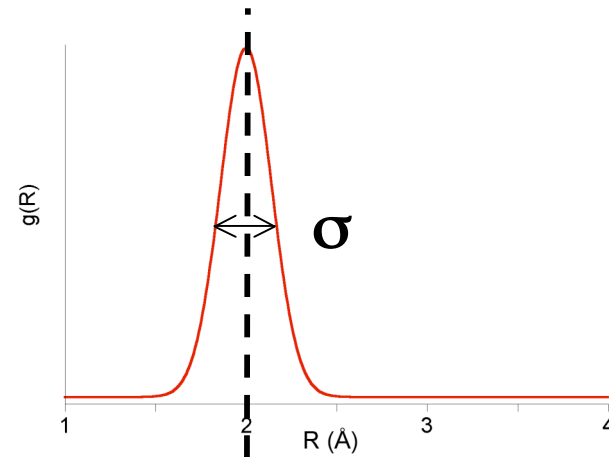
low temperature  
and/or covalent bonding



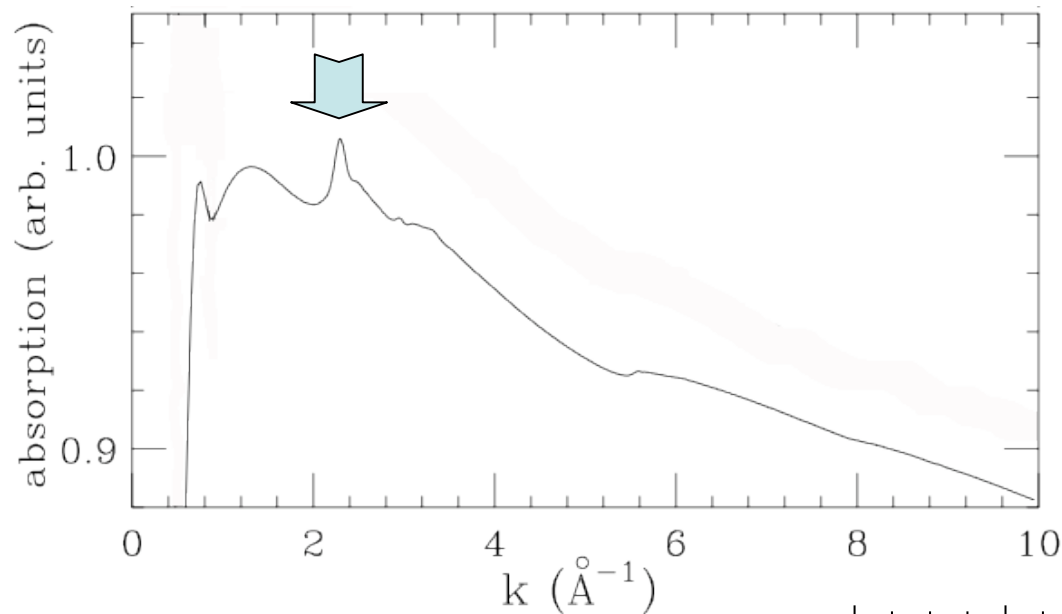
high temperature  
and/or ionic bonding



and when it shakes too much....



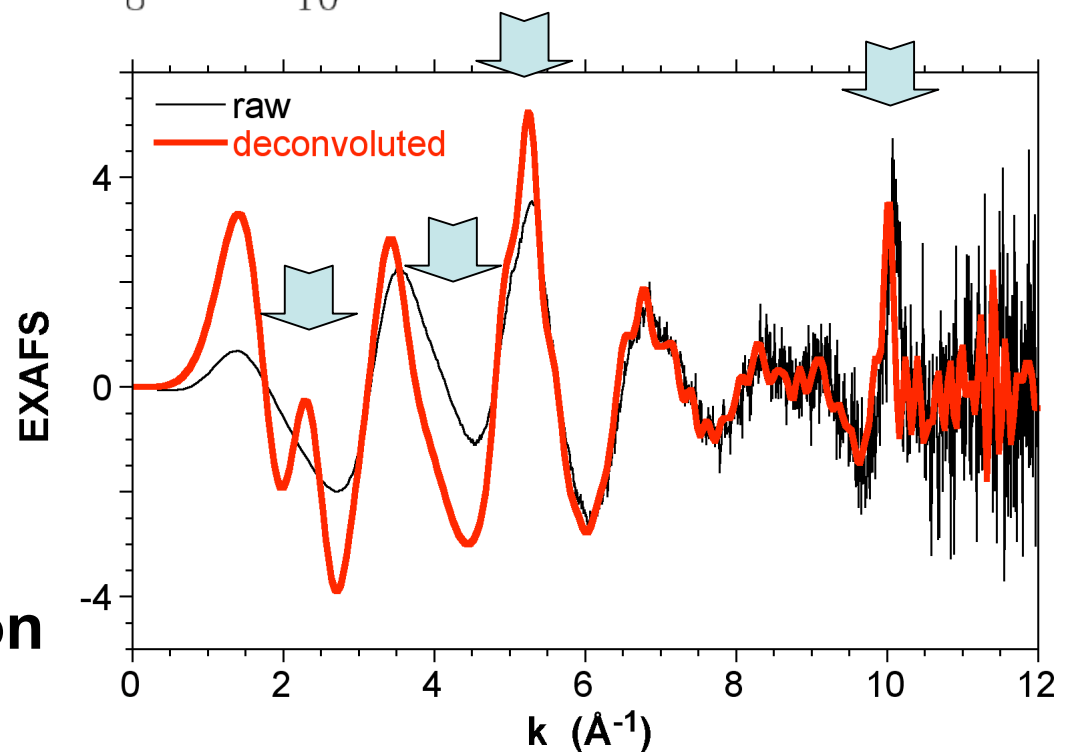
## complex : multi-electronic edges (MEE) :



krypton (gaz)  
au seuil K de Kr

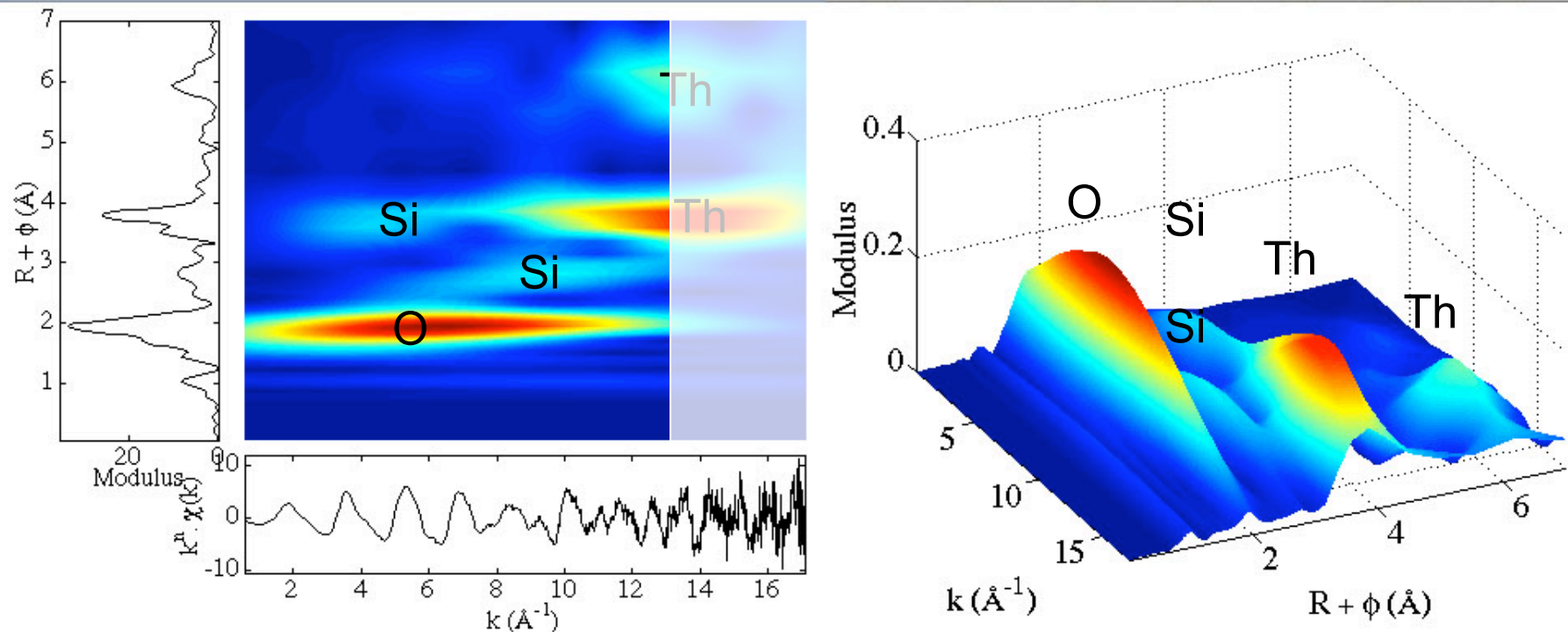
(after Fillipponi, 2000; Ito et al., 1992)

Th in  $\text{ZrSiO}_4$   
At the  $L_{III}$  of Th  
(after Farges et al., 2002)



- difficult to detect
- affects XAFS normalisation

- a new EZ tool to help : wavelets decompositions  
Th in  $\text{ThSiO}_4$  :



- **Si near 3 et 4  $\text{\AA}$  are detected until  $\sim 6 \text{ \AA}^{-1}$**
- **destructive interf. between Si and Th near ( $k = 8, R = 3.9$ )**
- **if data is collected to  $k = 13 \text{ \AA}^{-1}$  : not correct for Th !!**

Equation that we model (EXAFS) :



$$\langle \chi(k, R) \rangle = \int_0^{\infty} g(R) \chi(k, R) dR \quad \text{with :}$$

$$\Rightarrow \chi(k, R) = \left[ S_0^2 R f |F_{cw}| \right] \frac{N}{k R^2} e^{-2R/\lambda} \sin(2kR + \sum \phi)$$

$$\Rightarrow g(R) = \frac{U(R)}{k_B T} ; U(R) = U(R)_{\text{harmonic}} + U(R)_{\text{anharmonic}}$$

$$\int_0^{\infty} \frac{U(R)}{k_B T} dR$$

with  $S_0^2$  is the amplitude reduction « many body » factor);  $Rf$  is the amplitude reduction for the central atom loss;  $N$  is the number of backscattering neighbors ré;  $|F_{cw}|$  is the effective curve-wave amplitude of the neighbors;  $\lambda$  is the mean free path of the photoelectron AND  $\sum\phi$  is the sum of the phase-shifts (central and rbackscattering ).

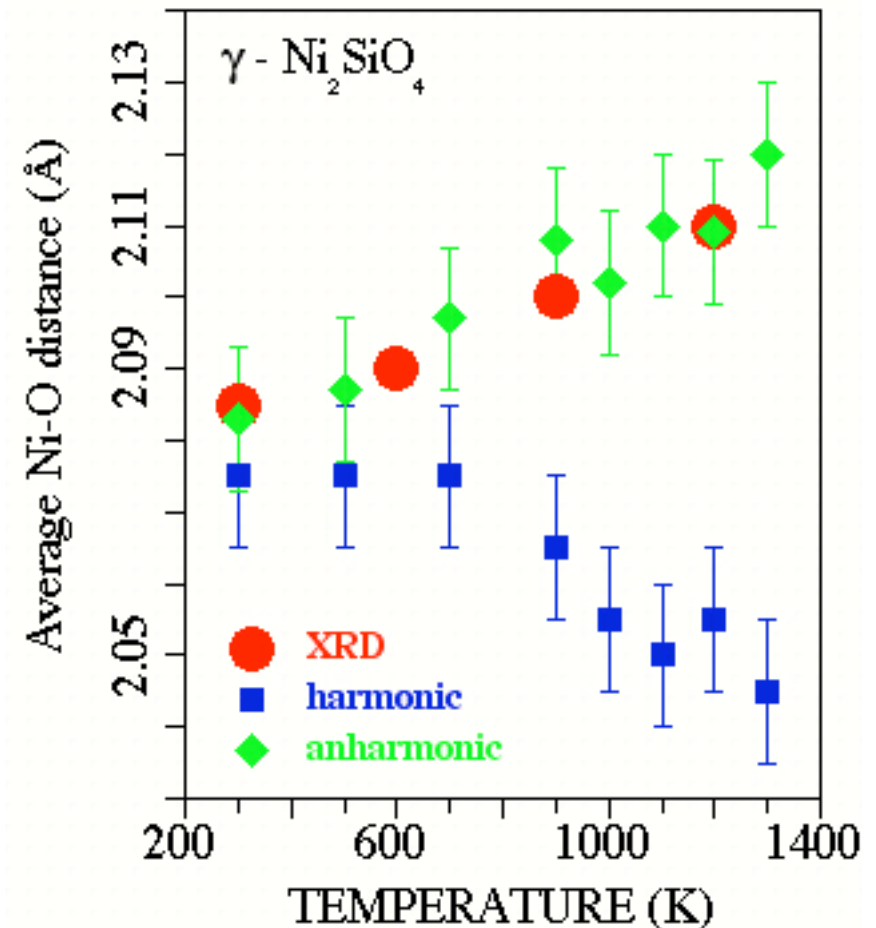
- if the studied dipole is an harmonic oscillator :

$$\Rightarrow \chi(k) = \left[ S_0^2 R_f |F_{cw}| \right] \frac{N}{k R_0^2} e^{-\left[ 2R_0/\lambda + 2k^2\sigma^2 \right]} \sin \left( 2kR_0 + \sum \phi \right)$$

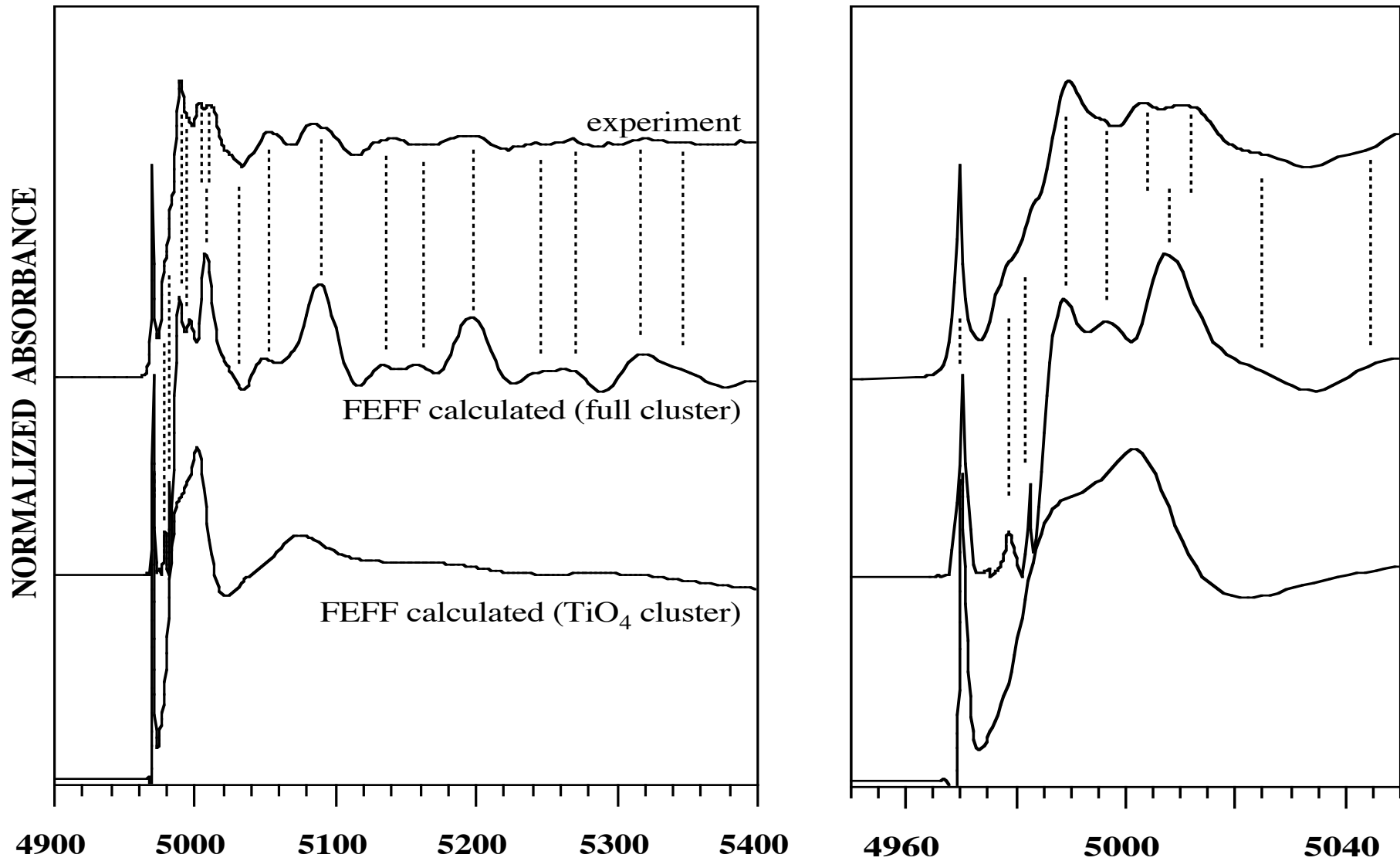
- otherwise, anharmonicity of the potential must be considered :














- cumulants
- asymmetric  $g(r)$
- asymmetric potential (incl. double-well etc.)

- otherwise, data reduction is not reliable :



- possibility to calculate rather easily all spectra :



software	pros	cons
<b>IFEFFIT (+FEFF)</b>   	powerful, modular, many updates, consider MS, anharmonicity, support	Tougher for beginners or occasional users
<b>WinXAS</b> 	interface, do dispersive EXAFS, many fonctions, support	costs \$100+, Windows only, some missing functions
<b>XAFS</b>   	intuitive, powerful, run on Win via emulation(*), does also RMC fits	can go just too fast easily think the software is bad
<b>SamPack</b> 	Easy interface, many modern methods (PCA, self-abs.)	Need a huge screen, portable to other platforms uneasy
<b>GNXAS</b> 	powerful, modular, uses state of the art MS, anharmonicity and MEE	UNIX only, support and documentation needs better
<b>Excurv98</b> 	powerful, modular, uses MS, good anharmonicity evaluations	commercial, UNIX, poor interface (to be better soon)
<b>EXAFSPAK</b>  	modular, robust models, runs on Win via Hummingbird	fairly VMS dedicated, missing anharmonicity
<b>XANDA - VIPER</b> 	powerful, numerous very unique possibilities, free	Stability recently improved, no documentation

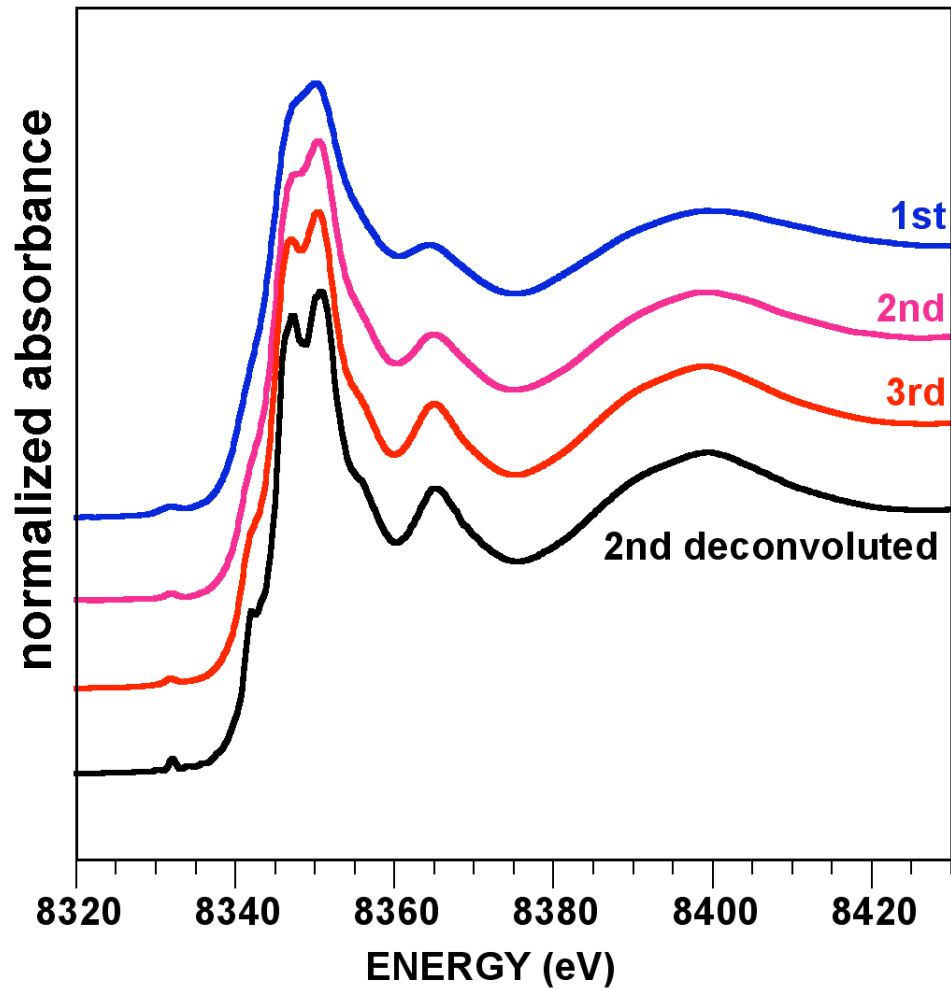
**Not or poorly tested : DARESURY, LASE, CDXAS, XFIT, SEDEM, EDA and many others. These opinions just reflect the author's use, sometimes in ratehr quick conditions.**

## **“COMMANDEMENTS” OF AN OK-XAFS USER :**

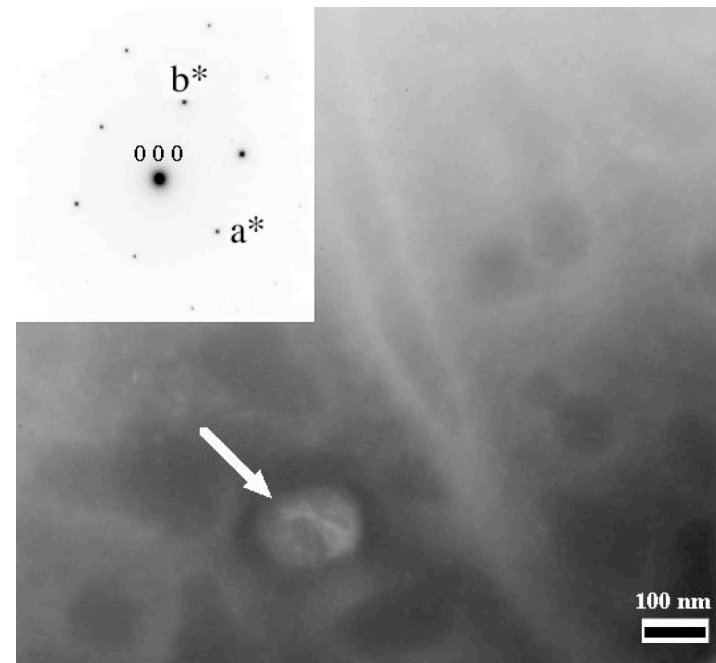
- 1) Physics always works, we sometimes don't**
- 2) Make up your mind first and double-check everything being told or “recommanded”. Track any bugs everywhere.**
- 3) Make sure other methods are not as good (Mossbauer, optical spectroscopy, EPR and many others) AND always do complementary work**
- 4) OK to spend as much time/efforts to (1) write proposal, (2) collect data, (3) reduce data or (4) discuss data**
- 5) XAFS methods are usually among the very last of a project AND not the first to start with (as microscopic, elemental, isotopic methods)**
- 6) Don't choose Win software because you can only run \*.exe's : the choice of a software is first and then you find the right computer afterwards. Otherwise, UNIX is, then, THE safest choice.**

***Failure(s) in any of these points might led to unreliable research***

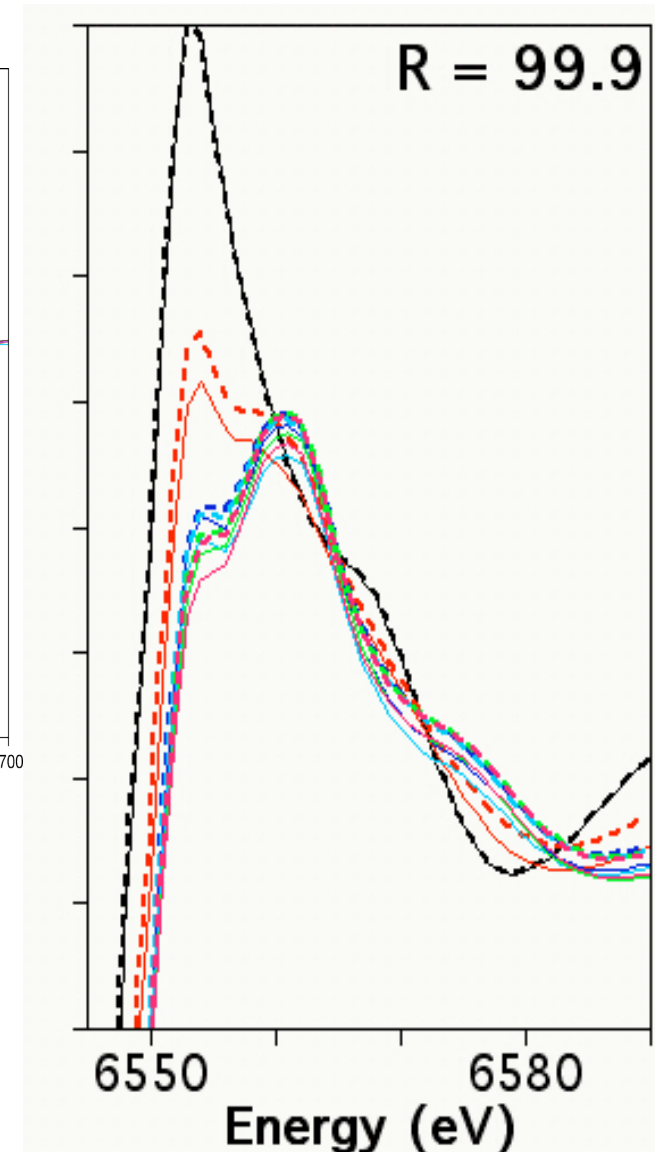
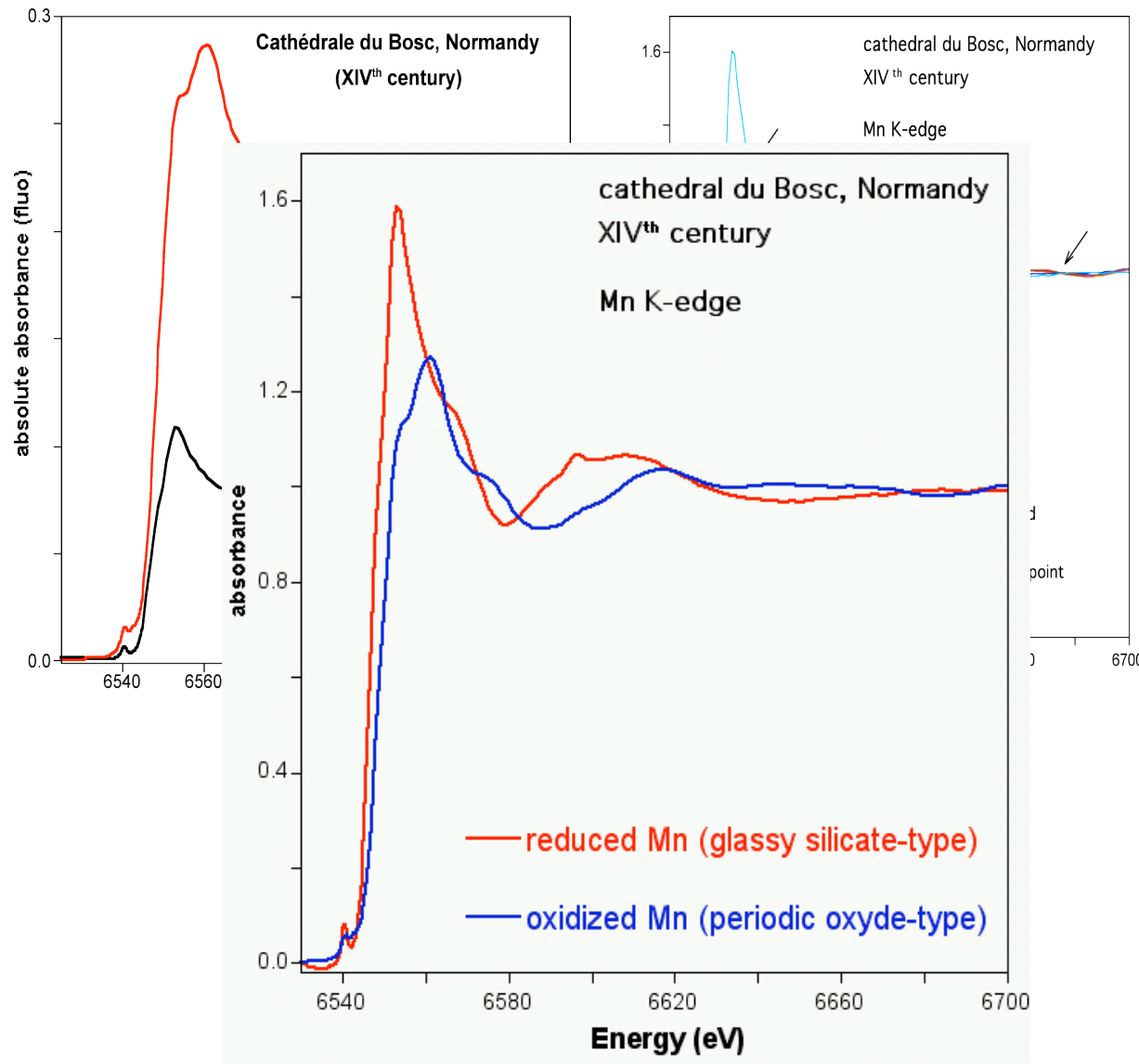
## Example 1 : deconvolutions



- important effect of experimental resolution on the spectra
- deconvolution evidences nanocrystalline domains, that TEM can evidence too :

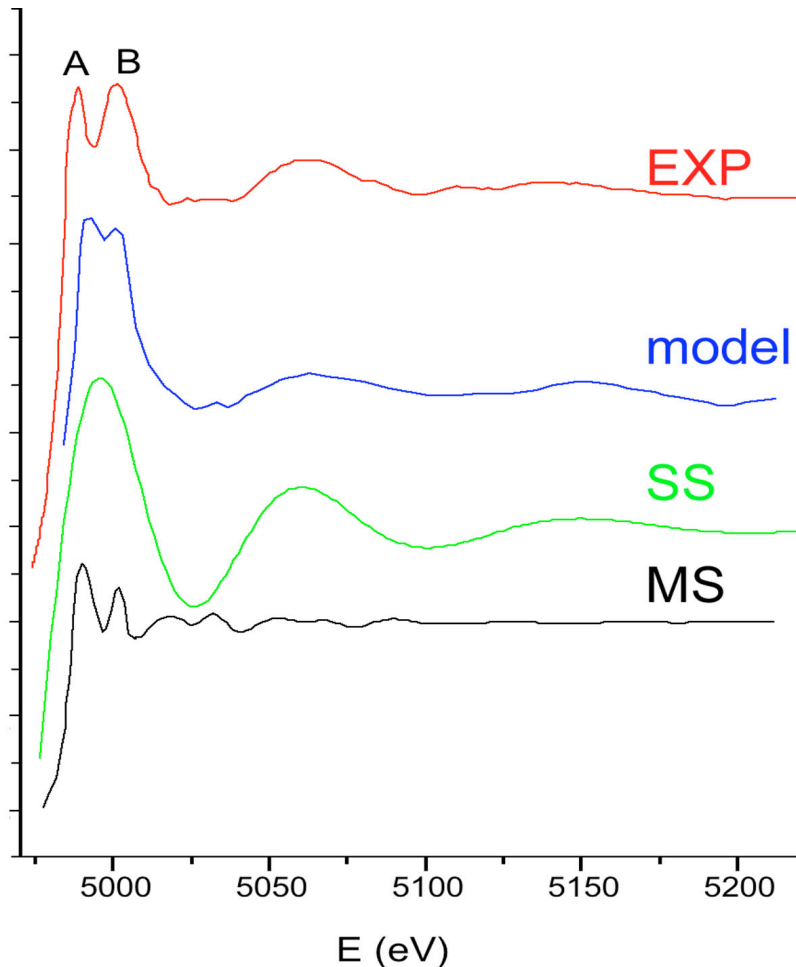


# Manganese in a stained glass (XIVth century, France) fresh versus weathered sides : a PCA analysis



### Exemple 3 : radiation damage in titanite ( $\text{CaTiSiO}_5$ )

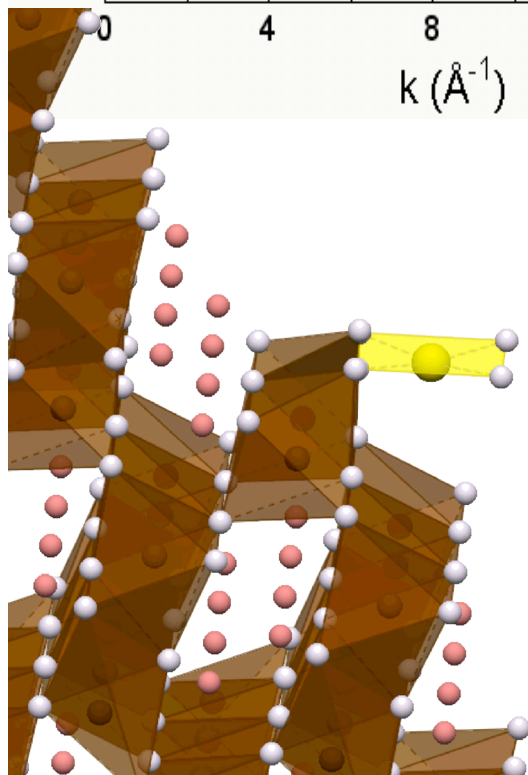
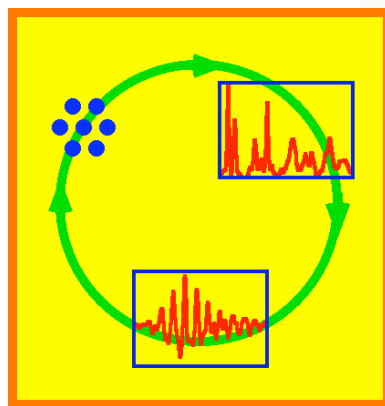
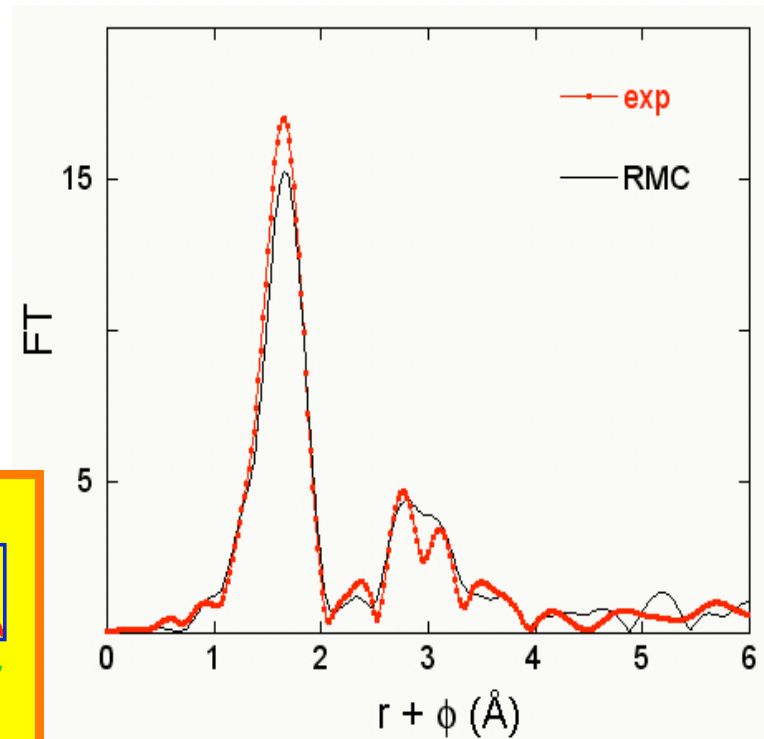
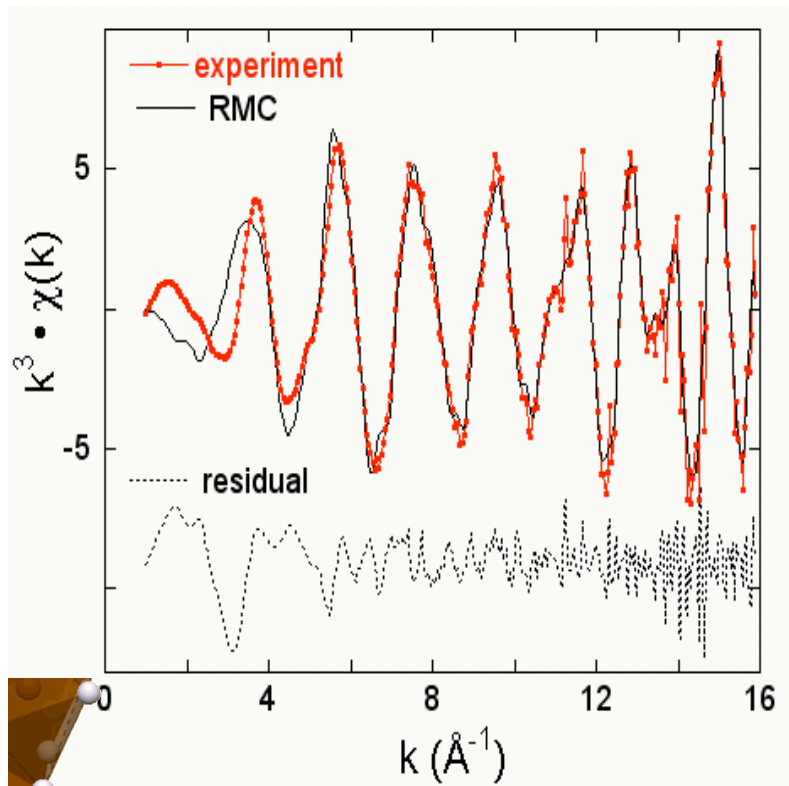
XANES of Ti in various titanites (crystalline and amorphous) can be modeled (CAUTIOUSLY) as EXAFS :



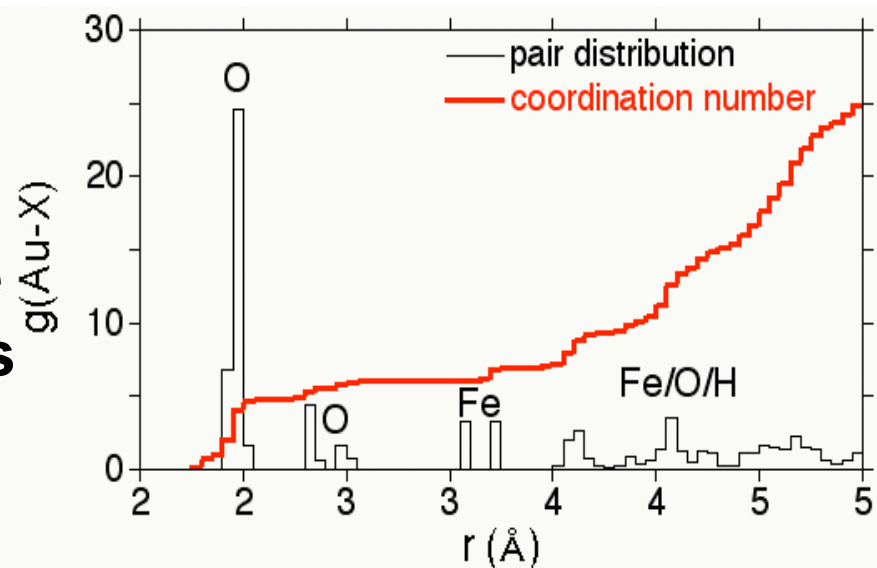
**SELFCOMP models**  
(Bugaev et al., 2002)

samples	N of oxygens	< Ti-O >	R-factor
<b>crystalline</b>	<b>6.1</b>	<b>1.99</b>	<b>0.005</b>
<b>glassy</b>	<b>4.9</b>	<b>1.95</b>	<b>0.001</b>
<b>irradiated +</b>	<b>5.7</b>	<b>1.98</b>	<b>0.005</b>
<b>irradiated ++</b>	<b>5.5</b>	<b>1.97</b>	<b>0.002</b>

**“quantitative” XANES**  
**modeled nearly as EXAFS**



**Reverse Monte-Carlo of the EXAFS coupled to bond valence => sorption sites and EXAFS limitations**



## References :

**There are so much papers that it is impossible to list them. In addition, outside the rather outdated book of Teo (1986), a good « EXAFS for dummies » kind of book (updated and extended to XANES) is really missing.**

### **Remains some « must read first » for the mineralogists/geochemists :**

Crozier E.D., Rehr J.J., and Ingalls R. (1988) Amorphous and liquid systems. In *X-ray absorption. Principles, applications, techniques of EXAFS, SEXAFS and XANES* (ed. D.C. Koningsberger and R. Prins) pp. 373-442, Chemical Analysis, vol. **92**, John Wiley, New-York.

Teo B.K. (1986) *EXAFS: basic principles and data analysis (Inorganic Chemistry Concepts, Vol. 9)*. Springer-Verlag, New York, 350 p.

Brown G.E., JR., Calas G., Waychunas G.A. and Petiau J. (1988) X-ray absorption spectroscopy: Applications in mineralogy and geochemistry. In *Spectroscopic Methods in Mineralogy and Geology* (ed. F.C. Hawthorne), pp. 431-512, *Reviews In Mineralogy*, Vol. 18, Mineralogical Society of America, Washington, DC.

Brown G.E., Jr., Farges F. and Calas G. (1995) X-ray scattering and x-ray spectroscopy studies of silicate melts. In *Structure, dynamics and properties of silicate melts* (ed. J.F. Stebbins, P.F. McMillan and D.B. Dingwell), pp. 317-410, *Reviews In Mineralogy*, Vol. 32, Mineralogical Society of America, Washington, DC.