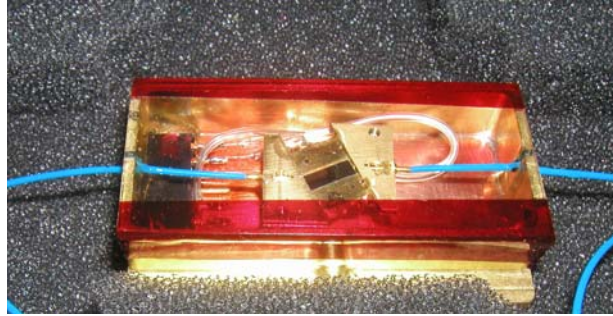


# Some Ideas for Photonic Approaches to LCLS timing, jitter, and xray temporal history measurements



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Presented to XFEL 2004

7/28/04

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# A system for fast single-transient radiation measurements

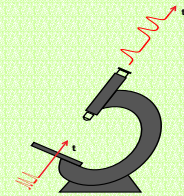


Radiation to optical converter



250 fs Resolution

Temporal Imaging



CW LASER

1550 nm

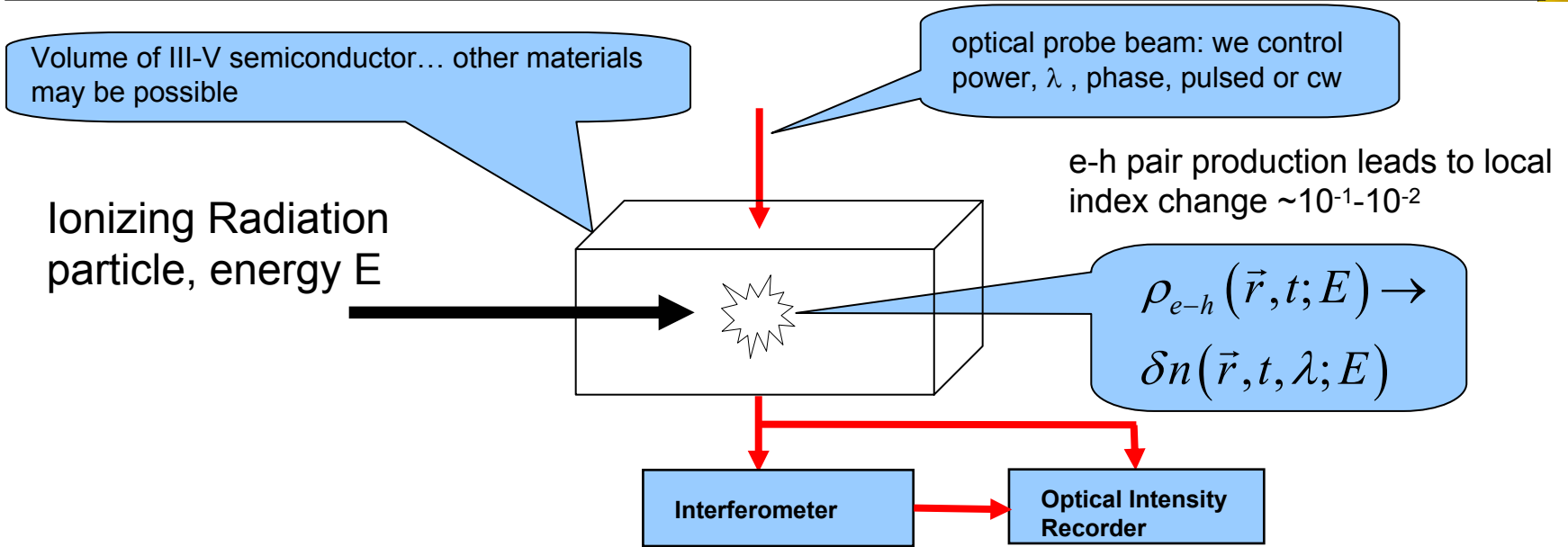
Upconverting Temporal Microscope

Stretched signal at 775 nm

Optical Streak Camera

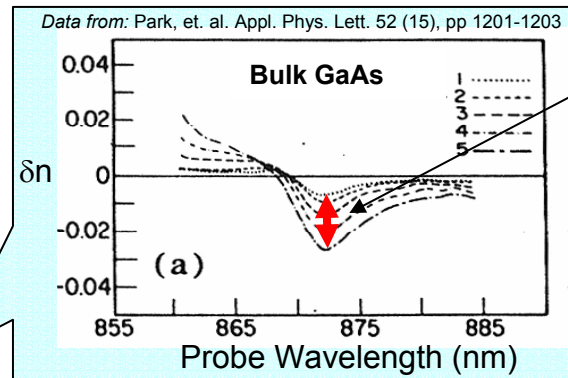
Single shot x-ray recorder for complex arbitrary waveforms. Sub-ps resolution and greater than THz instantaneous bandwidth.

# Radoptic Effect Radiation Detection



## •Why III-V semiconductors?

- Huge base of research in all optical switching for telecom applications
- established device technology
- index change depends on  $\rho_{e-h}$
- ~100 fs temporal response
- typical all-optical switching results



Index change increases with  $\rho_{e-h}$

$\rho_{e-h}$  and  $\lambda$  dependence

$$\delta n_{opt}(\lambda) = C \left[ \frac{\rho_{e-h} / \rho_{sat}}{1 + \rho_{e-h} / \rho_{sat}} \right] \left[ \frac{\Gamma}{1 + \Gamma^2} \right]$$

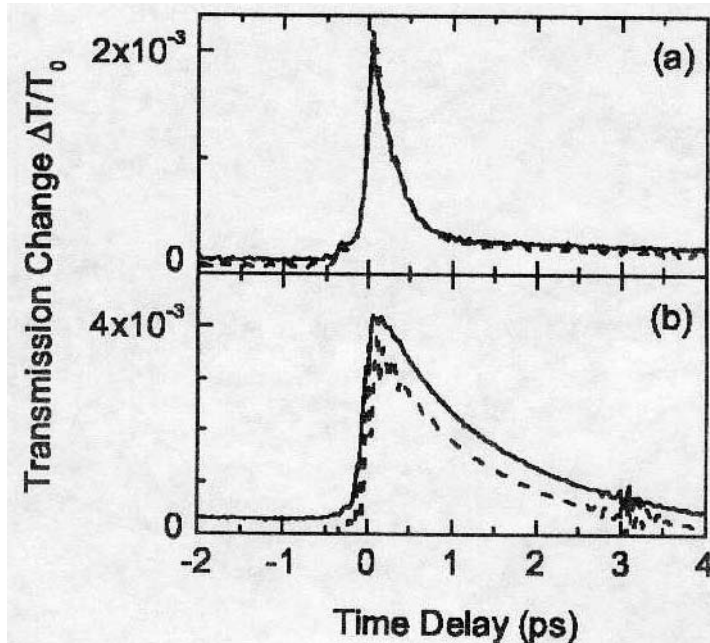
$$= CF(\rho)G(\lambda); \quad \Gamma = \frac{\lambda - \lambda_{edge}}{\delta \lambda}$$

- Ionizing radiation is the analog to the optical pump, index modulation physics the same
- The use of the optical probe is ideal for high-energy radiation particle detectors and high-speed operation: relatively high material volume required (no transport limitations)

# Results from the all-optical switching field show fast response

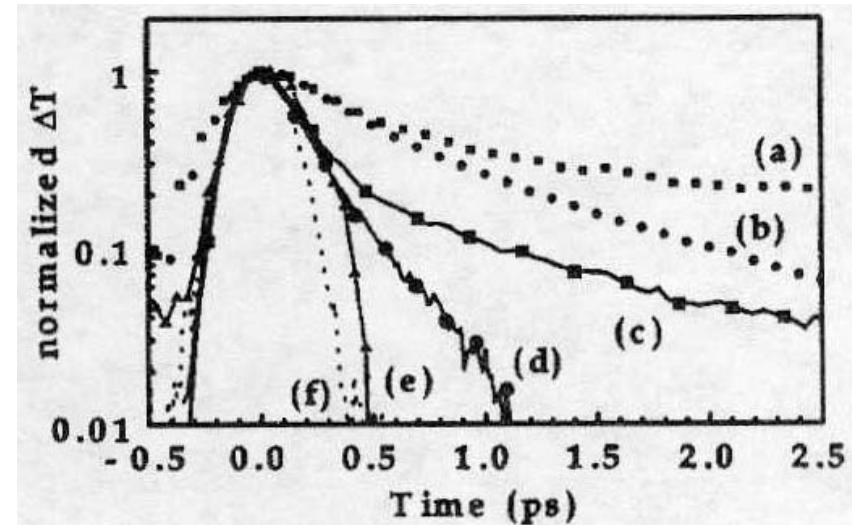


K. Biermann, et al., *Ultrafast optical nonlinearity of low-temperature-grown GaInAs/AlInAs quantum wells at wavelengths around 1.55  $\mu\text{m}$* , Appl. Phys. Lett., 80 (11), pp1936-1938 (2002)



(a) As grown; (b) annealed

T. Okuno, et al., *Femtosecond response time in beryllium-doped low-temperature-grown GaAs/AlAs multiple quantum wells*, Appl. Phys. Lett. 79 (6), pp 764-766 (2001)



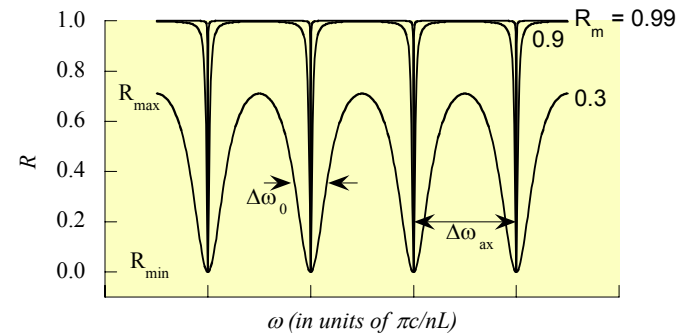
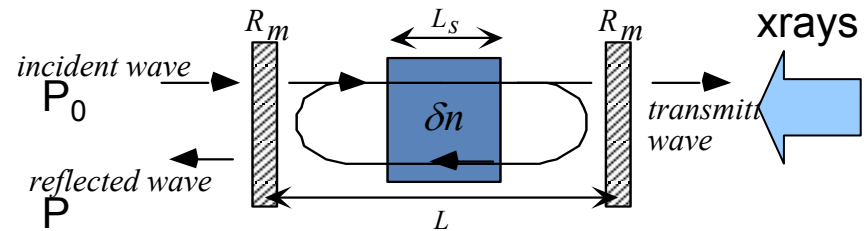
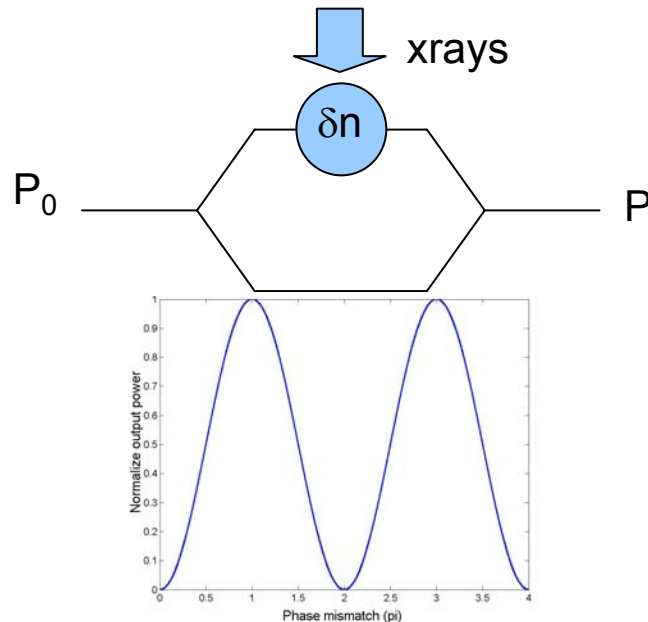
(d) MQW, 1/e fall time = 250 fs

(f) System response

- These devices are the optically-pumped analog of RadSensor.
- We expect similar temporal responses using appropriate epitaxial growth or neutrontdamaged epi

# To probe index change: Interferometry

## Mach-Zehnder and Fabry-Perot compared



$$\frac{\delta p}{P_0} = \pi \frac{L}{\lambda} \langle \delta n \rangle$$

Fringe-fraction

$$\frac{\delta p}{P_0} = 2 \frac{F}{\lambda} \langle \delta n \rangle L$$

$$\tau = \frac{nL}{c} \approx 40 \text{ ps} \quad \text{for } \begin{matrix} L = 3.4 \text{ mm,} \\ n = 3.5 \end{matrix}$$

Time-response

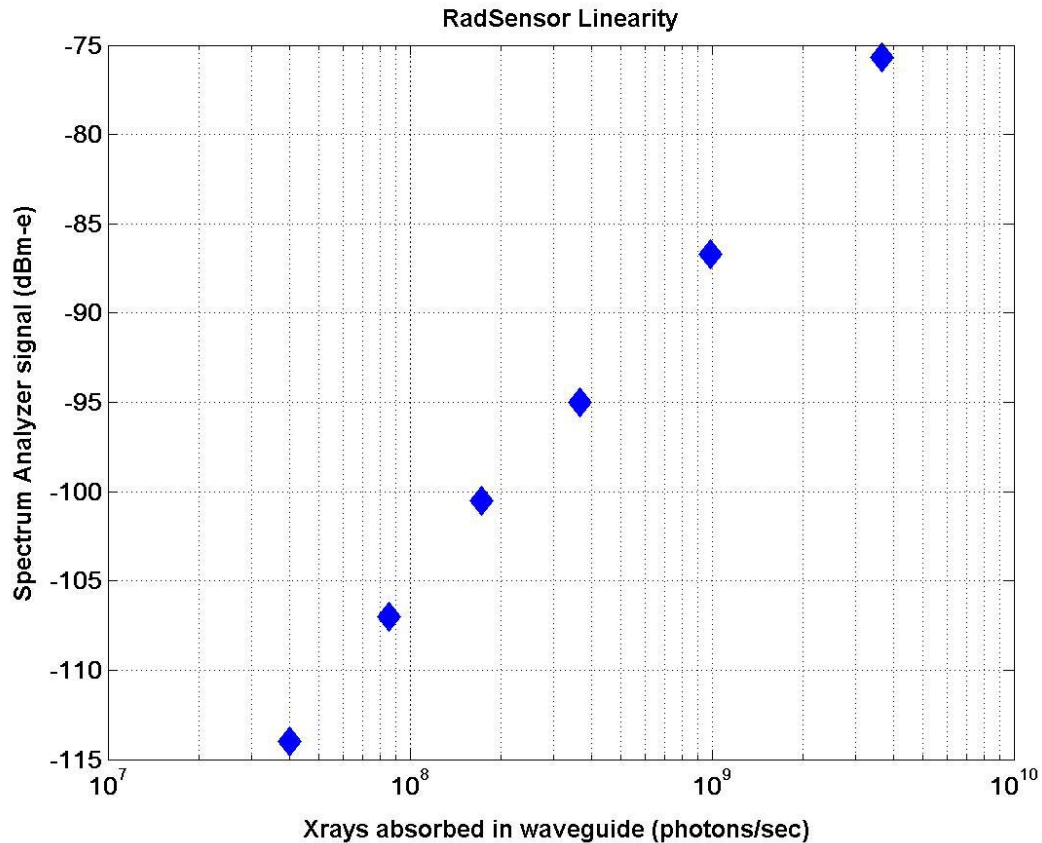
$$\tau_p = \frac{nL}{\pi c} F \cong 0.74 \text{ ps} \quad \text{for } \begin{matrix} F = 10, \\ \lambda = 1550 \text{ nm,} \\ n = 3.5 \\ \text{and } L = 20 \mu\text{m} \end{matrix}$$

- Sensitivity will be determined by how small a fringe-fraction we can measure (1-2% is reasonable); better sensitivity => higher fringe fraction
- The sensitivity of the FP is essentially that of the MZ, multiplied by  $2F/\pi$



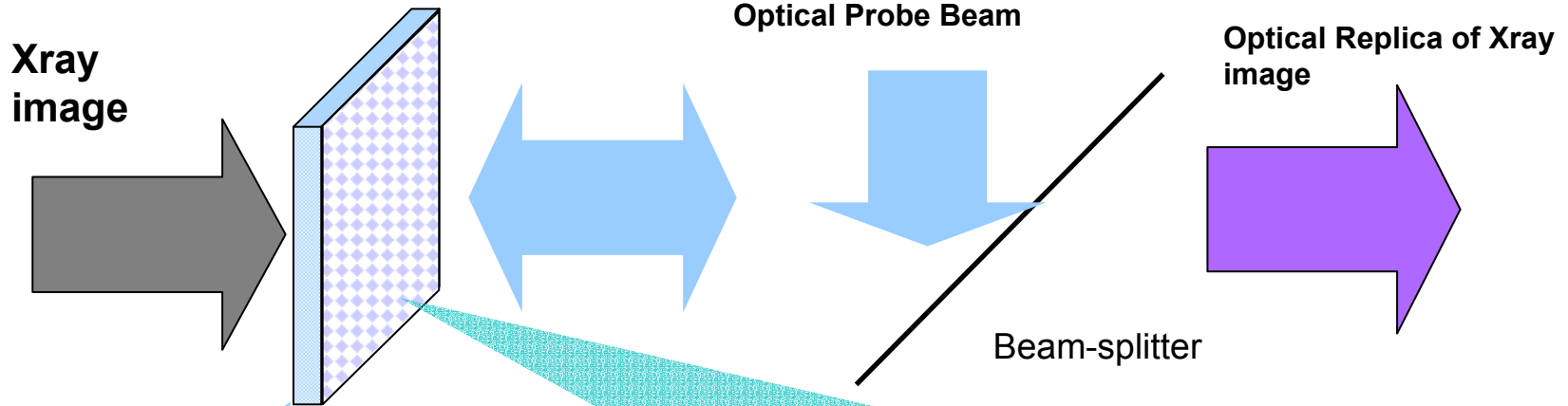


# RadSensor Linearity

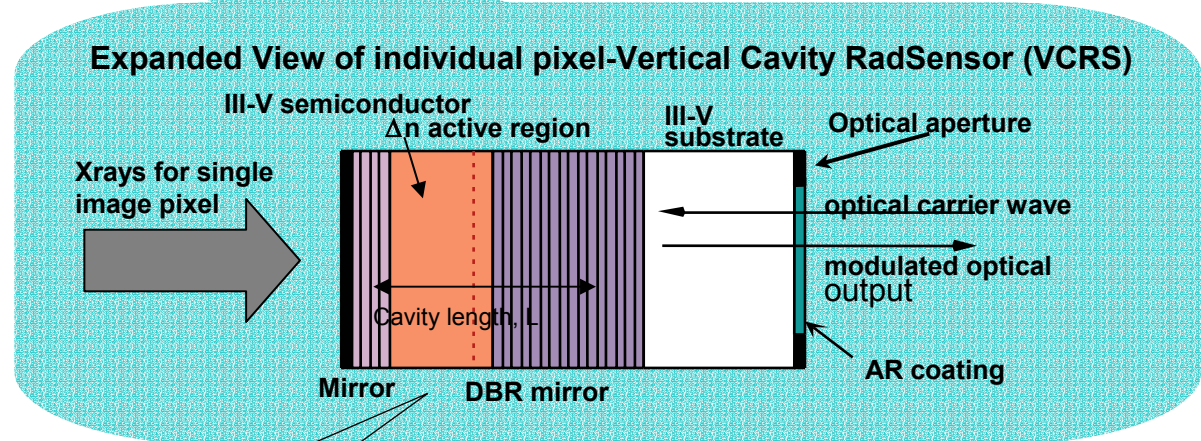


- RadSensor response appears fairly linear over 2 decades
- Note that lower amplitude signals correspond to single xray photon events (9 keV)

# Now We Are Focused on Imager Development in FY04



Monolithic RadSensor Reflection Modulation Array--Fabrication would be at the wafer level using well-established techniques used for VCSELs--  $10^6$  pixels achievable



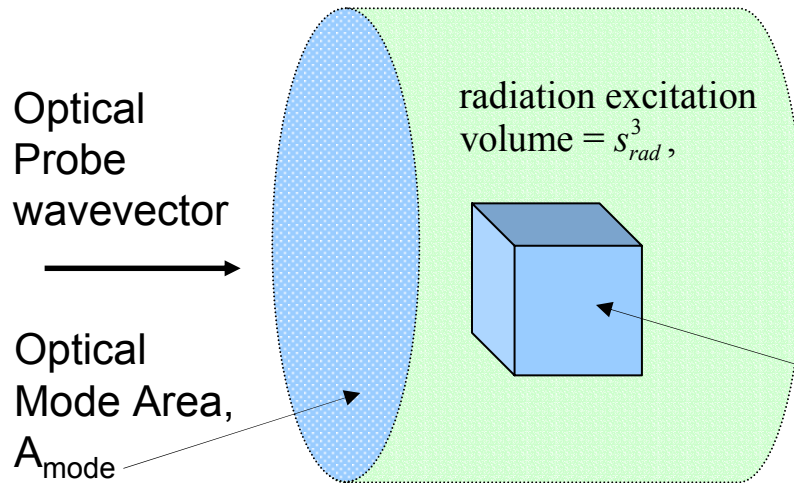
The cavity geometry is not only convenient for imagers it also allows for a sensitivity enhancement  $\sim$  cavity finesse

## R&D Challenges:

- Optimizing sensitivity
- epi for "thick" cavity (eventually)
- epi with short e-h pair lifetime (fast)
- containing scattering (if necessary)



# Optical phase shift from localized radiation excitation



$$\delta\phi_{\text{rad}} = \frac{2\pi s_{\text{rad}}}{\lambda} \langle \delta n \rangle; \quad (\text{Eq. 1})$$

$$\langle \delta n \rangle = \frac{s_{\text{rad}}^2}{A_{\text{mode}}} \delta n; \quad (\text{Eq. 2})$$

$$\delta\phi_{\text{rad}} = \frac{2\pi}{\lambda} \frac{s_{\text{rad}}^3}{A_{\text{mode}}} \delta n; \quad (\text{Eq. 3})$$

$\delta n$  = index change in excitation volume

$$\delta n = CF(\rho_{\text{rad}})G(\lambda) \quad (\text{Eq. 4})$$

**This inhomogeneity will also lead to scattering...**  
being quantified in Kallman's LDRD

From Nonlinear optical theory

$$\delta\phi_{\text{rad}} = \frac{2\pi C}{\lambda} \frac{s_{\text{rad}}^3}{A_{\text{mode}}} F(\rho_{\text{rad}})G(\lambda)$$

If  $\rho_{\text{rad}} \ll \rho_{\text{sat}}$  then,  $F(\rho_{\text{rad}}) \rightarrow \frac{\rho_{\text{rad}}}{\rho_{\text{sat}}}$

$$\rho_{\text{rad}} = \frac{N_{\text{rad}}}{s_{\text{rad}}^3}; N_{\text{rad}} = \frac{E_{\text{rad}}}{E_0};$$

$$E_0 = \frac{15}{5} E_{\text{gap}} + E' \approx 3.15 \text{ eV}$$

(for InGaAsP with  $E_{\text{gap}} = 0.857 \text{ eV}$ )

$$\delta\phi_{\text{rad}} = \frac{2\pi}{\lambda} \frac{C}{\rho_{\text{sat}}} \frac{s_{\text{rad}}^3}{A_{\text{mode}}} \left( \frac{E_{\text{rad}}/E_0}{s_{\text{rad}}^3} \right) G(\lambda)$$

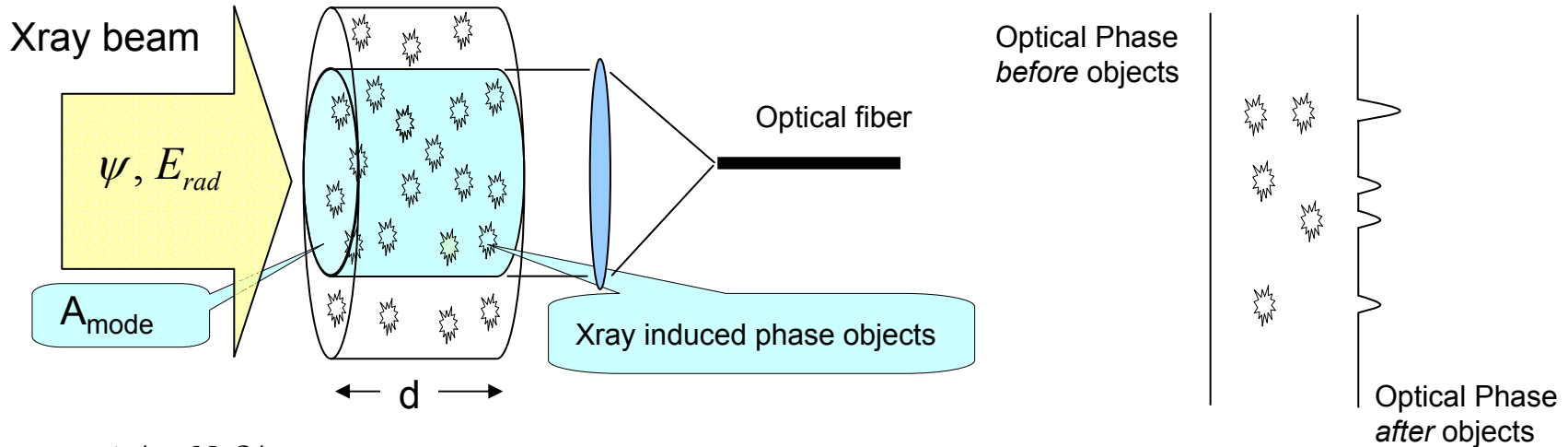
Excitation volume cancels

$$\delta\phi_{\text{rad}} = \frac{2\pi}{\lambda} \frac{C}{\rho_{\text{sat}}} \frac{1}{A_{\text{mode}}} \frac{E_{\text{rad}}}{E_0} G(\lambda)$$

From SSRL data we have (with  $A_{\text{mode}} = 4 \mu\text{m}^2$ ) ...

$$\frac{C}{\rho_{\text{sat}}} = 1.2 \times 10^{-6} \mu\text{m}^3$$

# RadSensor phase modulation is xray irradiance dependent



$$\Delta\phi = N_x \delta\phi;$$

$$\delta\phi = \frac{2\pi}{\lambda} \frac{1}{A_{mode}} \frac{C}{\rho_{sat}} \frac{E_{rad}}{E_0} G(\lambda); \quad \frac{C}{\rho_{sat}} \text{ calibrated from FY03 SSRL data}$$

$$N_x = \# \text{ of xrays absorbed in mode volume} = \psi A_{mode} (1 - \exp(-\mu d))$$

$\psi$  = xray irradiance (# xrays/ $\mu\text{m}^2$ ) into cavity volume

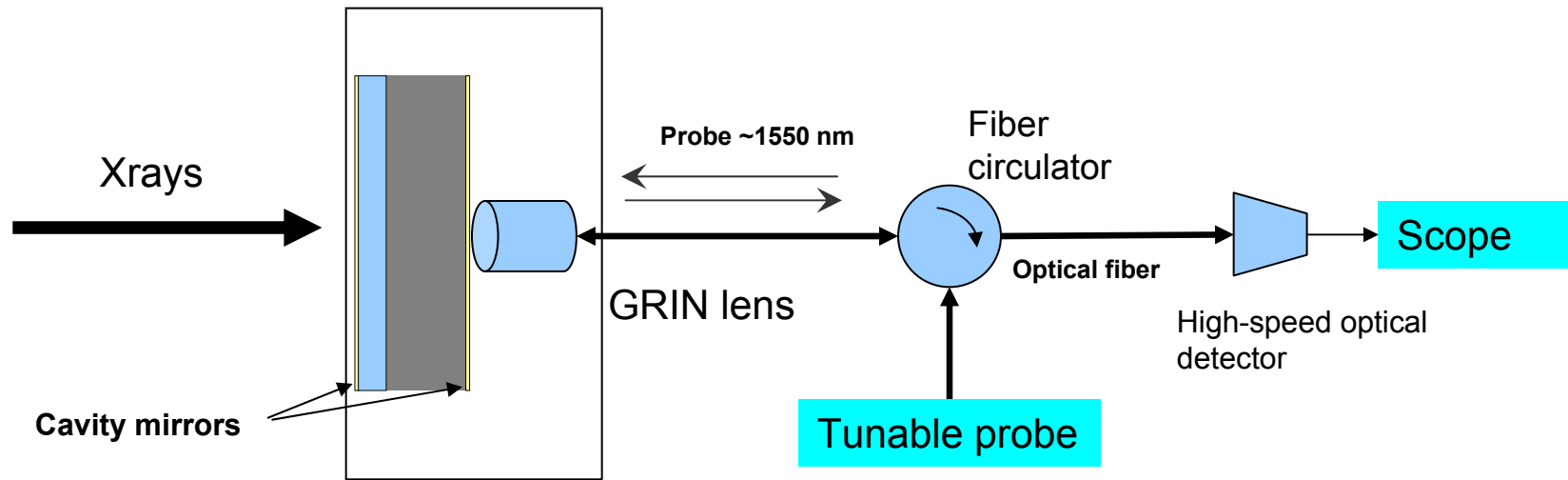
$$\Delta\phi = \psi A_{mode} (1 - \exp(-\mu d)) \frac{2\pi}{\lambda} \frac{1}{A_{mode}} \frac{C}{\rho_{sat}} \frac{E_{rad}}{E_0} G(\lambda)$$

$$\Delta\phi = \psi (1 - \exp(-\mu d)) \frac{2\pi}{\lambda} \frac{C}{\rho_{sat}} \frac{E_{rad}}{E_0} G(\lambda)$$

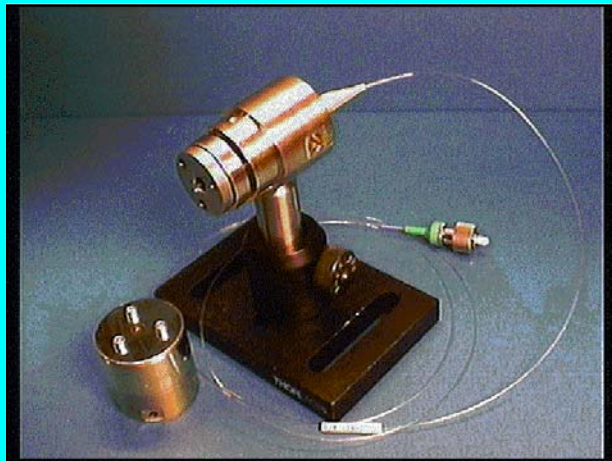
Since the areas cancel, the signal is only dependent on xray irradiance

signal is independent of pixel size... very different from conventional detectors

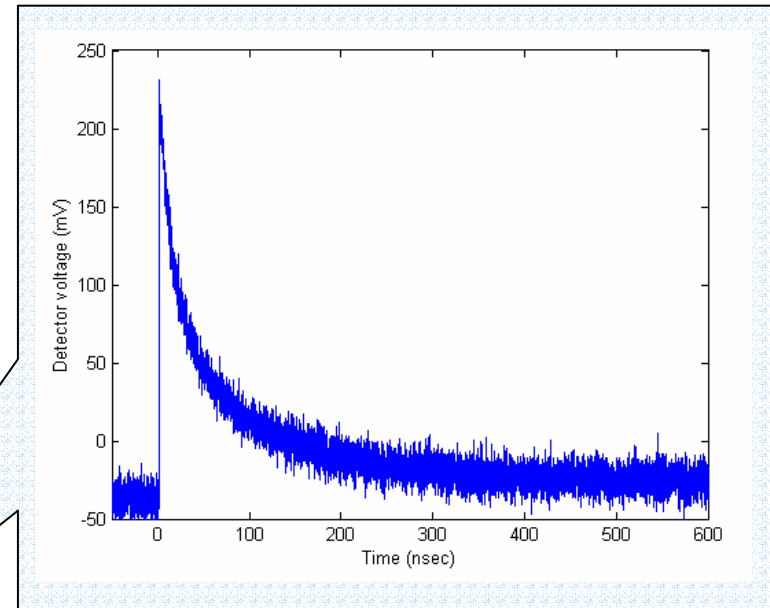
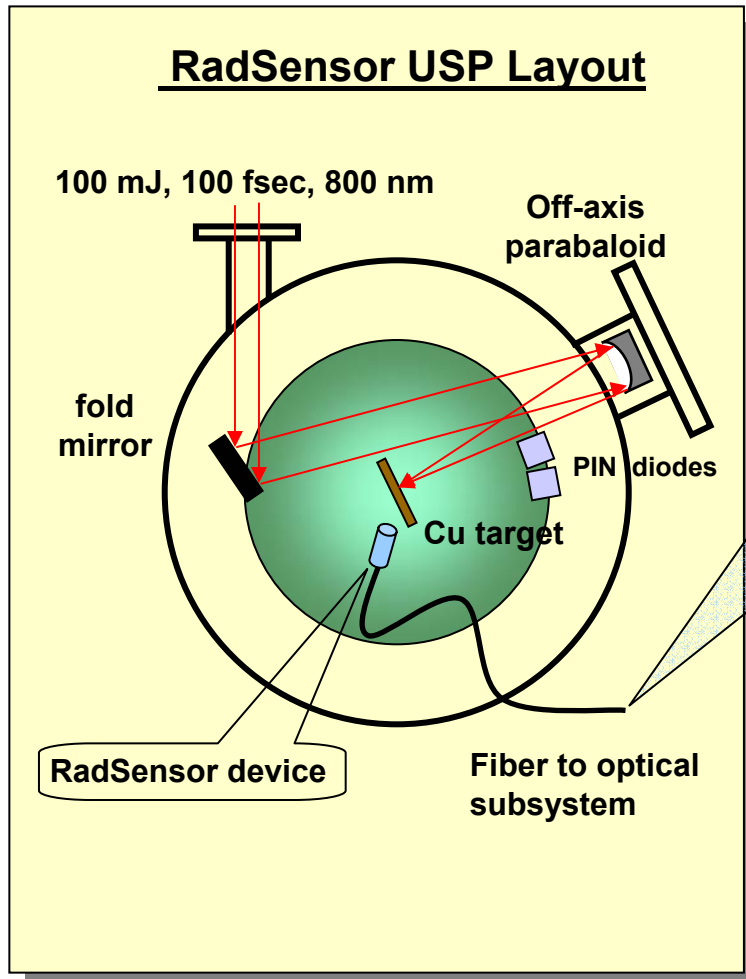
# RadSensor Optical System fielded at USP



Cavity RadSensor Proto Package



# First Single-Transient, Cavity RadSensor Data



- Standard Si Xray PIN diodes were used to monitor the xray output for each shot
- This shot had only Be filter
- Be and Cu filters were used to define and narrow xray spectrum later

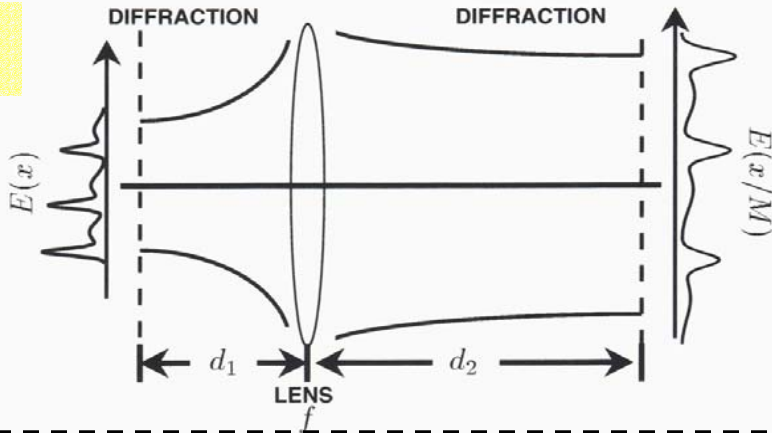
New geometry works with system-limited risetime

# Temporal Imaging Explained by Analogy to Spatial Imaging



Spatial Imaging

a)



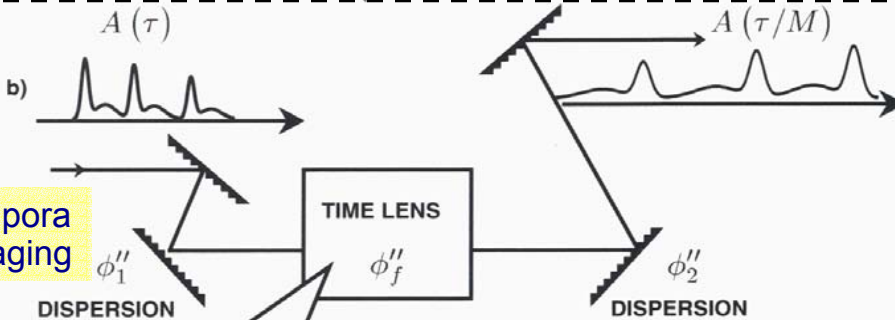
**Imaging Condition**

$$\frac{1}{d_1} + \frac{1}{d_2} = \frac{1}{f}$$

**Magnification**

$$M = -\frac{d_2}{d_1}$$

Temporal Imaging



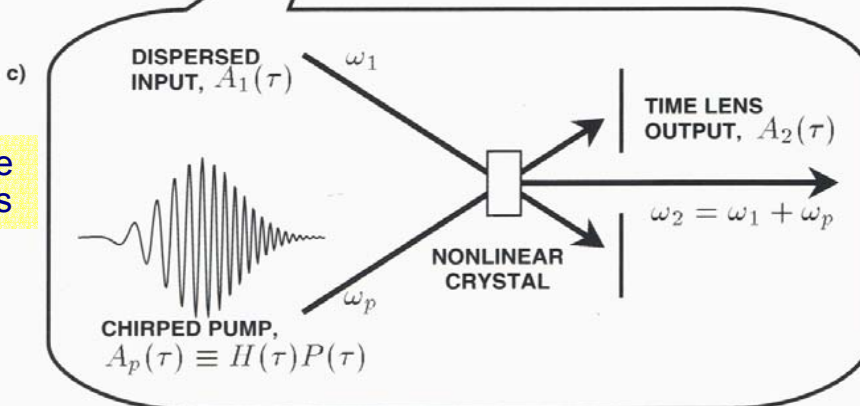
**Imaging Condition**

$$\frac{1}{\phi''_1} + \frac{1}{\phi''_2} = \frac{1}{\phi''_f}$$

**Magnification**

$$M = -\frac{\phi''_2}{\phi''_1}$$

Time Lens



**Group Delay Dispersion (GDD):**

$$\phi'' = \sum_n \xi_n \beta''_n = \left. \frac{d^2 \phi(\omega)}{d\omega^2} \right|_{\omega=\omega_0}$$

**Focal GDD**

$$\phi''_f \equiv \frac{-1}{d\omega/d\tau}$$

**Resolution**

$$\delta\tau_{in} \approx \frac{.44}{\Delta f_{pump}}$$

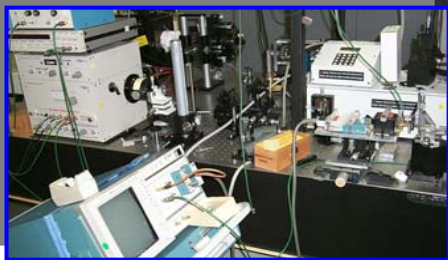


# It Works But Old System Had Many Problems

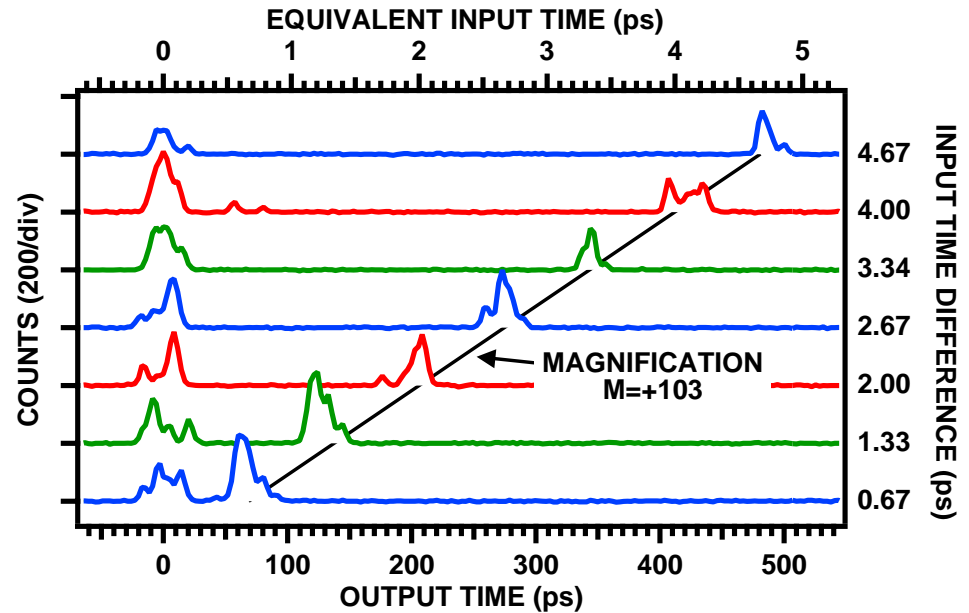
(from DNT LDRD 98-ERD-027)



## Past System Setup



## Streak Camera **Single-Shot** Recording



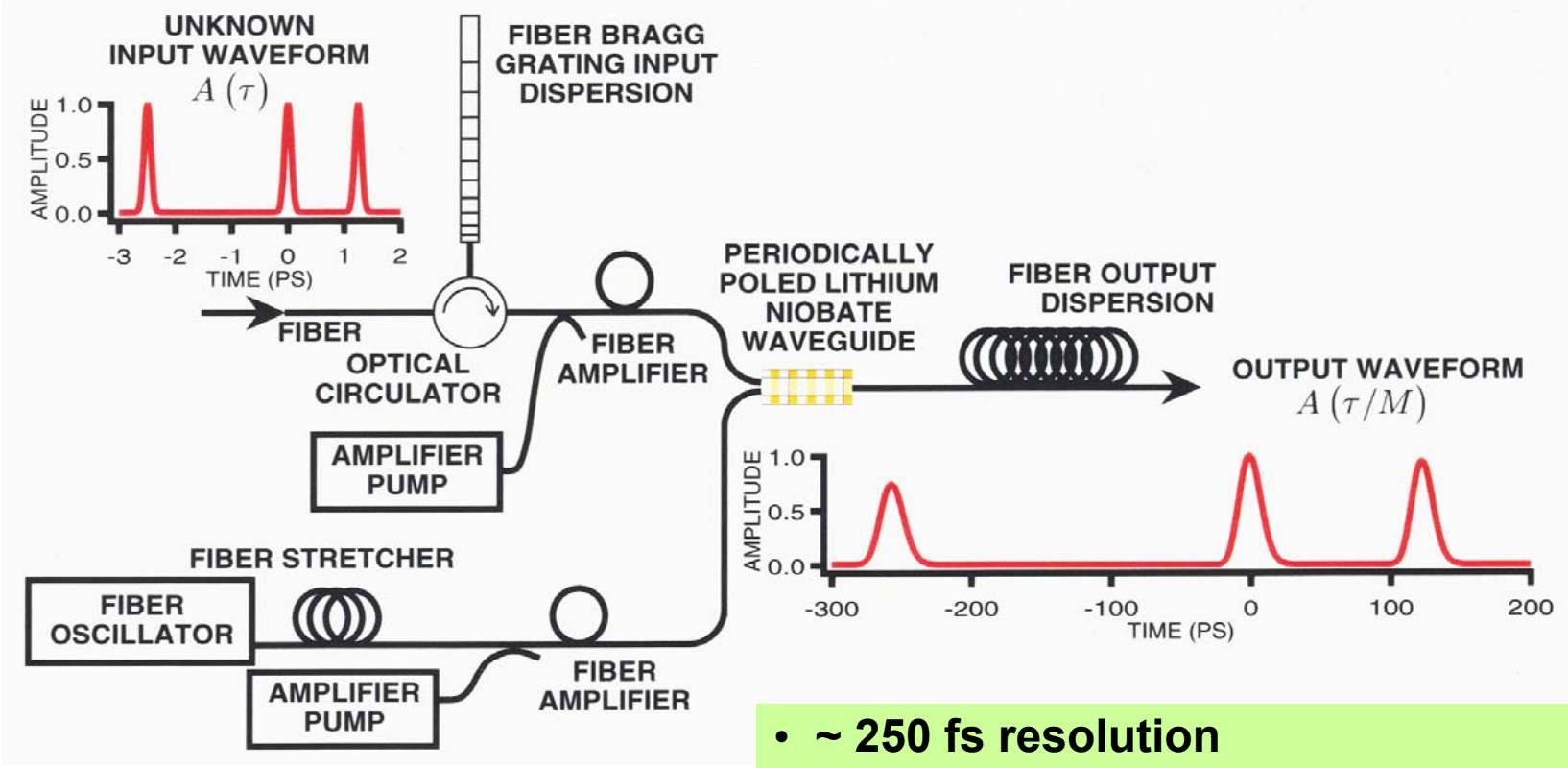
- **VERY LARGE** free space system
  - Filled 5 x 12 ft optical table
- Many mechanical stability problems
- Not practical for imaging

- Two pulse test pattern, changed in 670 fs steps
- 68.8 ps changes at output, demonstrated M=+103 magnification
- **Fundamental problem was low efficiency, producing poor Dynamic Range (DR)**

**A Practical Instrument Requires a Complete Redesign**  
(Introducing new challenges)

(See backup slides for addition past results and publications from LDRD 98-ERD-027)



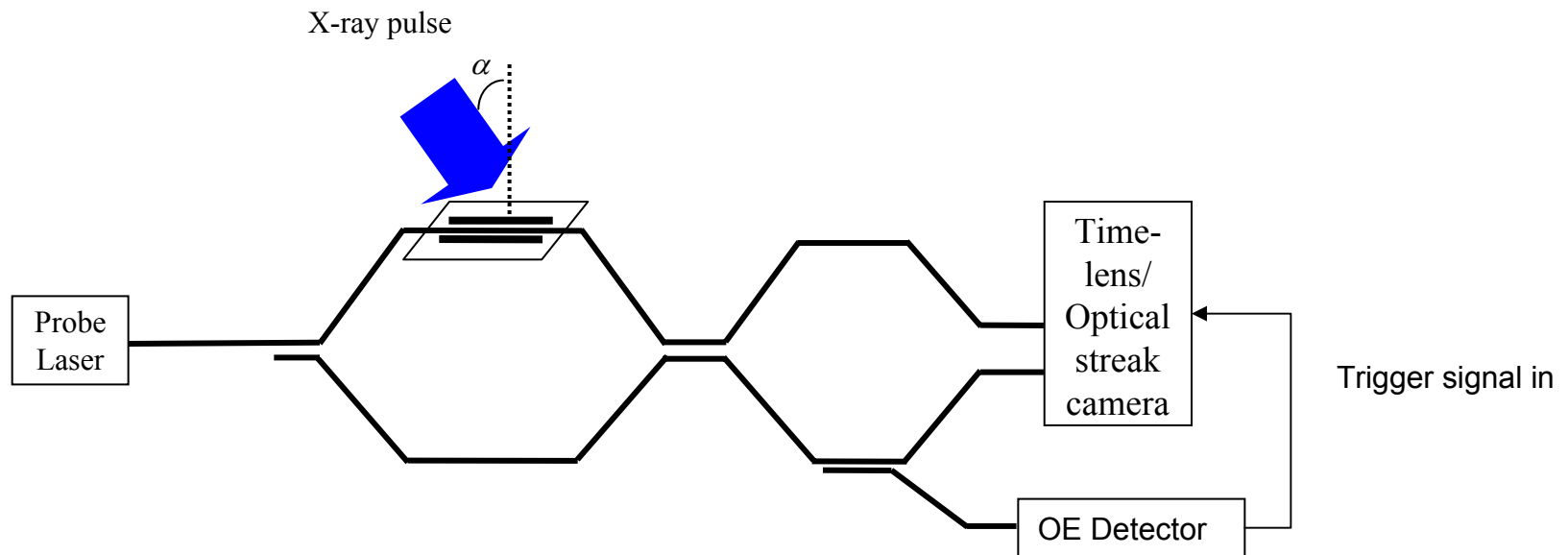


• **New Challenges:**

- Noise due to Amplified Spontaneous Emission
- Aberrations due to higher order dispersion terms and possible self phase modulation
- Polarization Mode Dispersion
- Packaging of nonlinear crystal with fiber input & output

- ~ 250 fs resolution
- Dynamic Range > 100
- Practical record length (100ps – 1 ns)
- Compact and Robust

# RadSensor/Time lens approach to xray pulse measurement



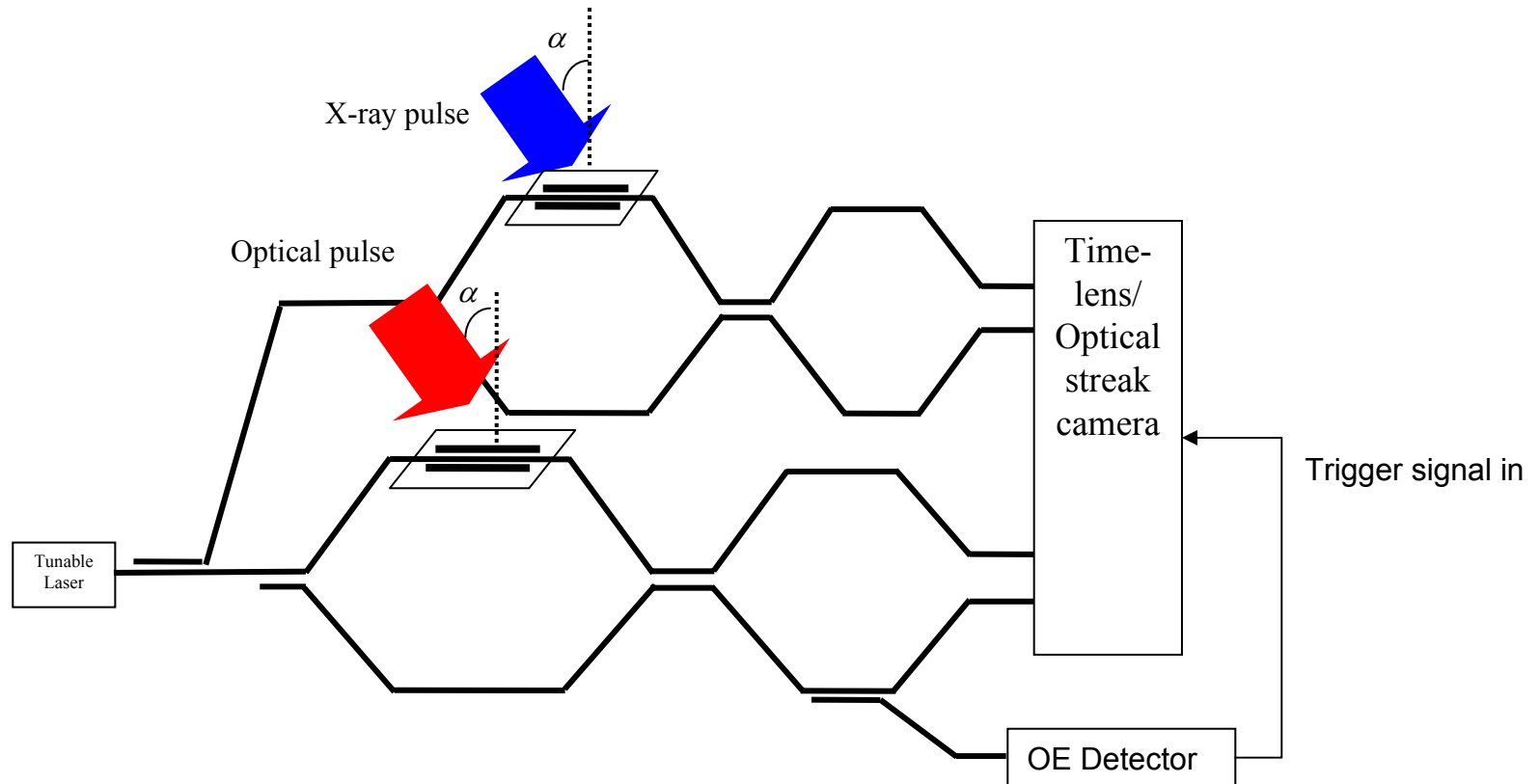
Phase matching condition: 
$$\frac{v_{\text{probe}}}{c} = \sin \alpha = \frac{1}{n} = \frac{1}{3.5} \rightarrow \alpha = 16.6^\circ$$

For slow-recovery material (integrating detector), signal can be differentiated to obtain pulse shape.

Fast recovery material will probably yield better dynamic range

**~100 fs temporal resolutions are possible**

# Potential RadSensor Based Cross-timing Scheme



Phase matching condition: 
$$\frac{v_{\text{probe}}}{c} = \sin \alpha = \frac{1}{n} = \frac{1}{3.5} \rightarrow \alpha = 16.6^\circ$$

**Cross-timing ~ 100 fs is possible using just the rising edge of the RadSensor signal**

# Conclusions and future work



- We have
  - demonstrated that xrays can be produce an optical phase modulation for detection purposes, that should scale to  $< 1$ ps.
  - investigated the xray sensitivity as a function of wavelength separation from the band-edge...  $1.0 \times 10^{-4}$  fringe-fractions/xray photon is best measured
  - Measured the linearity over 2 decades of xray fluence.
  - Developed model in reasonable agreement with measurements
  - Recently demonstrated single-shot results with new cavity geometry
- We plan to:
  - Improve the sensitivity using optimized cavity structures (USP experiments)
    - Goal is single xray photon sensitivity
  - make fast devices and characterize temporal response ( $< \text{ps}$ )
  - Develop imaging versions
  - Develop companion optical recording technologies (Time lens/streaker)

**We believe these approaches are capable of 100 fs temporal resolution and reasonable dynamic range**

# Backup slides

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# Our device design model is aimed at optimizing sensitivity

Model output

$$\frac{\delta P_{Rx}}{P_{Rx}^0} = \frac{1}{R(\phi)} \frac{dR}{d\phi} \delta \hat{\phi}(\lambda)_{rad}$$

The “Contrast Ratio” is maximized for optimum detector sensitivity

Model inputs

$$n = H(E_{gap}, \lambda)$$

The linear index is a function of material composition which is directly related to the energy gap. We use empirically derived polynomial expressions from Amman and Buus

$$\delta \hat{\phi}(\lambda)_{rad} = \frac{2\pi}{\lambda} \frac{C}{\rho_{sat}} \frac{1}{A_{mode}} \frac{E_{rad}}{E_0} G(\lambda)$$
$$\alpha = \alpha_g \exp\left[\frac{hc/\lambda - E_{gap}}{E_{urb}}\right]$$

The resonant nature of the nonlinearity implies close to the bandedge is good, more phase shift

The Urbach absorption tail is higher, closer to the bandedge, thus close to the bandedge is bad

We are exploring the device design parameter space: mirror reflectivities, thickness, wavelength offset to quantify the trade-offs to maximize sensitivity



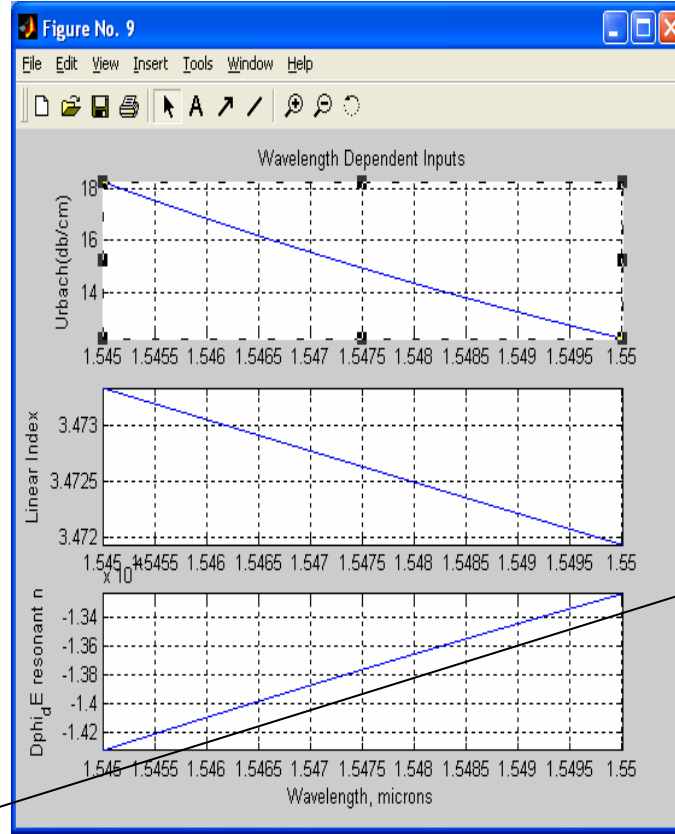
# Design model examples for sensitivity optimization



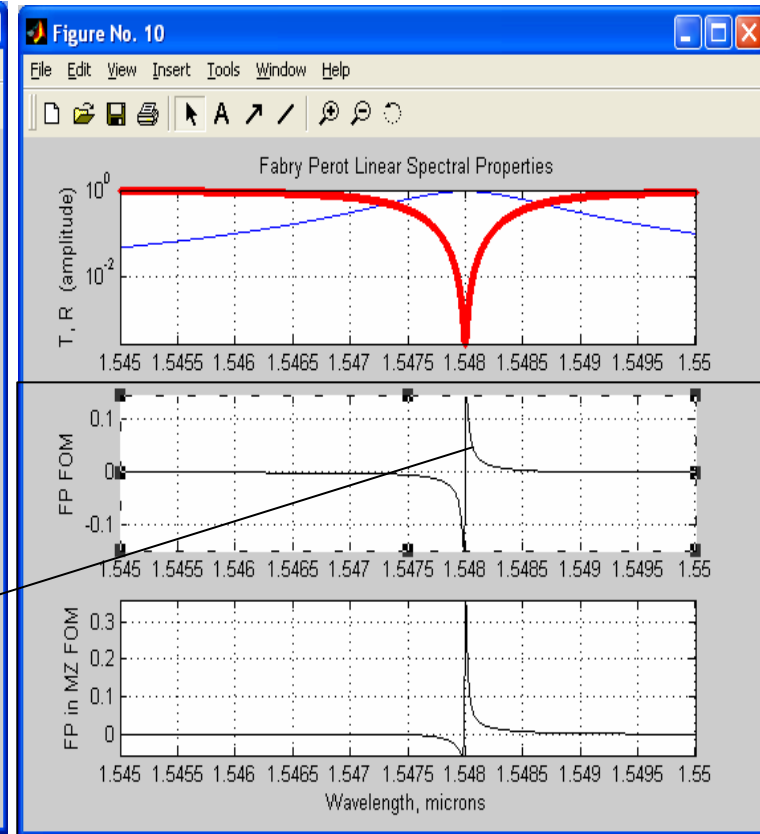
## Model inputs

Parameters		
lam	1.55	um
lam1	1.545	um
lam2	1.55	um
R1	0.945	
phi1	0	radians
R2	0.95	
phi2	0	radians
d	5.16	um
n	3	
alphadB	0	dB/um
T_lump	1	RT trans
Aeff	1	um <sup>2</sup>
n2	1e-017	m <sup>2</sup> /W
alpha2	0	m/W
Pmax	1	W
phi0	0	pi
lambda_edge	1.47	um
Exb	0.003	ev
Lexw	0.004	um
E0	0.0065	ev
alpha_g	1.0e+004	dB/cm

## "Materials Physics"



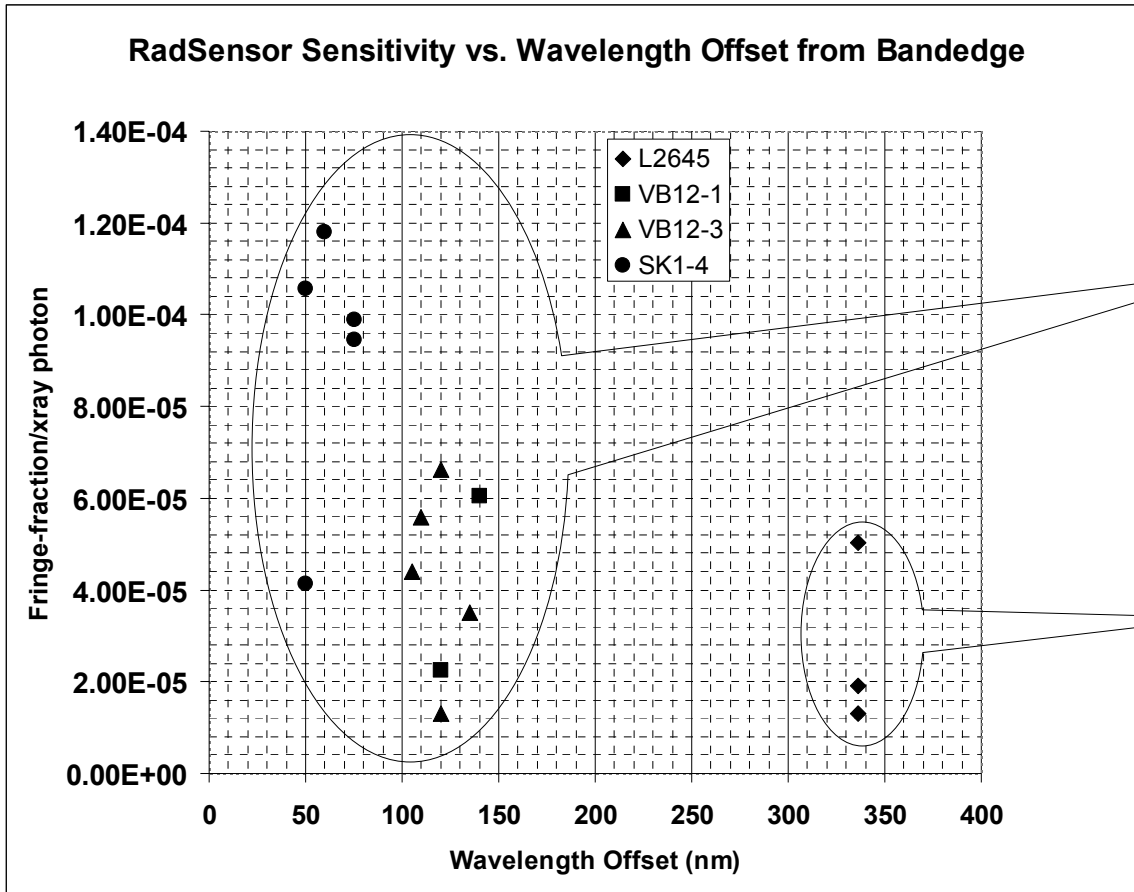
## Cavity physics; predicted responsivity



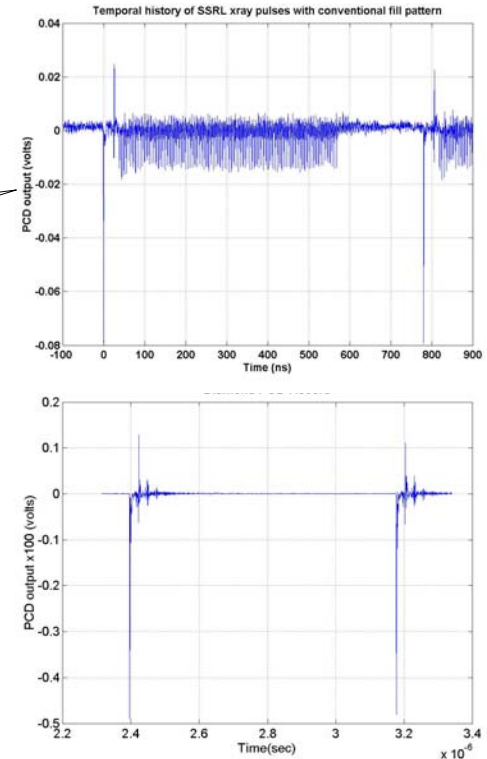
**5% contrast for a single xray photon at 8 keV, this should be detectable**

**The model also outputs cavity results in 3D (vs. wavelength and thickness)...**

# Summary of RadSensor sensitivity data



## Differing Xray excitation dynamics

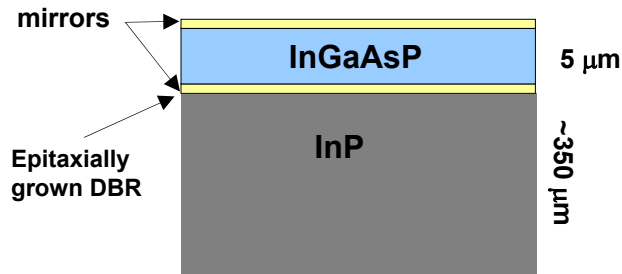


- Scatter in data is primarily due to polarization instability in the interferometer caused by packaging induced birefringence in the RadSensor– higher values probably more accurate
- We should see more resonant enhancement. Trap-filling effects may be causing the as measured “normal fill-pattern” fringe-fractions to be preferentially saturated (x10 xray photons/SPEAR period)

# We considered several cavity design approaches to mitigate risk



## Epi-DBR-thin cavity



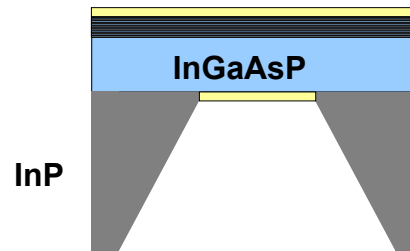
### Pros:

- Very VCSEL-like know how to produce high-quality laser cavities
- Material in the cavity is only InGaAsP

### Cons:

- MOCVD shutdown forced reliance on vendors– no takers
- Difficult to match resonances to probe range (10 μm OK)

## Etched-back, thin cavity



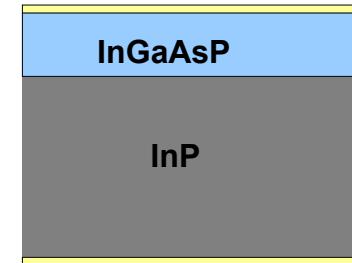
### Pros:

- We have control over final cavity thickness and mirror reflectivities

### Cons:

- Cavity trimming (not quite working)
- Membrane is fragile and stressed

## Thick cavity



### Pros:

- Robust (thick)
- Relatively easy to make

### Cons:

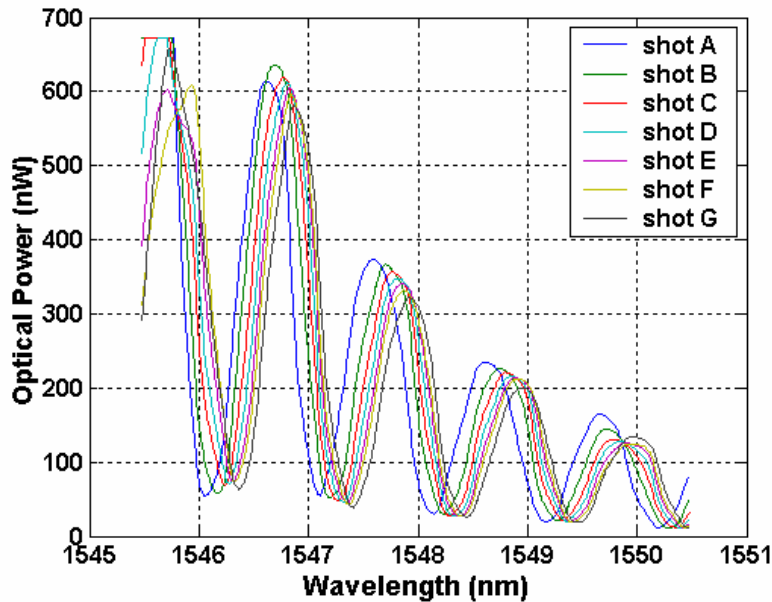
- Cavity includes InP and InGaAsP
- Difficult to get good effective finesse
- Impossible to get fast response

To meet our USP fielding schedule we had to go with the thick cavity

# First generation RadSensor Cavities Presented Challenges



Cavity resonances showing thermal drift

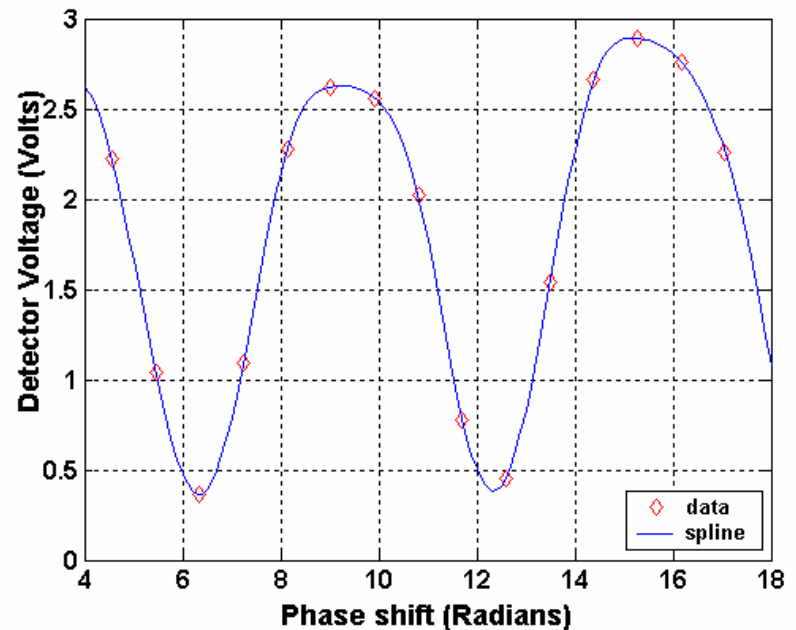


- Finesse is very poor ( $\sim 1$ )
- Measured using ASE from EDFA
- Drift went away when optical input power was lowered to  $\sim 0$  dBm
- We suspected this would be a problem—locking circuit not ready for this fielding... will be next time

Temporary cavity drift solution:

Map detector voltage vs. phase by varying wavelength and modulating light on/off – Optical receiver is ac-coupled

Detector Voltage (Volts) vs. Phase shift (Radians)

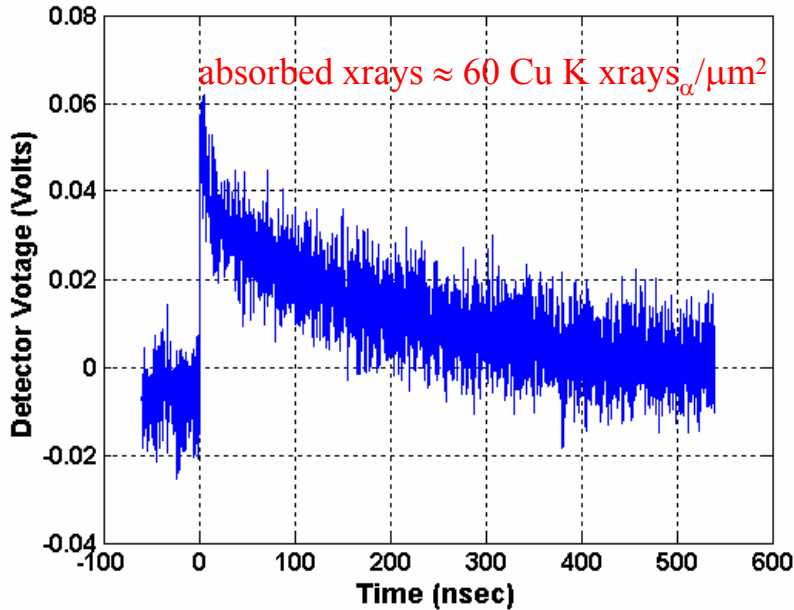


# Typical Results Compared to Prediction from SSRL derived empirical constants

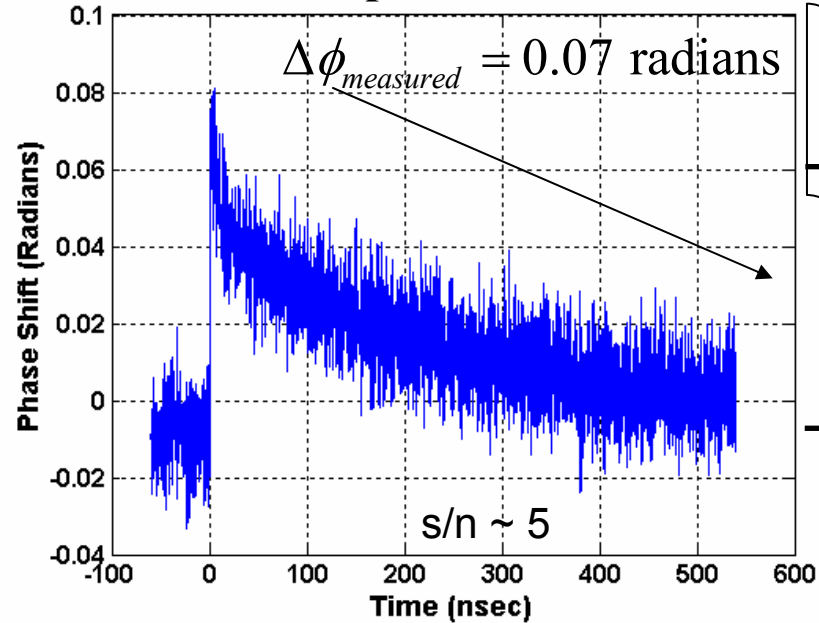


These data used Be and Cu filters to narrow xray spectrum

## Detector Voltage



## Optical Phase



$$\Delta\phi_{\text{predicted}} = \psi (1 - \exp(-\mu d)) \frac{2\pi}{\lambda} \frac{C}{\rho_{\text{sat}}} \frac{E_{\text{rad}}}{E_0} G(\lambda)$$

$$= (60 \text{ xrays}/\mu\text{m}^2) \left( \frac{2\pi}{1.55 \mu\text{m}} \right) (1.2 \times 10^{-6} \mu\text{m}^3) \left( \frac{8.05 \text{ keV}}{3.15 \text{ eV}} \right) (0.04) = 0.030 \text{ radians}$$

Reasonably good comparison between SSRL experiments and USP... please note— these are very preliminary results, analysis still underway